

**Robust uncertainty quantification of tsunami
ionospheric holes for the 2011 Tohoku-Oki
earthquake using satellite data.**

Ryuichi KANAI

University College London
Alan Turing Institute

[Collaborated researchers]

Prof. Alan Smith

University College London

Prof. Serge Guillas

University College London

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University of Shizuoka

Prof. Toshiyasu Nagao

Tokai University

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Space-based system using GPS satellites could warn of incoming tsunamis

28 April 2022

A new method for detecting tsunamis using existing GPS satellites orbiting Earth could serve as an effective warning system for countries worldwide, according to a new study by an international team led by UCL researchers.



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A new method to detect tsunamis using GPS satellites orbiting Earth could be an effective warning system for countries worldwide, according to a new study by an international team led by PhD student Ryuichi Kanai & Prof [@GuillasSerge](#) [@stats_UCL](#) [@uclmaps](#)
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SPACE-BASED SYSTEM USING GPS SATELLITES COULD WARN OF INCOMING TSUNAMIS

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RESEARCH BRIEFS

28 April 2022

Source(s): University College London



**TC** TechCrunch

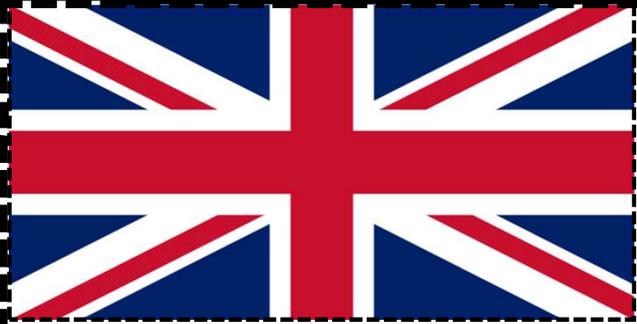
GPS signals could detect tsunamis better and faster than seismic sensors

**Stefanie Waldek**

May 5, 2022 · 3 min read



GPS networks are already a crucial part of everyday life around the world, but an international team of scientists has found a new, potentially life-saving use for them: tsunami warnings.

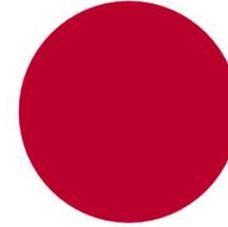


Space Science

UCL's Mullard Space
Science Laboratory

Statistics

UCL's Statistical Science
department



Seismology

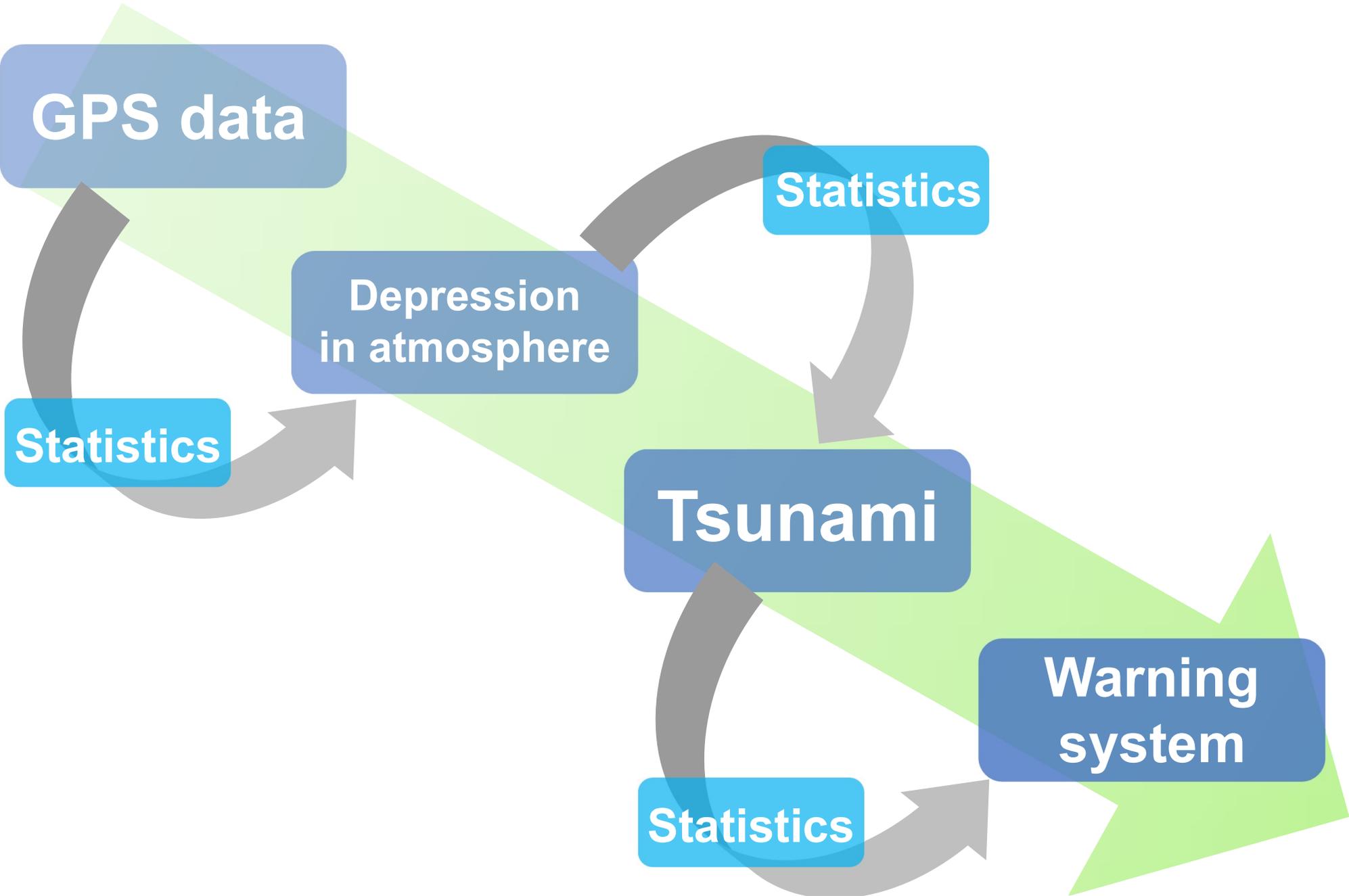
Tokai University

Electrodynamics

University of Shizuoka

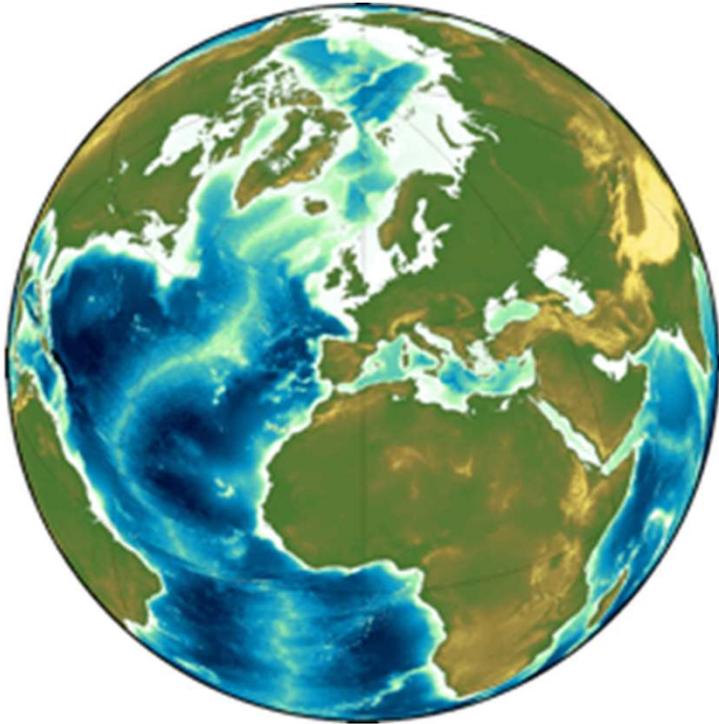
Space-based analysis

Rough sketch



1

Background



2

Application of statistics

3

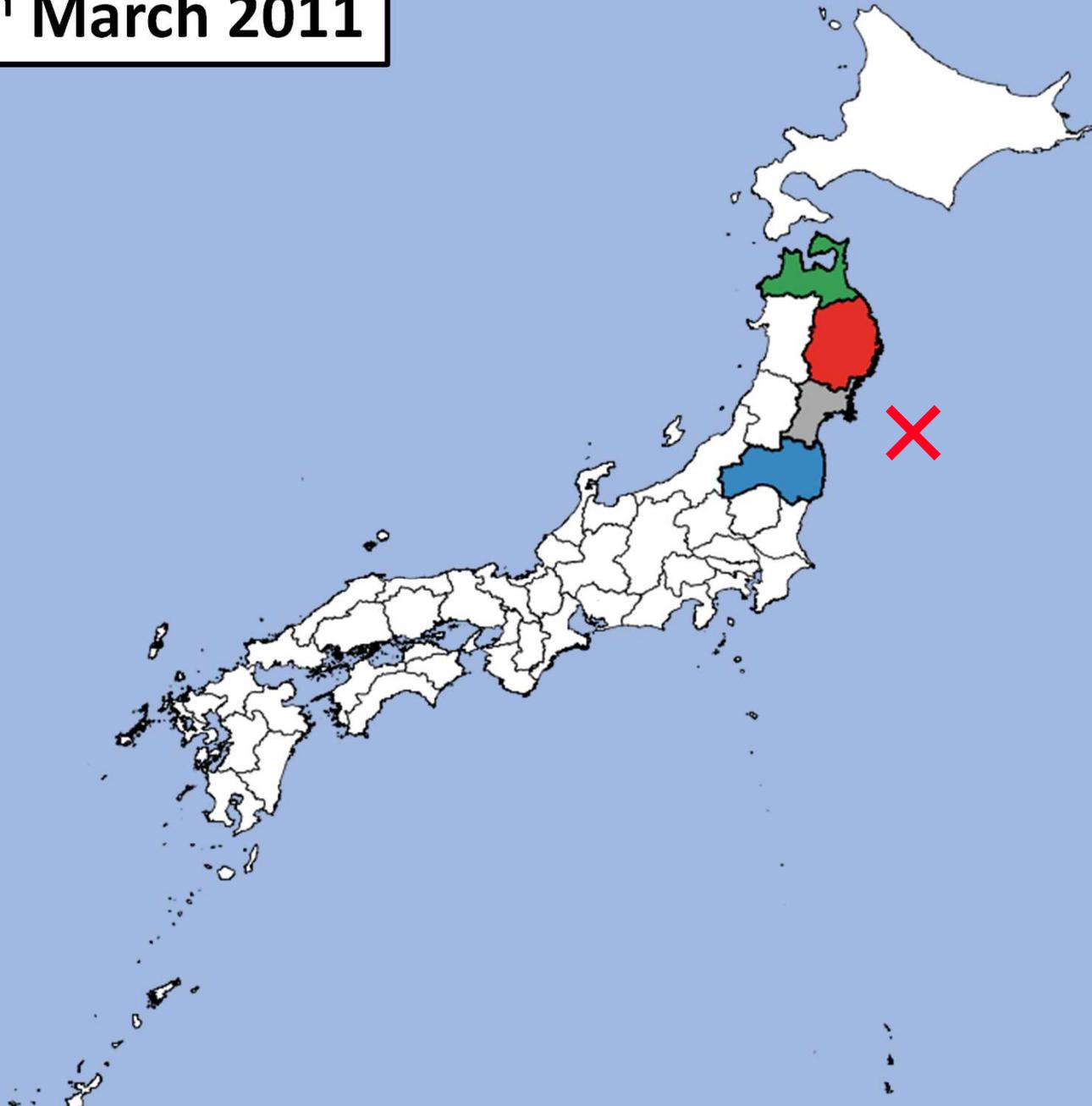
Results

4

Future step

Background

11th March 2011



[Time]
05:46:24 (UTC)

[Location]
38.297°N 142.373°E

[Depth]
29.0 km

[Magnitude]
9.1

Background

11th March 2011

Aomori

Iwate

Miyagi

Fukushima

Tsunami warning

	3 minutes after	28 minutes after	44 minutes after
Aomori	1 meter	3 meters	8 meters
Iwate	3 meters	6 meters	More than <u>10 meters</u>
Miyagi	6 meters	More than <u>10 meters</u>	More than <u>10 meters</u>
Fukushima	3 meters	6 meters	More than <u>10 meters</u>

How to overcome?

