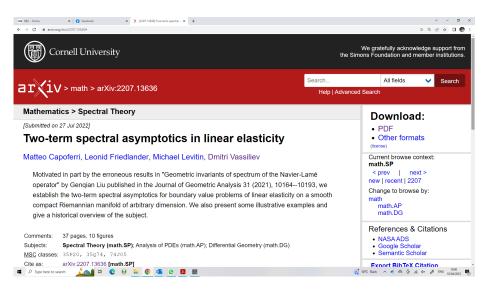
# Two-term spectral asymptotics in linear elasticity

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9 May 2023

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May 31	<u>Maciej Zworski</u> <u>Spectral theory of internal waves in fluids</u> <u>Video</u>	
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La forza del destino

### Contents [hide] Article Talk Read Edit View history Tools ~ (Top) From Wikipedia, the free encyclopedia Performance history La forza del destino (Italian pronunciation: Ila 'fortsa del de'sti:no]: The Power of Fate.<sup>[1]</sup> often La forza del destino translated The Force of Destiny) is an Italian opera by Giuseppe Verdi. The libretto was written Opera by Giuseppe Verdi Recent critical editions by Francesco Maria Piave based on a Spanish drama. Don Álvaro o la fuerza del sino (1835). Roles by Ångel de Saavedra, 3rd Duke of Rivas, with a scene adapted from Friedrich Schiller's Wallensteins Lager (Wallenstein's Camp). It was first performed in the Bolshoi Kamenny Theatre Instrumentation of Saint Petersburg, Russia, on 29 November 1862 O.S. (N.S. 10 November). Synopsis La forza del destino is frequently performed, and there have been a number of complete Overture recordings. In addition, the overture (to the revised version of the opera) is part of the standard Act 1 repertoire for orchestras, often played as the opening piece at concerts. Act 2 Performance history [edit] Act 3 Act 4 Revisions [edit] Superstition After its premiere in Russia, La forza underwent some revisions and made its debut abroad with Other media performances in Rome in 1863 under the title Don Alvaro. Performances followed in Madrid Recordings (with the Duke of Rivas, the play's author, in attendance) and the opera subsequently travelled to New York, Vienna (1865), Buenos Aires (1866), and London (1867). [citation needed] References c. 1870 poster by Charles Lecoco Librettist Francesco Maria Piave Following these productions, Verdi made further, more extensive revisions to the opera with External links

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# Playing field

Let (M, g) be a closed Riemannian *d*-manifold.

Consider a diffeomorphism  $\varphi: M \to M$ . This is the unknown quantity of elasticity theory.

Second Riemannian metric  $h := \varphi^* g$ , the pullback of g.

A pair of metrics, g and h, allows us to write down an action (variational functional).

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# Strain tensor

Linear algebra: a pair of non-degenerate symmetric bilinear forms  $g, h: V \times V \to \mathbb{R}$  in a real finite-dimensional vector space V defines an invertible linear operator  $L: V \to V$  via the formula

$$h(u,v) = g(Lu,v), \quad \forall u,v \in V.$$

Convenient to subtract the identity operator,

$$S := L - \mathrm{Id}$$

Definition of strain tensor:

$$S^{\alpha}{}_{\beta}(x) := [g^{\alpha\gamma}(x)] [h_{\gamma\beta}(x)] - \delta^{\alpha}{}_{\beta}.$$

Describes, pointwise, linear map in the fibres of the tangent bundle

$$v^{\alpha} \mapsto S^{\alpha}{}_{\beta} v^{\beta}.$$

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### Scalar invariants of the strain tensor

Obvious choice: 
$$tr(S^k)$$
,  $k = 1, 2, ..., d$ .

More convenient choice:

$$e_1(\varphi) := \operatorname{tr} S = \lambda_1 + \lambda_2 + \ldots + \lambda_d ,$$

$$e_2(\varphi) := \frac{1}{2} \left[ (\operatorname{tr} S)^2 - \operatorname{tr}(S^2) \right] = \lambda_1 \lambda_2 + \lambda_1 \lambda_3 + \ldots + \lambda_{d-1} \lambda_d ,$$

$$\vdots$$

$$e_d(\varphi) := \det S = \lambda_1 \lambda_2 \ldots \lambda_d .$$

Elementary symmetric polynomials. The  $\lambda_i$  are eigenvalues of S.

# Action (potential energy of elastic deformation)

$$\int_{M} \mathcal{L}(e_1(\varphi), e_2(\varphi), \dots, e_d(\varphi)) \sqrt{\det g} \, dx \, ,$$

where  $\mathcal{L}$  is some prescribed smooth real-valued function of d real variables and  $dx := dx^1 dx^2 \dots dx^d$ .

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## Describing diffeomorphism in terms of a vector field

**First approach** Use integral curves of a vector field. Impossible: J.Milnor 1983.

### Second approach Use geodesics.

Connect a point  $P \in M$  with the point  $\varphi(P) \in M$  by a geodesic  $\gamma : [0,1] \to M$ , so that  $\gamma(0) = P$  and  $\gamma(1) = \varphi(P)$ . Parameterise the geodesic in such a way that  $\gamma(t)$  is a solution of the equation

$$\ddot{\gamma}^{\lambda} + \left\{ egin{smallmatrix} \lambda \ \mu
u \end{array} 
ight\} \dot{\gamma}^{\mu} \dot{\gamma}^{
u} = \mathbf{0},$$

where the dot stands for differentiation in t.

Define the vector field of displacements as

$$u: M \ni P \mapsto \dot{\gamma}(0) \in TM.$$

# Linear elasticity

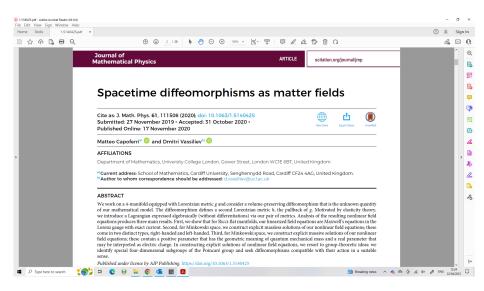
- Linearise the strain tensor with respect to the vector field of displacements u.
- Choose action quadratic in u.

Action reads

$$\frac{1}{2}\int_{\mathcal{M}}\left(\lambda\left(\nabla_{\alpha}u^{\alpha}\right)^{2}+\mu\left(\nabla_{\alpha}u_{\beta}+\nabla_{\beta}u_{\alpha}\right)\nabla^{\alpha}u^{\beta}\right)\sqrt{\det g} \, dx\,,$$

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where  $\lambda$  and  $\mu$  are Lamé coefficients.



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Spectral problem for linear elasticity:

$$-\mu\left(\nabla_{\beta}\nabla^{\beta}u^{\alpha}+\operatorname{Ric}^{\alpha}{}_{\beta}u^{\beta}\right)-(\lambda+\mu)\nabla^{\alpha}\nabla_{\beta}u^{\beta}=\Lambda u^{\alpha}.$$

Possible boundary conditions.

- Dirichlet.
- Free boundary. This is not the Neumann boundary condition.

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### Historical overview 1

**1885** Lord Rayleigh discovers *Rayleigh wave*. Wave runs along free boundary and exponentially decays towards interior. Let

$$R_{\alpha}(w) := w^{3} - 8w^{2} + 8(3 - 2\alpha)w + 16(\alpha - 1),$$

where

$$\alpha := \frac{\mu}{\lambda + 2\mu}.$$

The cubic equation  $R_{\alpha}(w) = 0$  has three roots  $w_j$ , j = 1, 2, 3, over  $\mathbb{C}$ , where  $w_1$  is the distinguished real root in the interval (0, 1). Put

$$\gamma_{R} := \sqrt{w_1} \, .$$

The subscript R in  $\gamma_R$  stands for "Rayleigh". The quantity

$$c_{R} := \sqrt{\mu} \gamma_{R}$$

has the physical meaning of velocity of Rayleigh's surface wave.

### Historical overview 2

**1912** Peter Debye writes down one-term asymptotic formula for the eigenvalue counting function

$$\mathcal{N}(\Lambda) = a \operatorname{Vol}_d(M) \Lambda^{d/2} + o\left(\Lambda^{d/2}\right) \quad \text{as} \quad \Lambda \to +\infty,$$

where

$$a = rac{1}{(4\pi)^{d/2} \Gamma\left(1+rac{d}{2}
ight)} \left(rac{d-1}{\mu^{d/2}} + rac{1}{(\lambda+2\mu)^{d/2}}
ight) \, .$$

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1915 Hermann Weyl provides rigorous proof.

### Historical overview 3

Search for two-term asymptotic formula

$$\mathcal{N}(\Lambda) = a \operatorname{Vol}_d(M) \Lambda^{d/2} + b \operatorname{Vol}_{d-1}(\partial M) \Lambda^{(d-1)/2} + o\left(\Lambda^{(d-1)/2}
ight)$$
 as  $\Lambda o +\infty$ 

Second Weyl coefficient *b* should depend on boundary conditions.

**1950** E. W. Montroll publishes incorrect formulae for second Weyl coefficient. Same incorrect formulae as Genquian Liu in 2021.

**1960** Lars Onsager and coauthors publish correct formulae for second Weyl coefficient for d = 3.

**1997** Safarov and Vassiliev book (only results, without details).

- Onsager's results for d = 3 checked and confirmed.
- Formulae for second Weyl coefficient for d = 2 written down.