Traditional Seismology



Magnitude cannot be estimated.

**Distribute buoys
densely in the ocean**



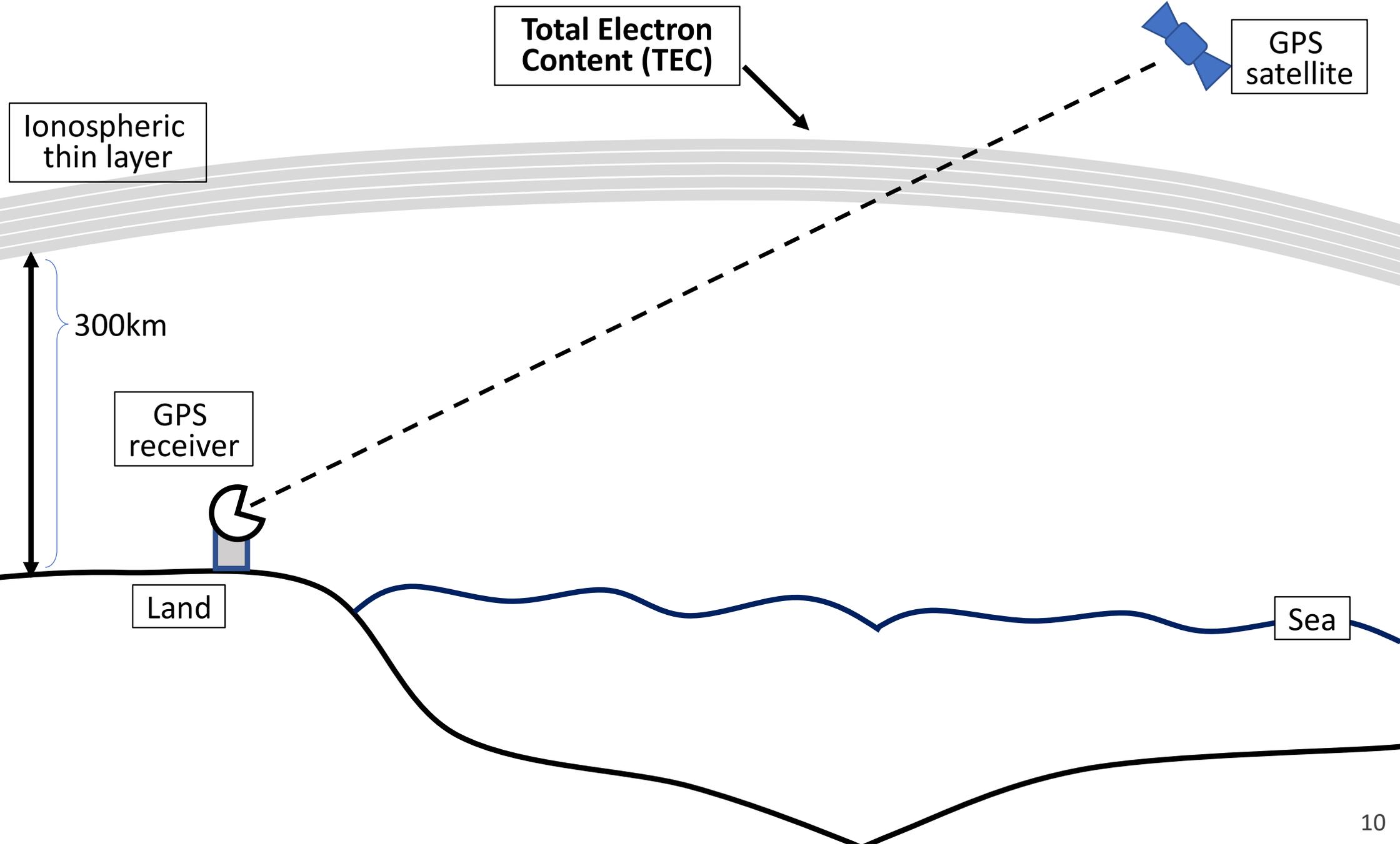
**Impossible due to budget
constraints. No ship can sail.**

**Distribute instruments
on the seafloor**

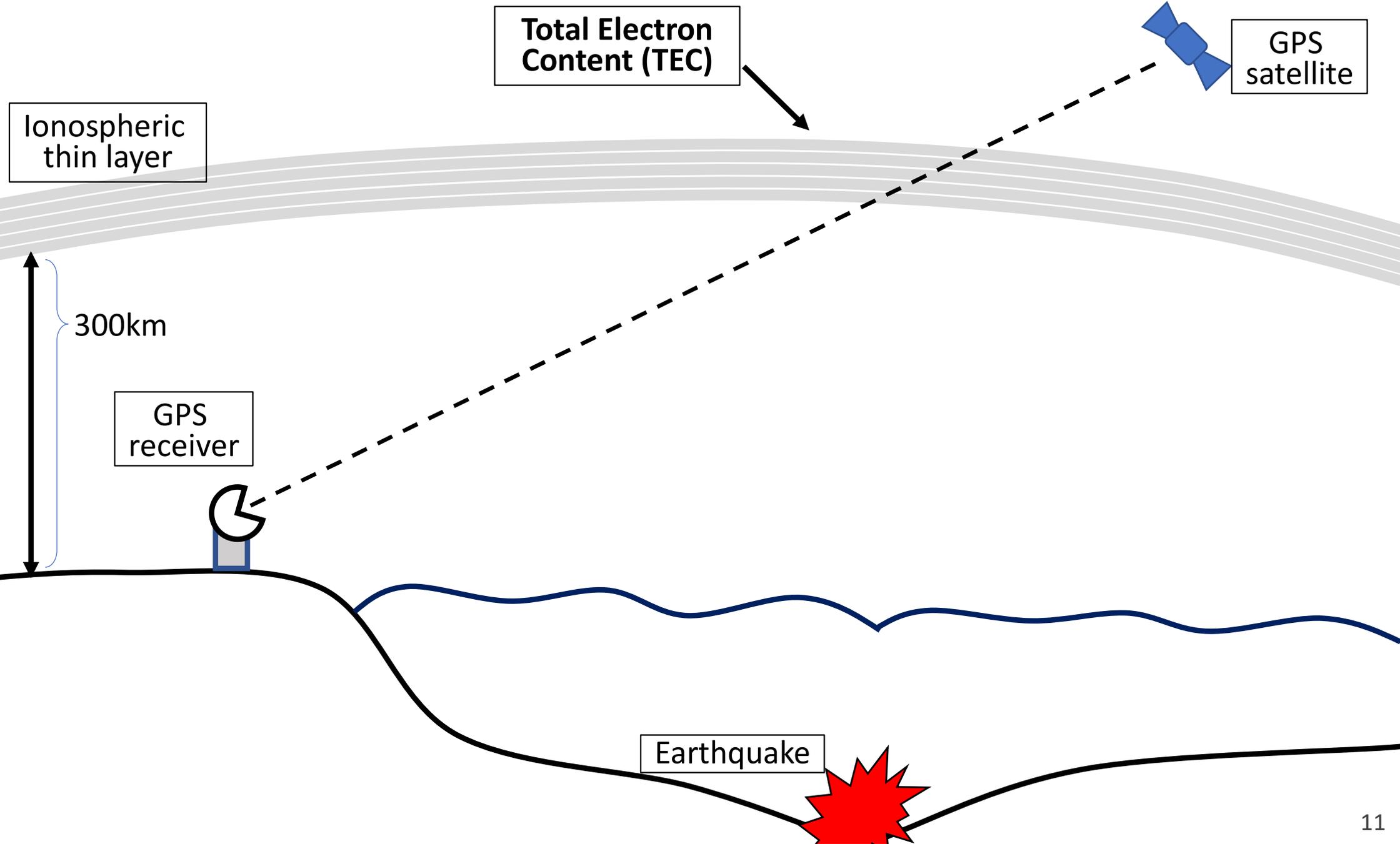


**Impossible to know where
earthquakes occur in advance.**

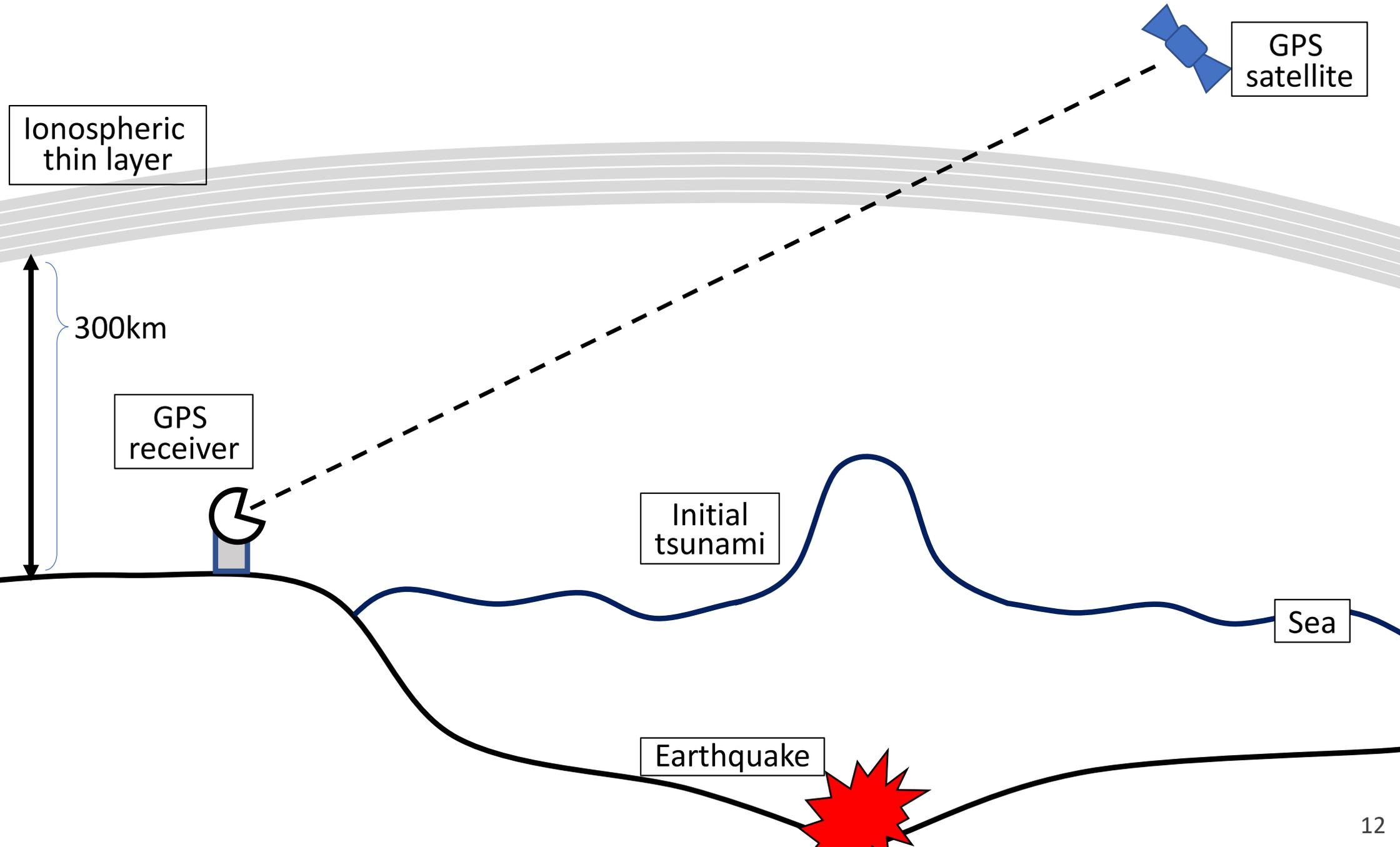
A possible solution



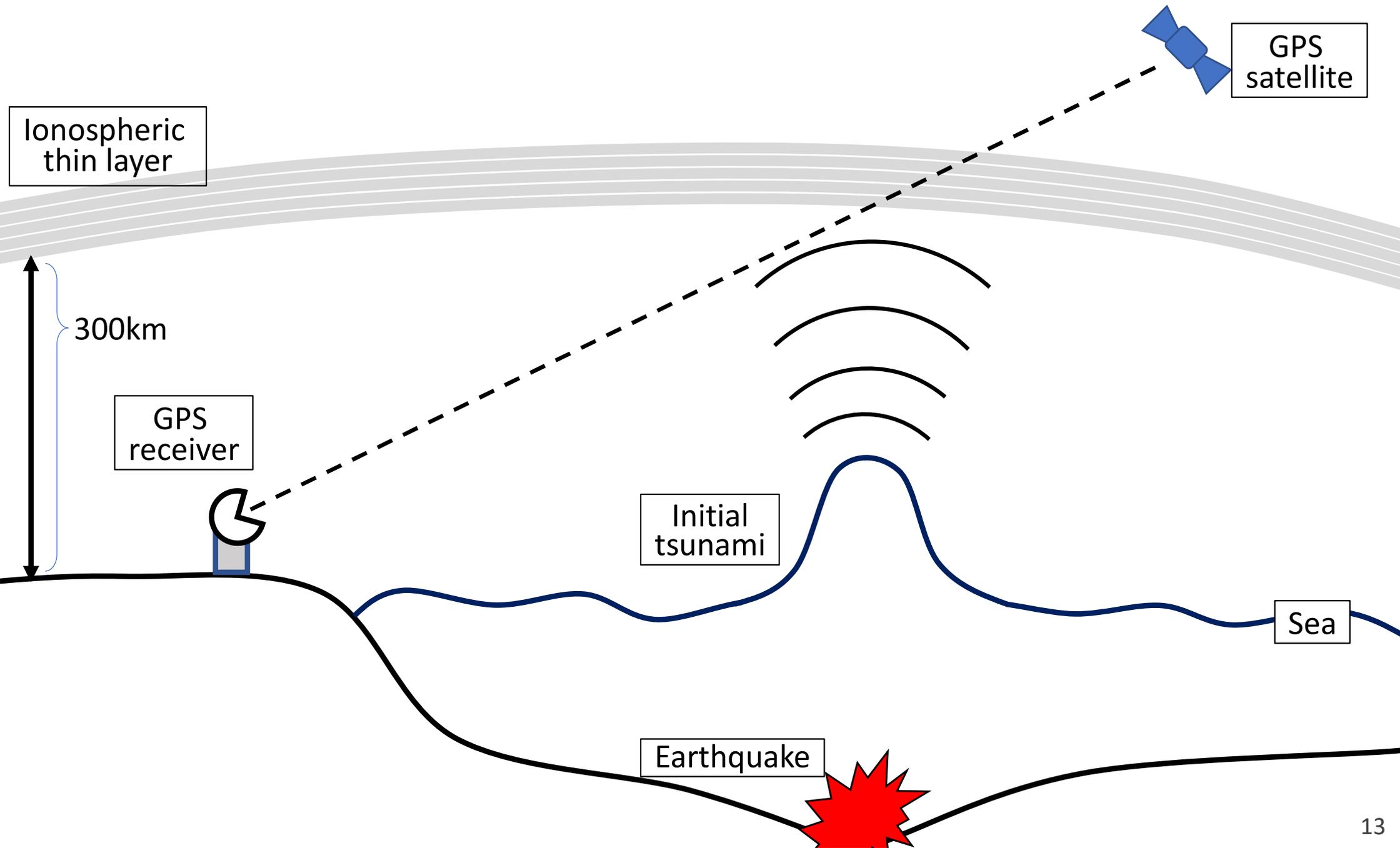
A possible solution



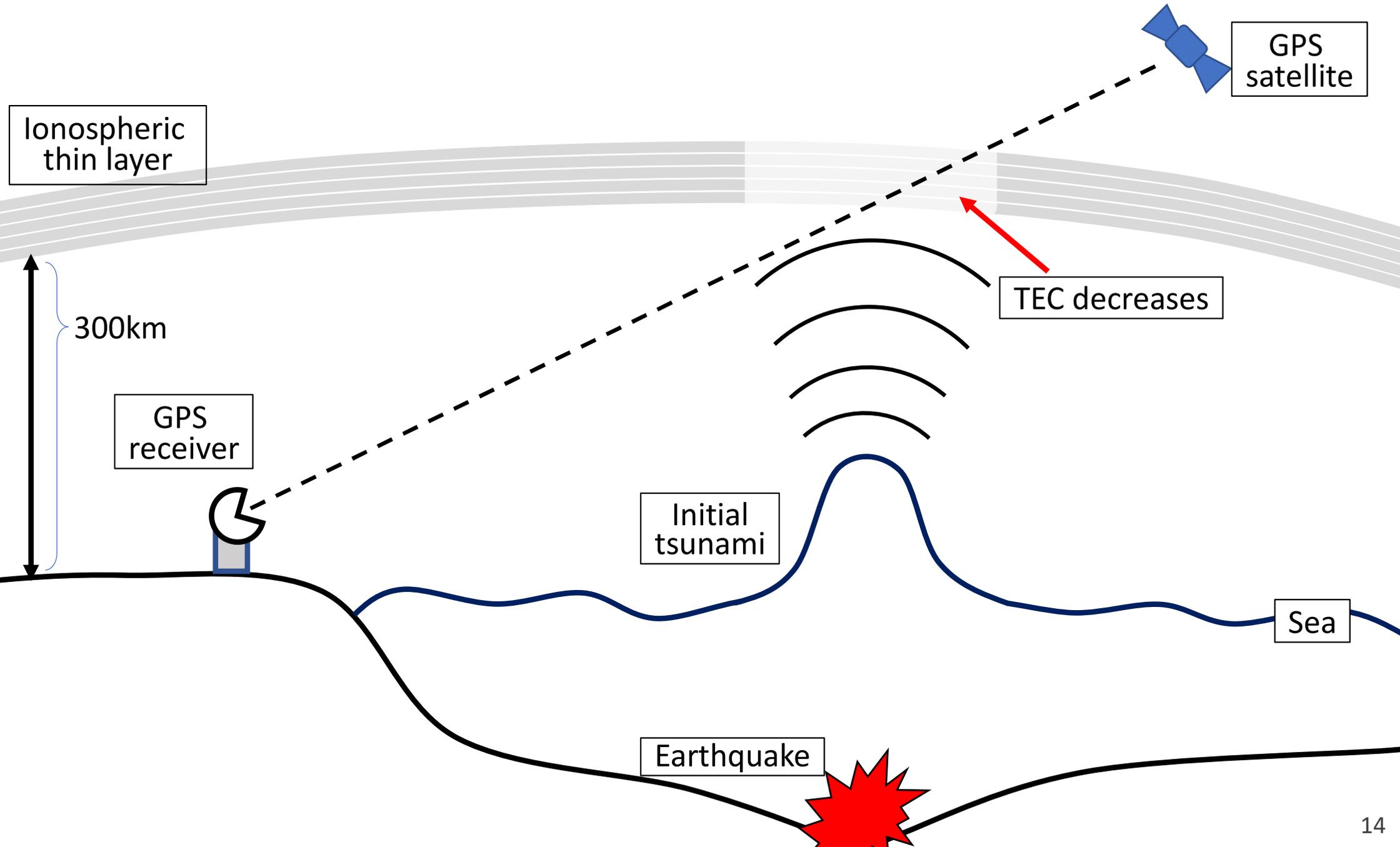
A possible solution



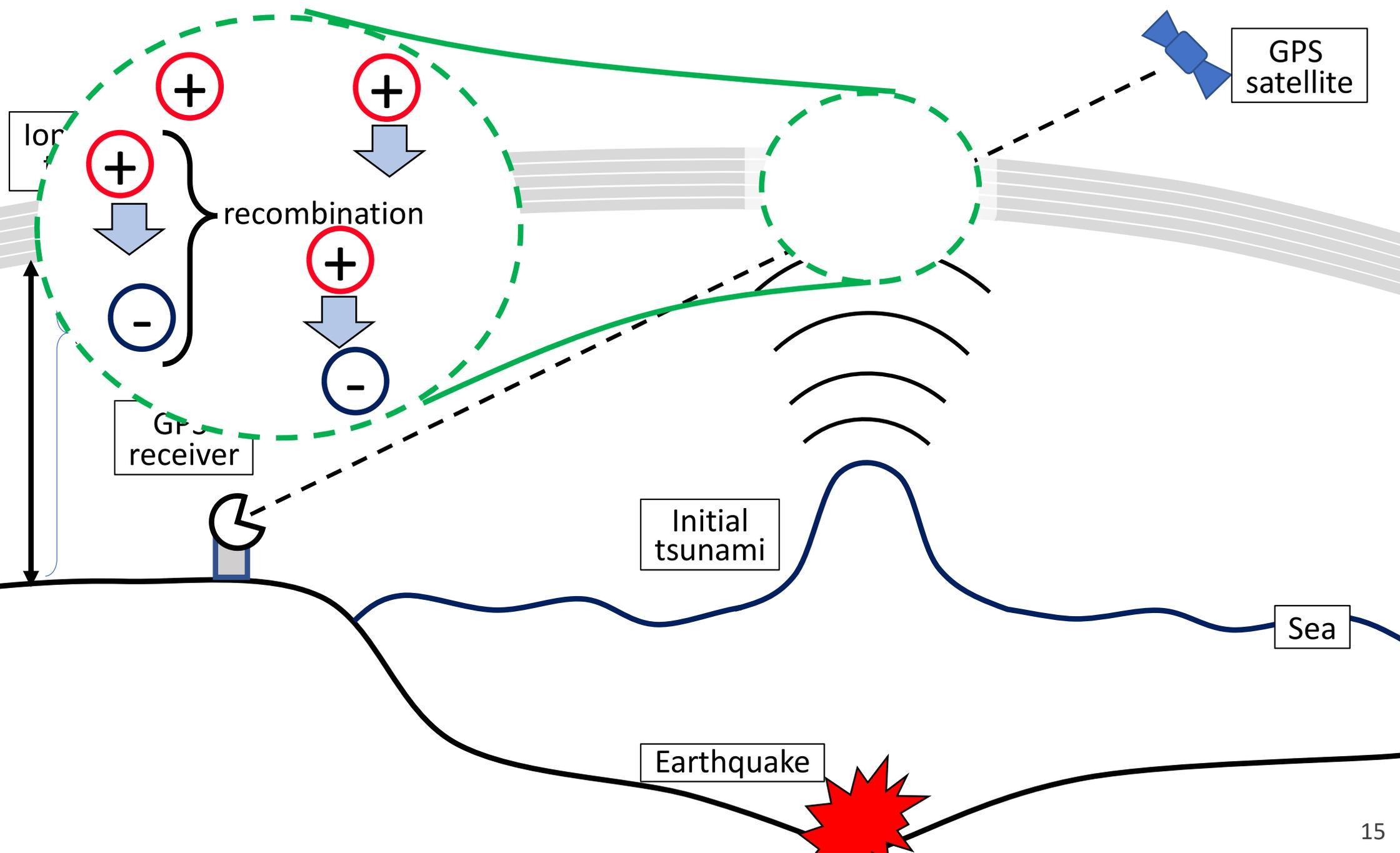
A possible solution



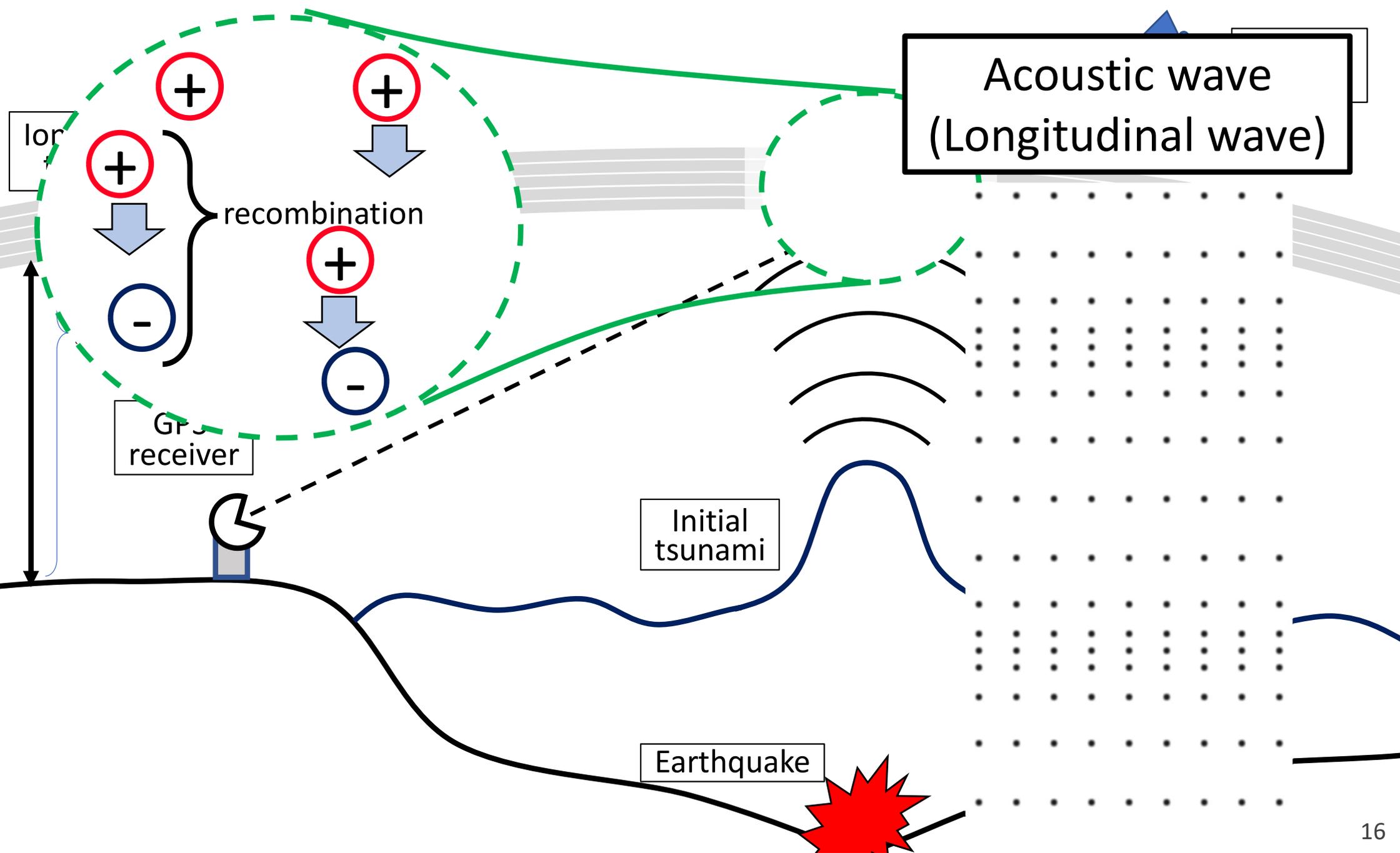
A possible solution



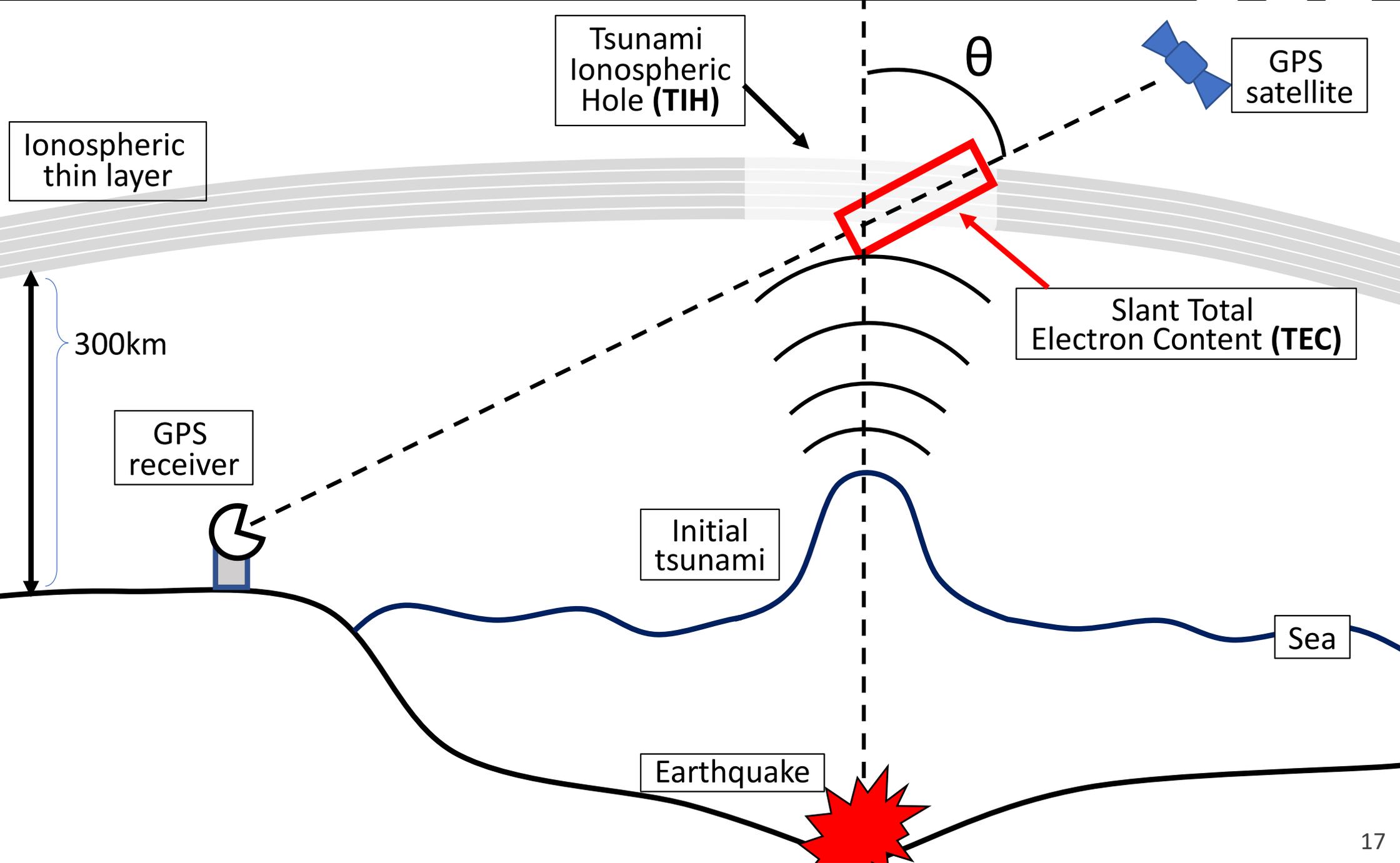
A possible solution



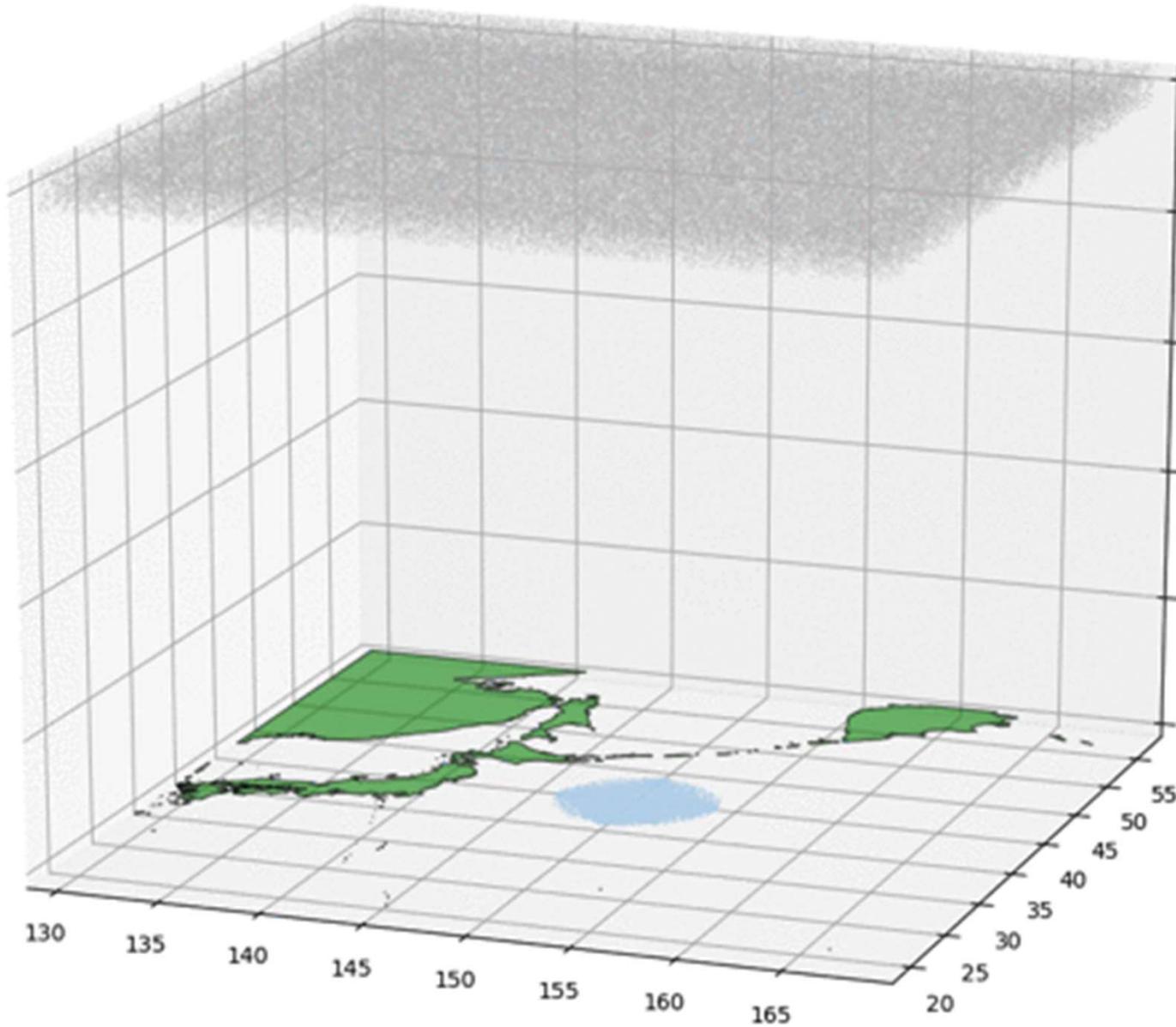
A possible solution