# Algorithm for the calculation of second Weyl coefficient

- Fix point x' ∈ ∂M, freeze coefficients in operator and boundary conditions and perform Fourier transform ∂x' → iξ' along ∂M. Gives spectral problem for a system of ODEs with constant coefficients on semi-axis [0, +∞). This 1-dimensional spectral problem depends on (x', ξ') ∈ T\*∂M as a parameter.
- Need to calculate the spectral shift function shift(x', ξ', Λ) (regularised trace of spectral projection).
- Use ideas from scattering theory:

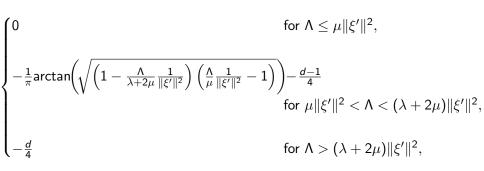
$${\sf shift}(x',\xi',{\sf \Lambda}):=rac{arphi(x',\xi',{\sf \Lambda})}{2\pi}+{\sf N}(x',\xi',{\sf \Lambda})\,,$$

where  $\varphi(x', \xi', \Lambda)$  is scattering phase (phase shift) and  $N(x', \xi', \Lambda)$  is the eigenvalue counting function of the 1-dimensional spectral problem.

$$b = \frac{1}{(2\pi)^{d-1}} \int_{T^* \partial M} \operatorname{shift}(x', \xi', 1) \, dx' \, d\xi' \, .$$

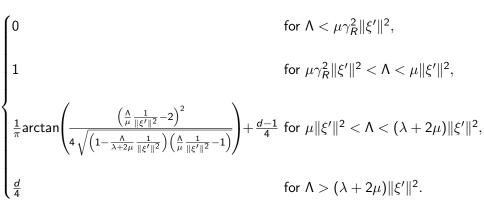
Spectral shift function for Dirichlet boundary conditions

 $\operatorname{shift}_{\operatorname{Dir}}(\xi', \Lambda) =$ 



## Spectral shift function for free boundary conditions

 $\operatorname{shift}_{\operatorname{free}}(\xi', \Lambda) =$ 

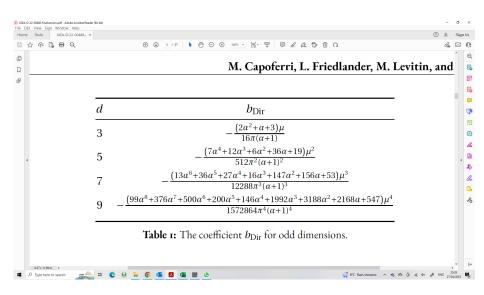


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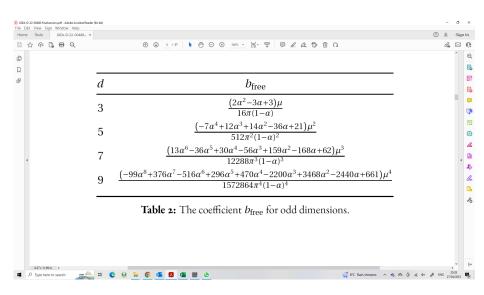
Main result (for  $\alpha$  and  $\gamma_R$  see one of previous slides)

$$b_{\text{Dir}} = -\frac{\mu^{\frac{1-d}{2}}}{2^{d+1}\pi^{\frac{d-1}{2}}\Gamma\left(\frac{d+1}{2}\right)} \left(\frac{4(d-1)}{\pi} \int_{\sqrt{\alpha}}^{1} \tau^{d-2} \arctan\left(\sqrt{(1-\alpha\tau^{-2})(\tau^{-2}-1)}\right) d\tau + \alpha^{\frac{d-1}{2}} + d - 1\right),$$

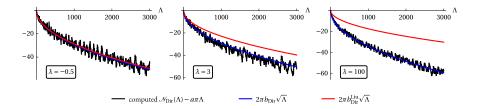
$$b_{\text{free}} = \frac{\mu^{\frac{1-d}{2}}}{2^{d+1}\pi^{\frac{d-1}{2}}\Gamma\left(\frac{d+1}{2}\right)} \left(\frac{4(d-1)}{\pi} \int_{\sqrt{\alpha}}^{1} \tau^{d-2} \arctan\left(\frac{(\tau^{-2}-2)^{2}}{4\sqrt{(1-\alpha\tau^{-2})(\tau^{-2}-1)}}\right) d\tau + \alpha^{\frac{d-1}{2}} + d - 5 + 4\gamma_{R}^{1-d}\right).$$



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