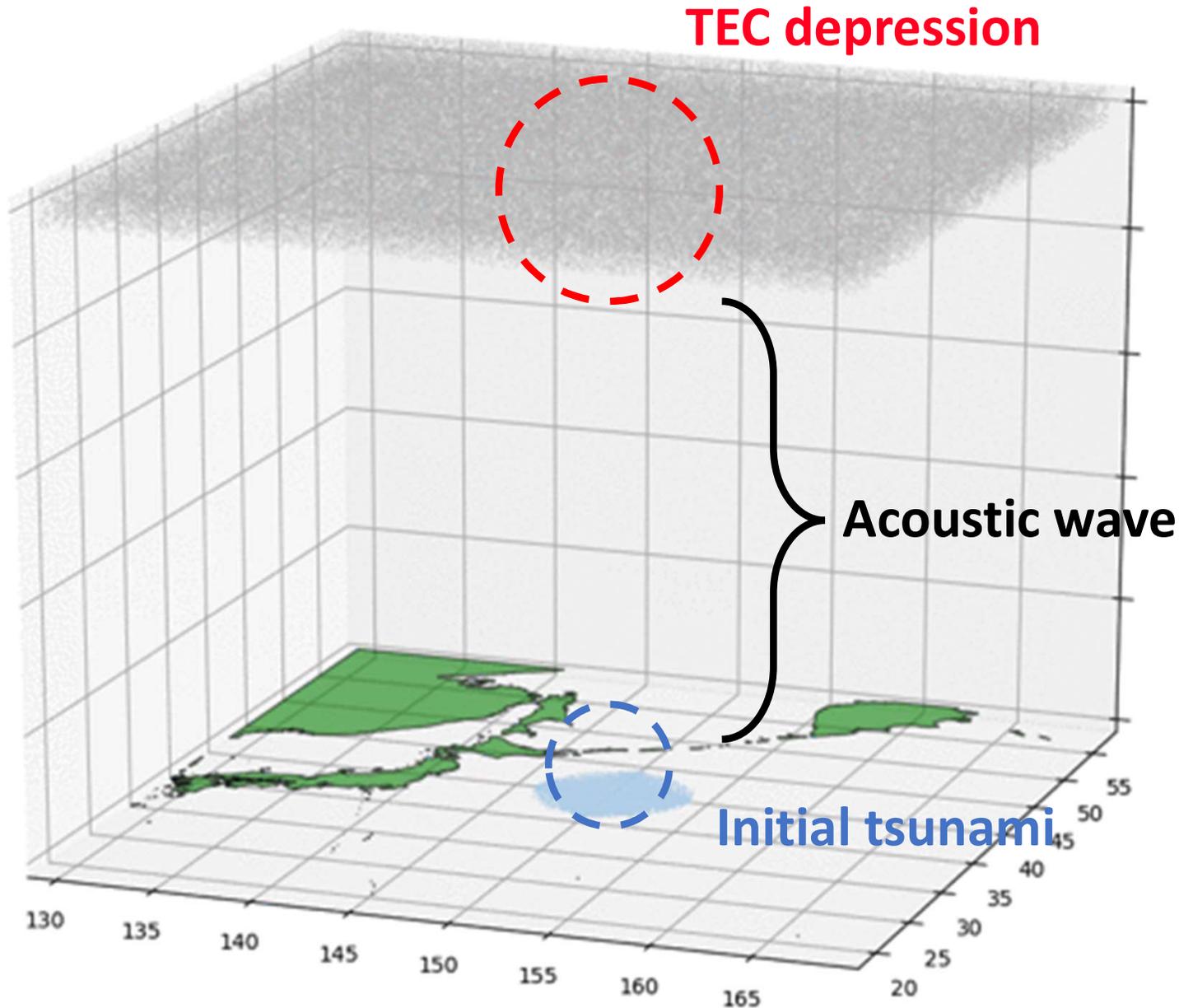
A possible solution



A possible solution



A possible solution

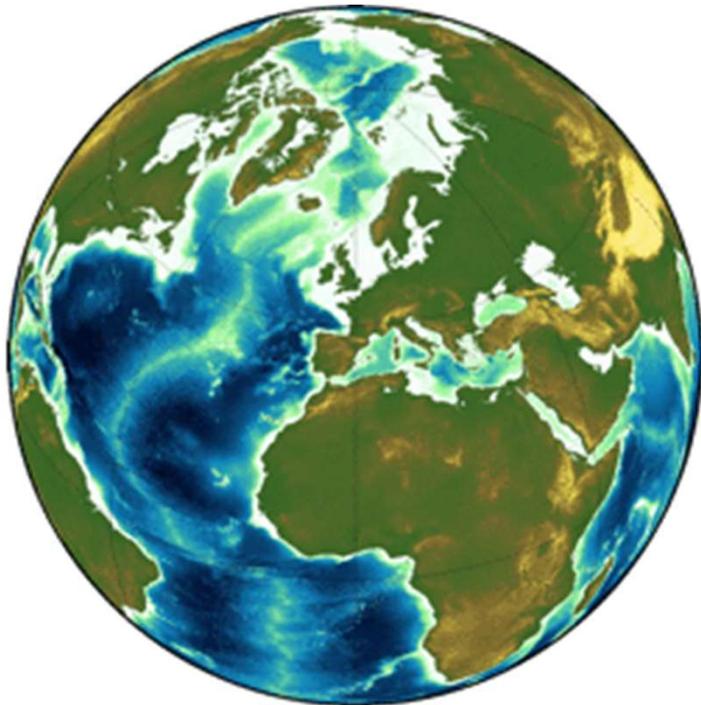


GPS data



Statistics

1 Background



2 Application of statistics

3 Results

4 Future step

Our approach

Data preprocessing

- Low pass filter
- Outlier detection

Surface fitting

- GP regression
- Uncertainty evaluation

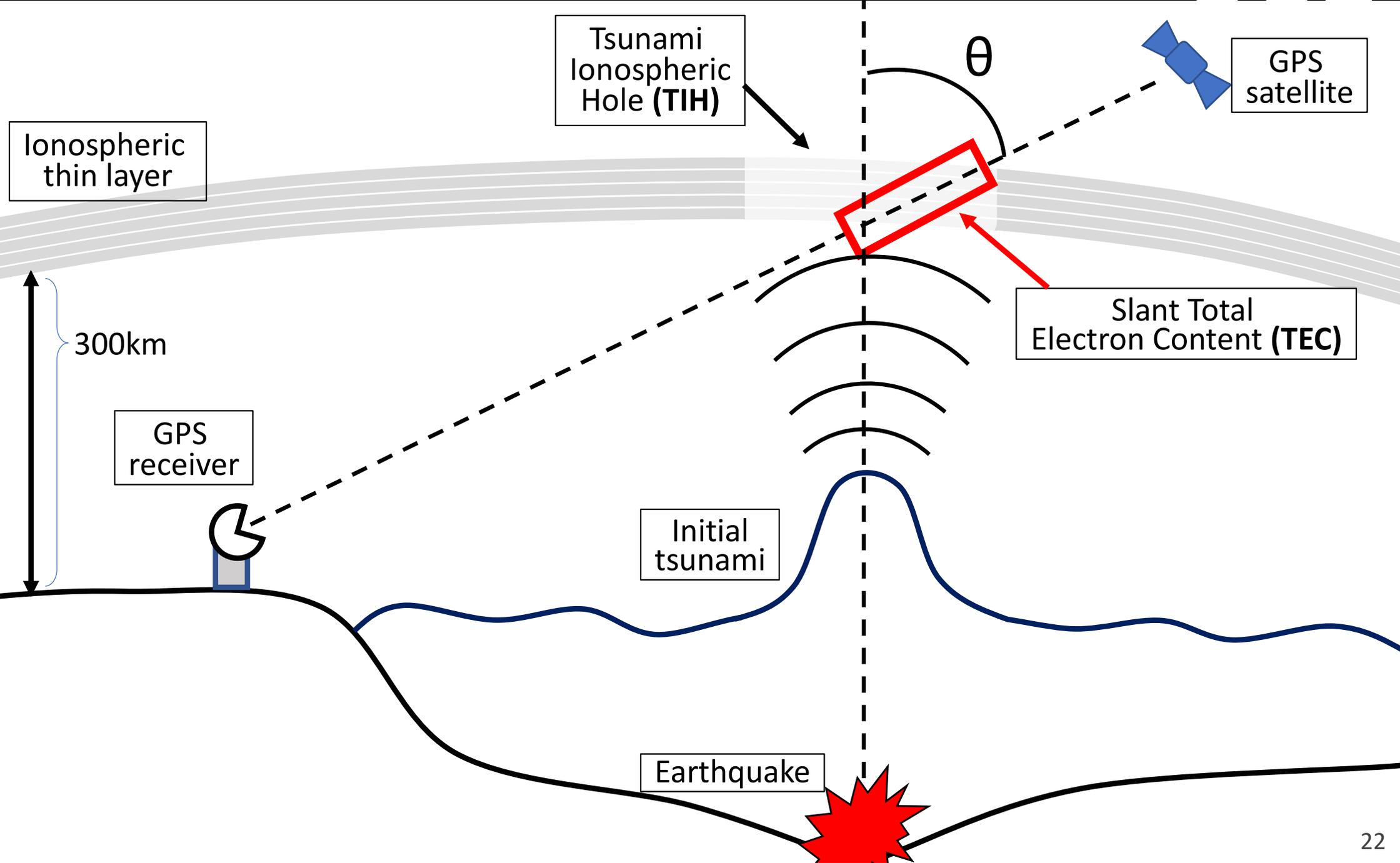
Analysis

- Depression propagation
- Overlapping with tsunami

Proposal for
a new measure

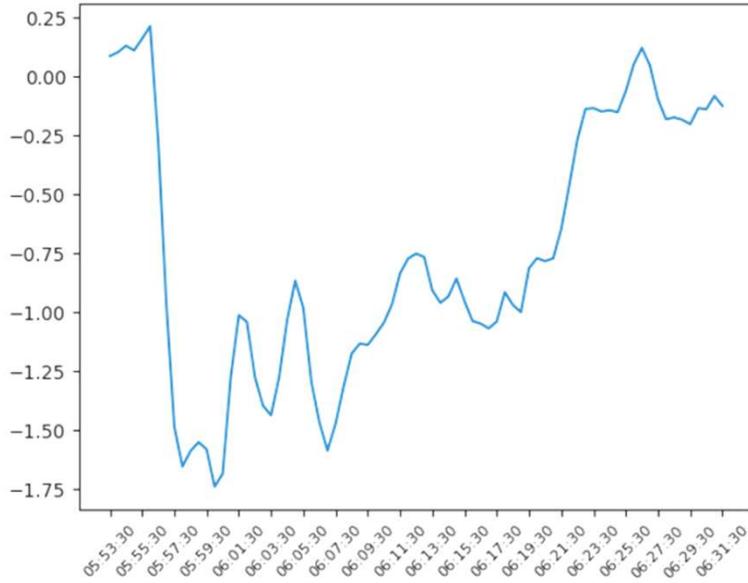
- Volume of TIH
- Early warning system

[Redisplaying] Detected data

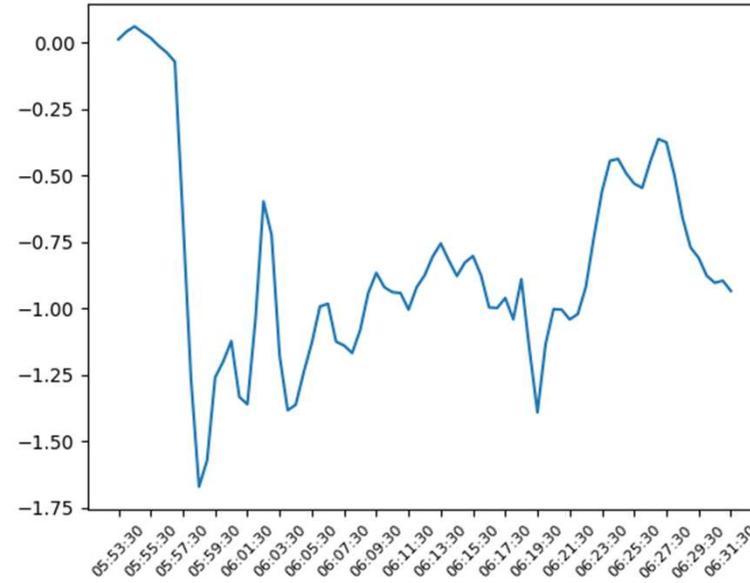


Detected data

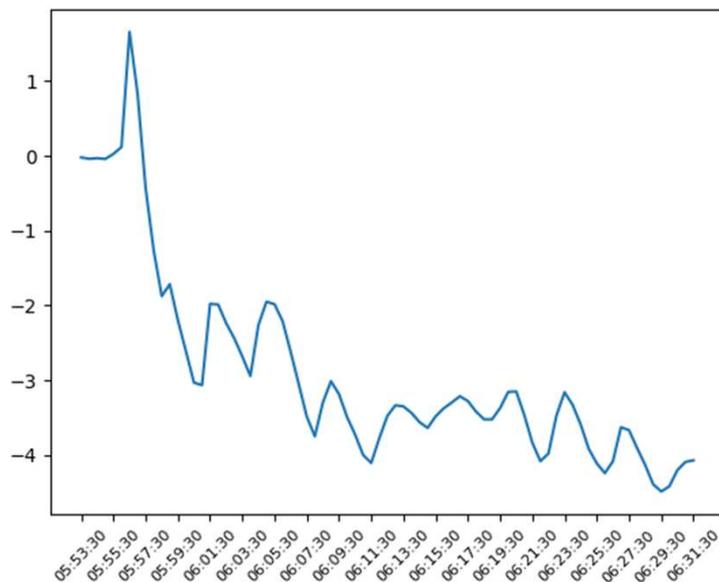
Receiver 564



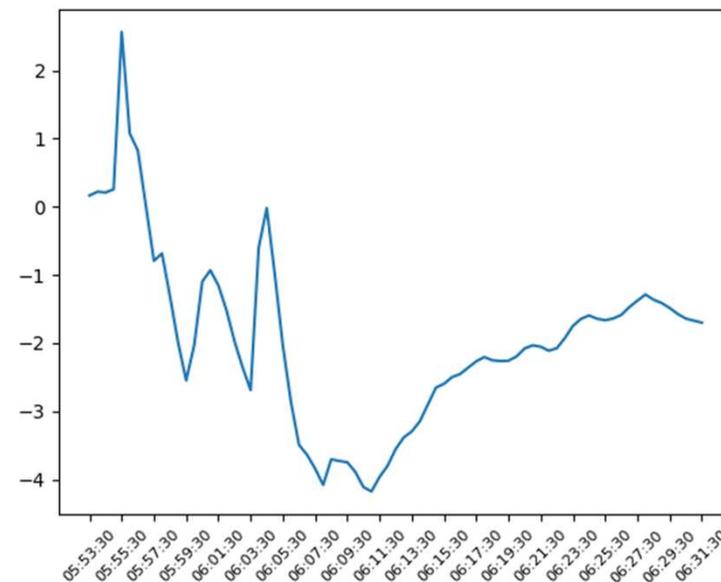
Receiver 565



Receiver 610



Receiver 3025



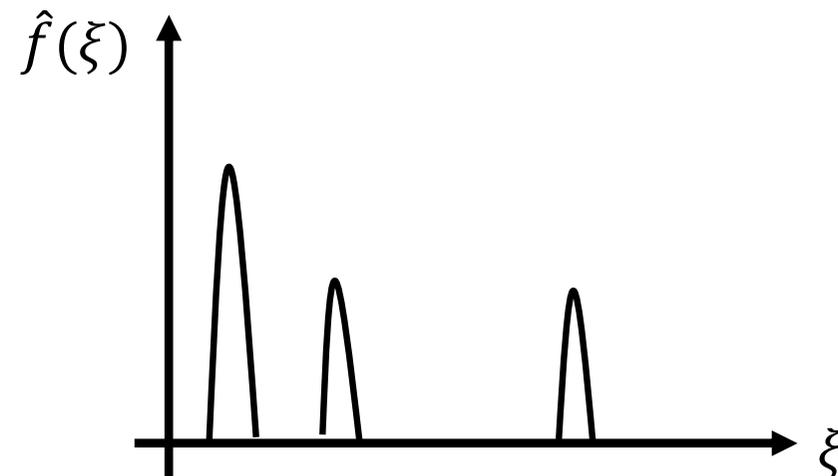
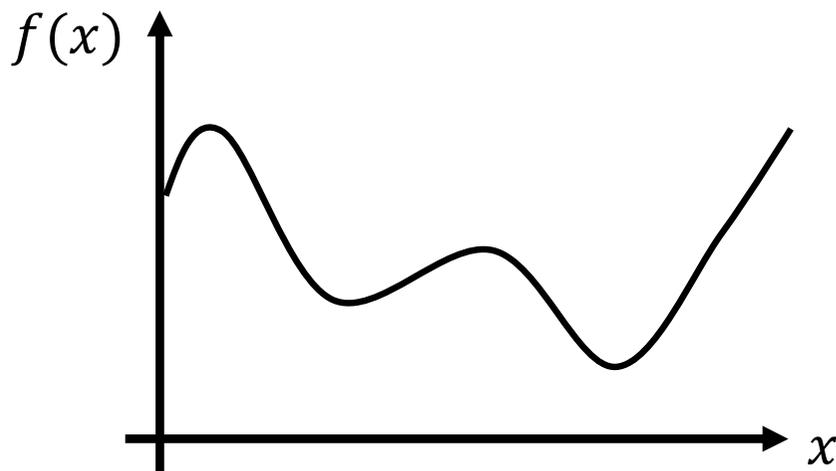
Satellite 26

This satellite is the one which is located near the epicenter. All the graphs displayed here are electron density depression data captured by the satellite 26. The data can be purchased by the Japanese government.

Fourier Transform

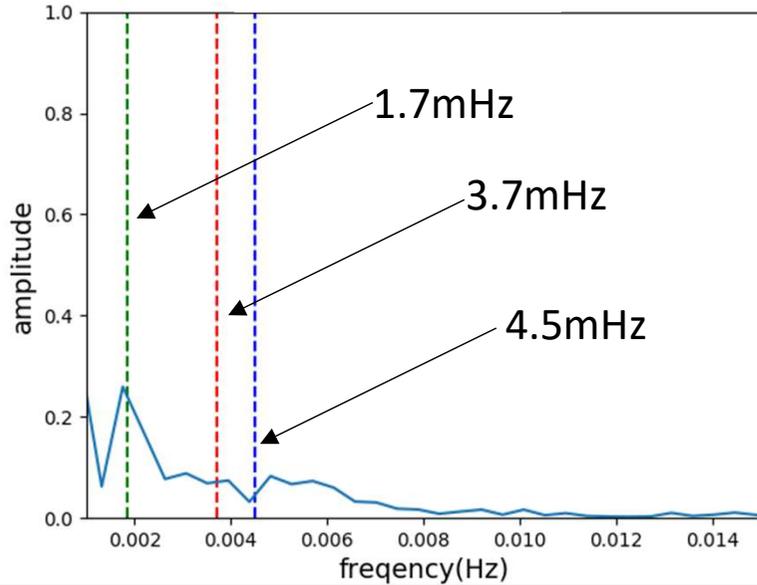
The Fourier Transform transforms a function f into a function \hat{f} that describes the frequencies contained in f . In other words, the Fourier Transform is a frequency domain representation of f .

$$\hat{f}(\xi) := \int_{-\infty}^{\infty} f(x) e^{-2\pi i x \xi} dx$$

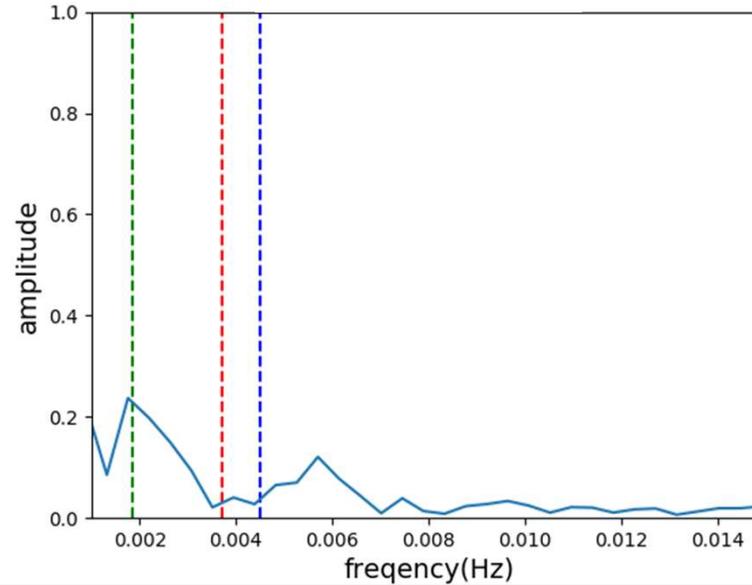


Detected data

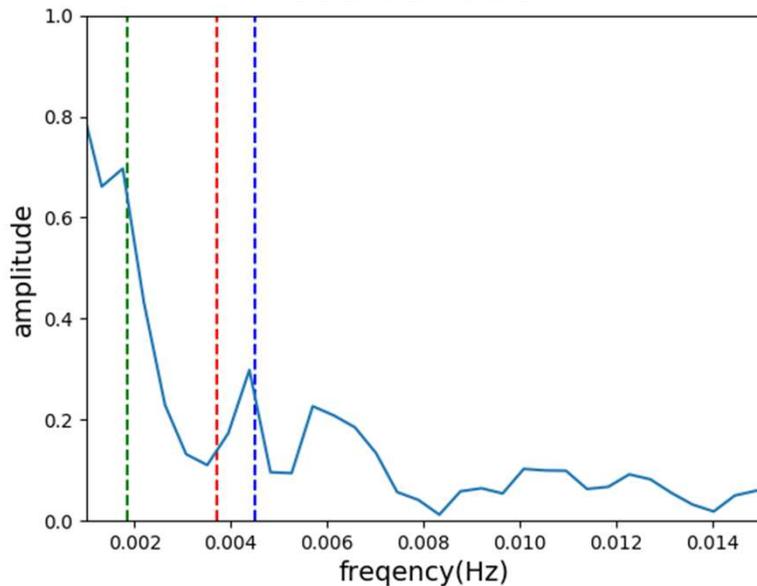
Receiver 564



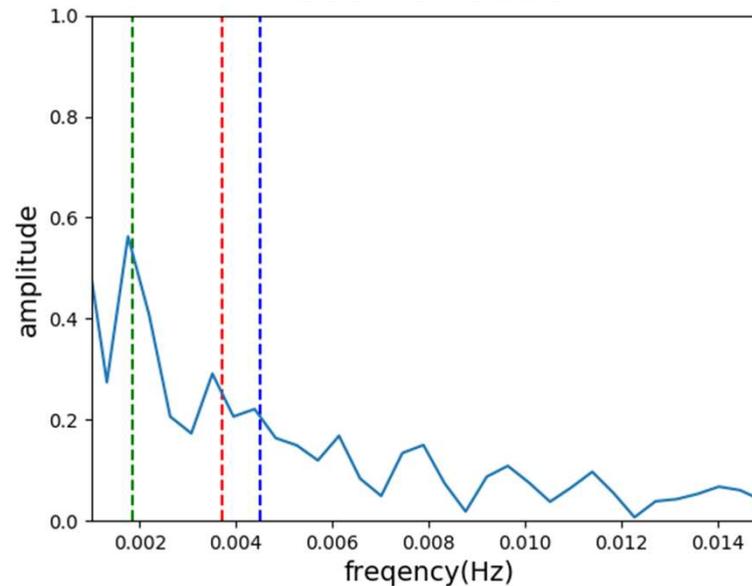
Receiver 565



Receiver 610



Receiver 3025



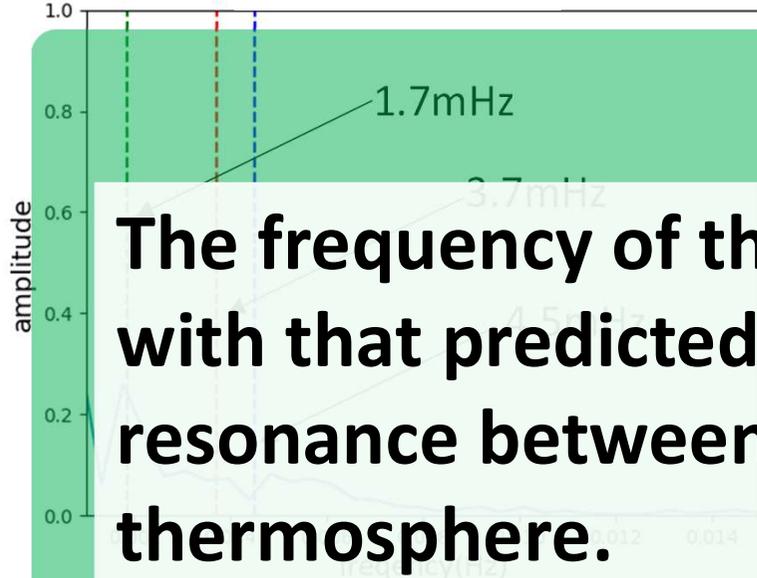
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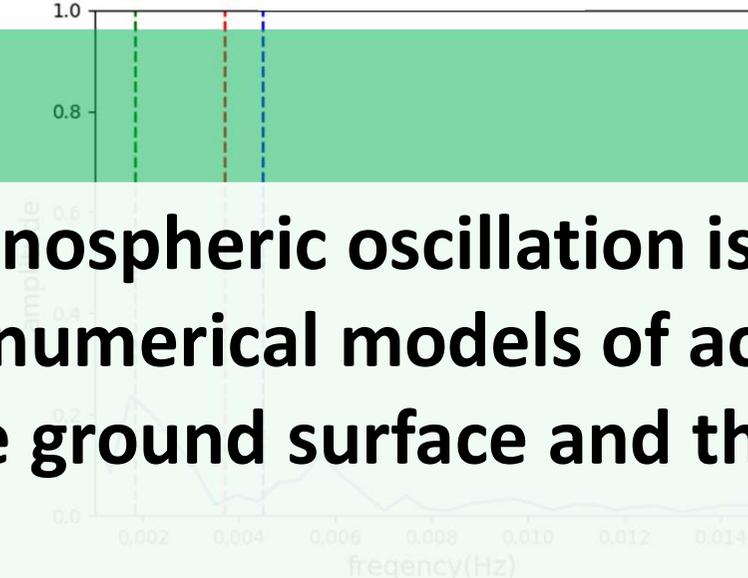
The data can be purchased by the Japanese government.

Detected data

Receiver 564



Receiver 565



Satellite 26

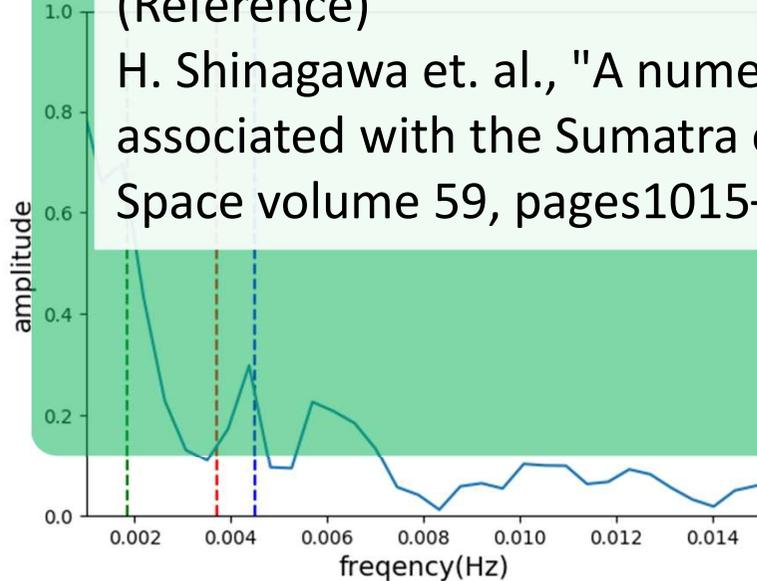
This satellite is the one which

The frequency of the ionospheric oscillation is consistent with that predicted by numerical models of acoustic resonance between the ground surface and the lower thermosphere.

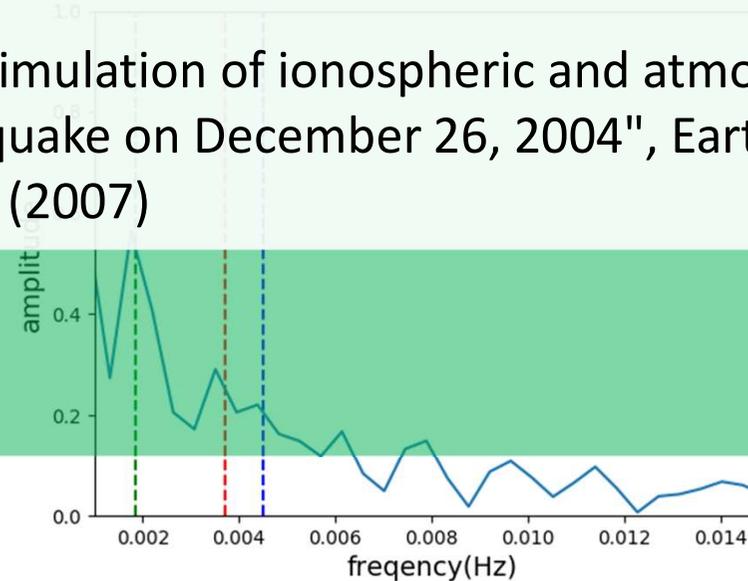
(Reference)

H. Shinagawa et. al., "A numerical simulation of ionospheric and atmospheric variations associated with the Sumatra earthquake on December 26, 2004", Earth, Planets and Space volume 59, pages1015–1026 (2007)

Receiver 610



Receiver 3025



is located near the epicenter.

The graphs displayed here

are electron

density

depression

data captured

by the satellite

26.

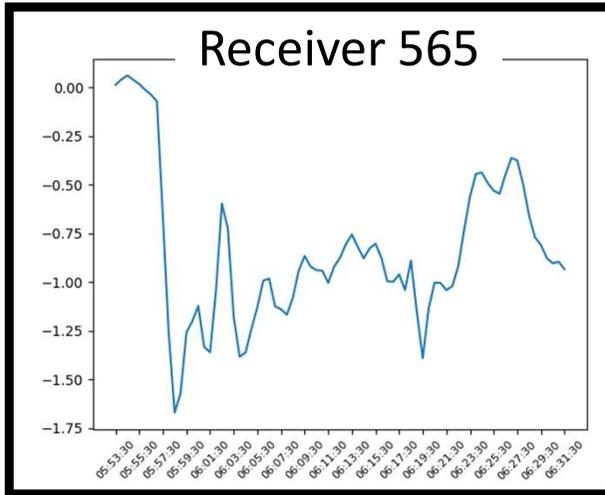
The data can

be purchased

by the

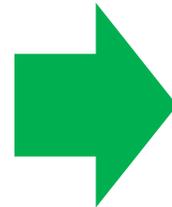
Japanese

government.



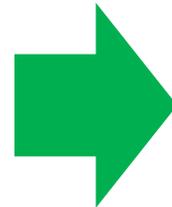
- The data shows the electron density depression.
- Many fluctuations are observed.
- High frequency components should be removed.

Fourier Transform



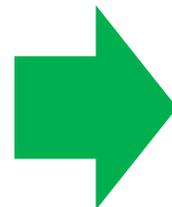
Not applicable to Real-time analysis

Linear Regression



Not applicable to Real-time analysis

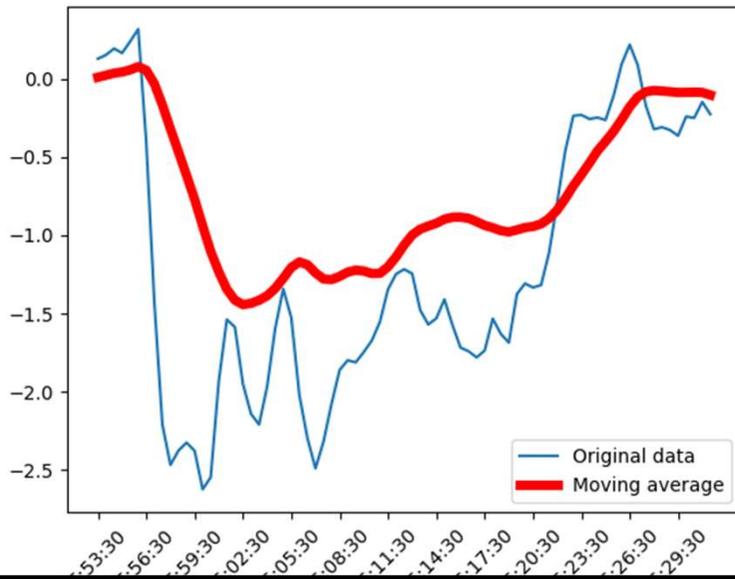
Backward Moving Average



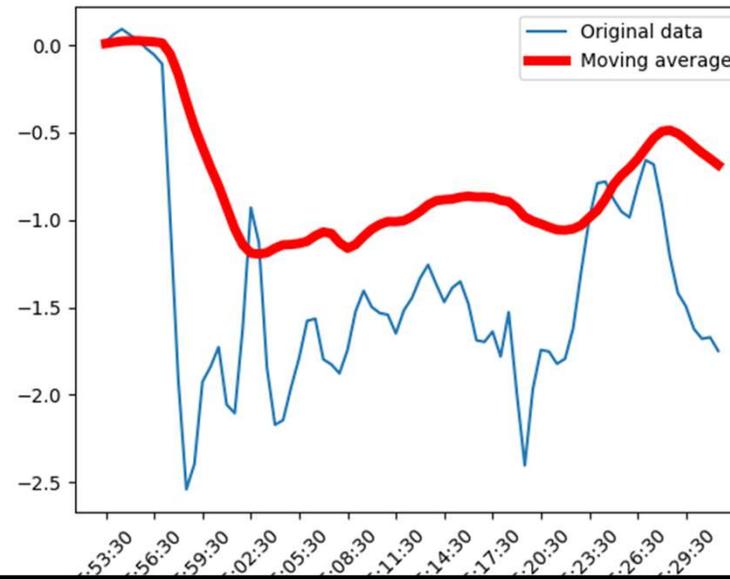
Applicable to Real-time analysis

Low-pass filtered data

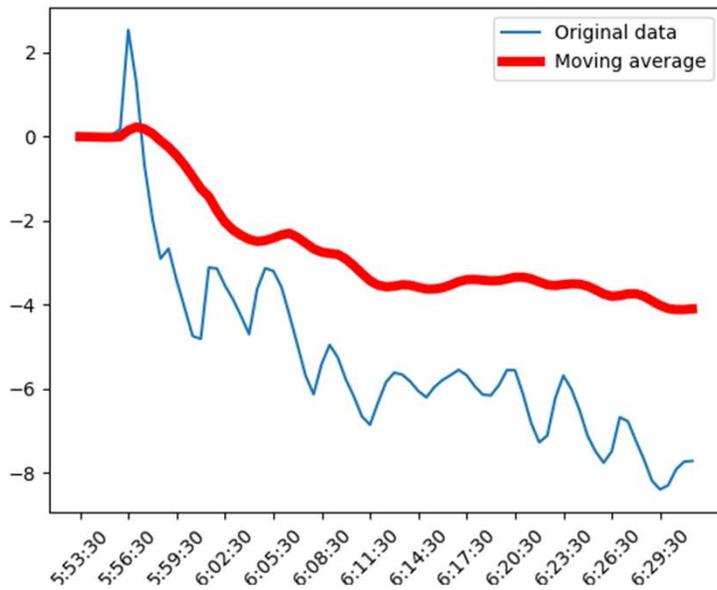
Receiver 564



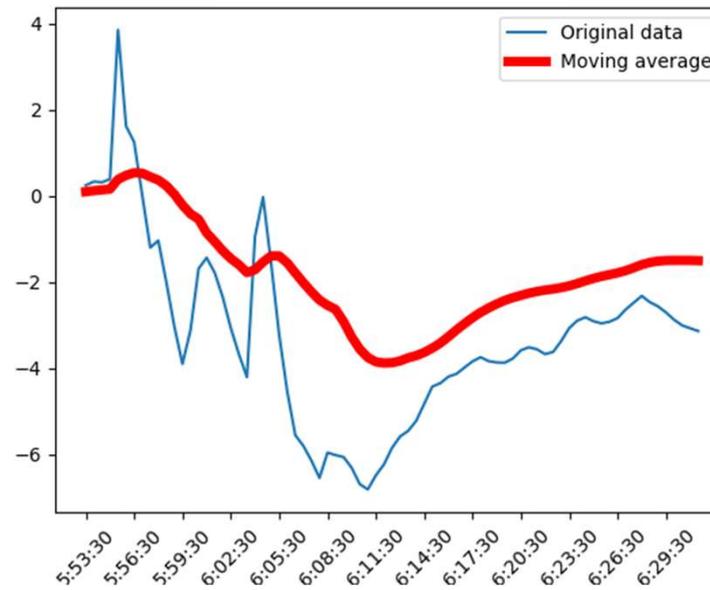
Receiver 565



Receiver 610



Receiver 3025



Low-pass filter

Only low frequency modes can be obtained through the low-pass filter based on the backward moving average.

Data preprocessing

- Low pass filter
- **Outlier detection**

Surface fitting

- GP regression
- Uncertainty evaluation

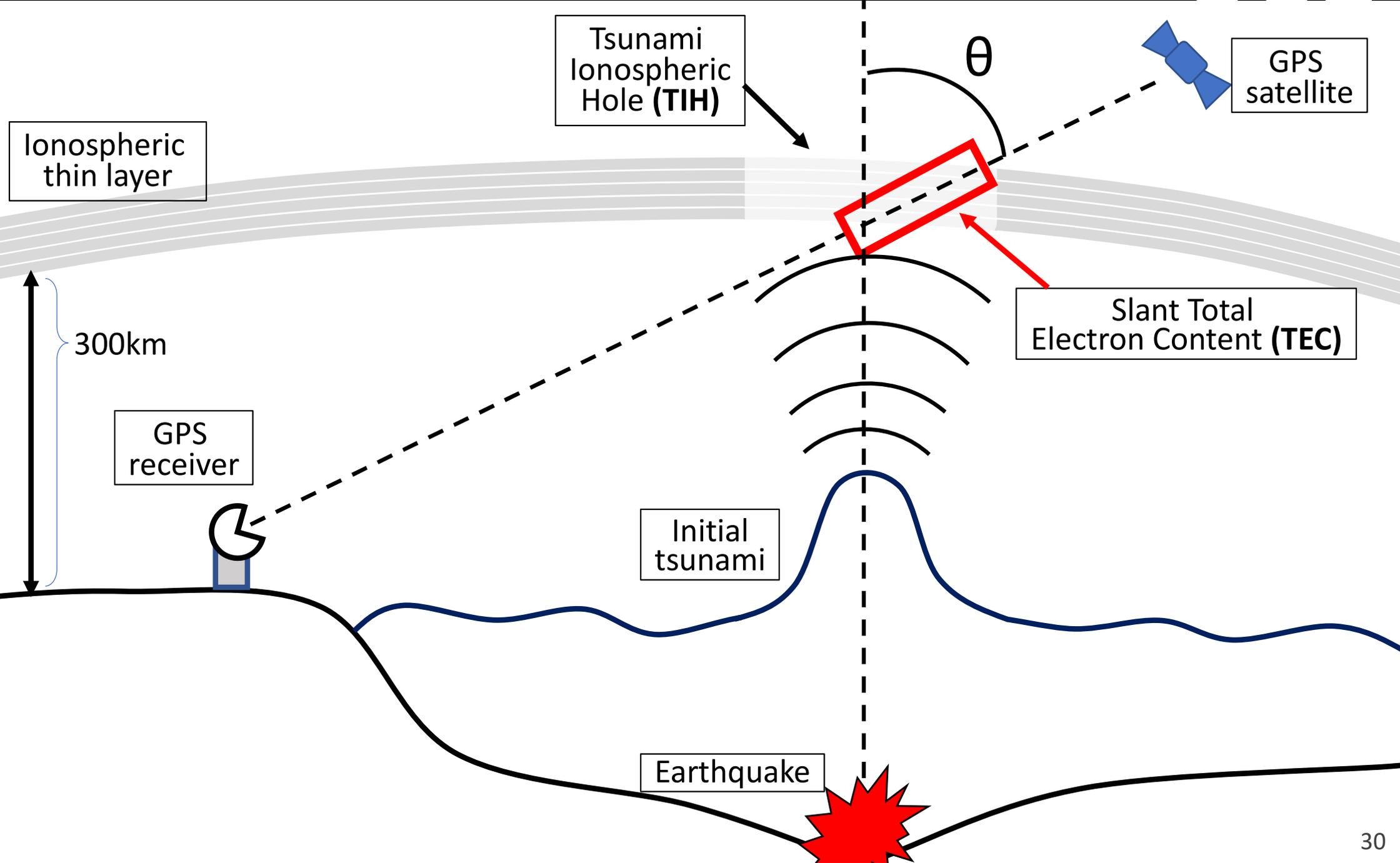
Analysis

- Depression propagation
- Overlapping with tsunami

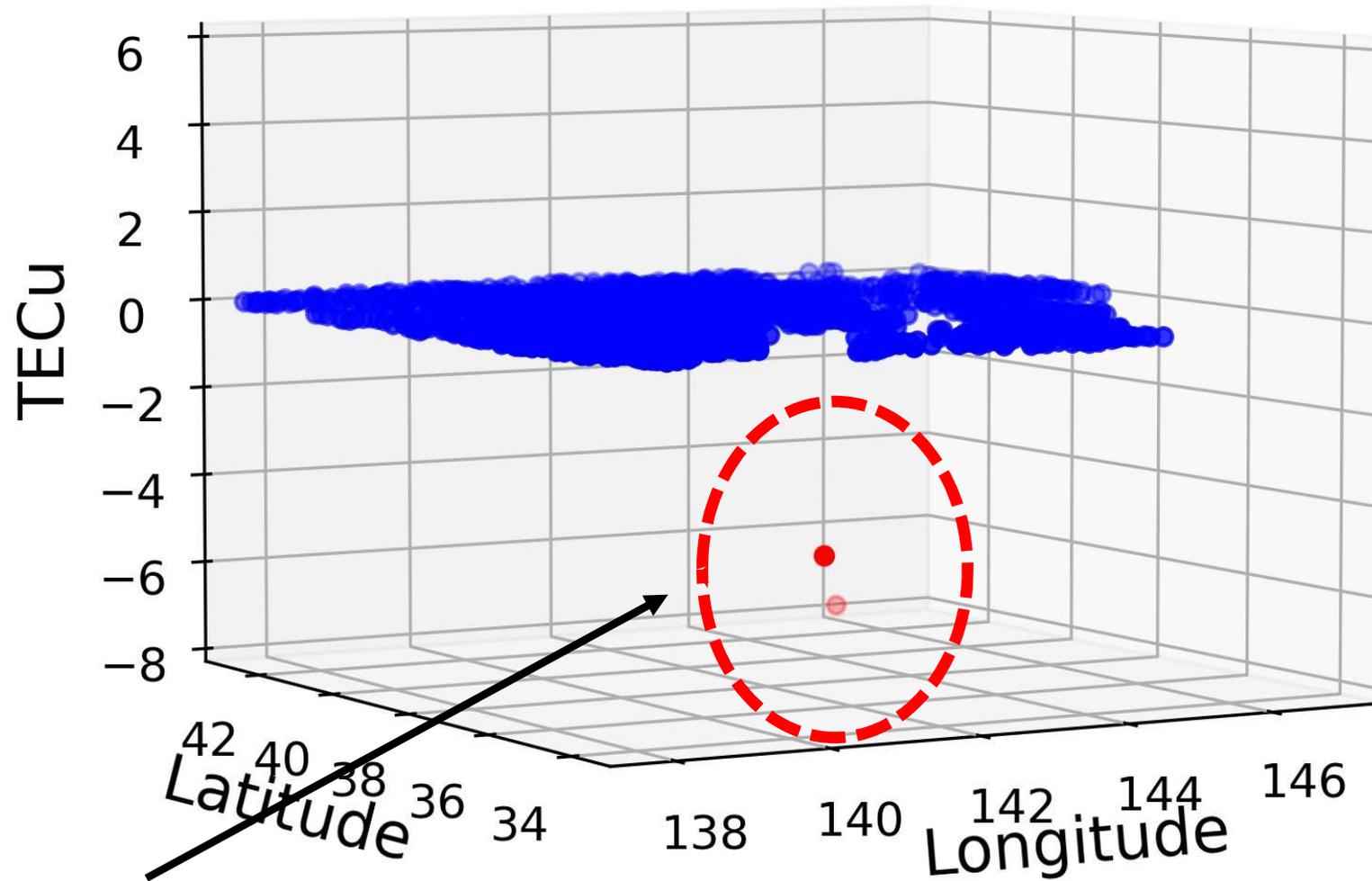
Proposal for
a new measure

- Volume of TIH
- Early warning system

[Redisplaying] Detected data



3D plot of low-pass filtered data



These two data points seem to be outlier

TECu

1. Original Data

(Low-pass filtered)

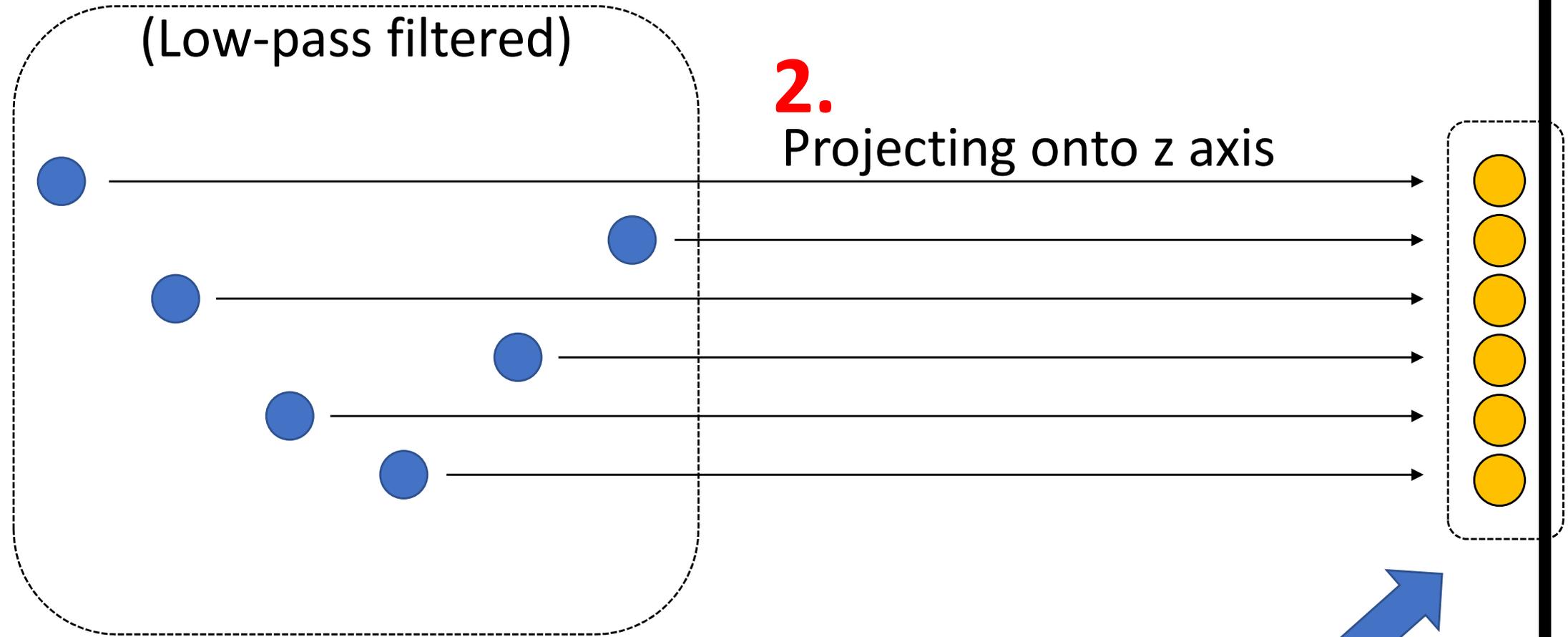
Mapping to feature space

2.

Projecting onto z axis

3.

Outlier detection based on KNN-method

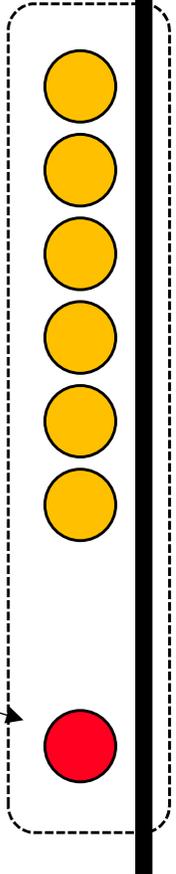


TECu

Original Data
(Low-pass filtered)

Projecting onto z axis

This point can be
detected as outlier

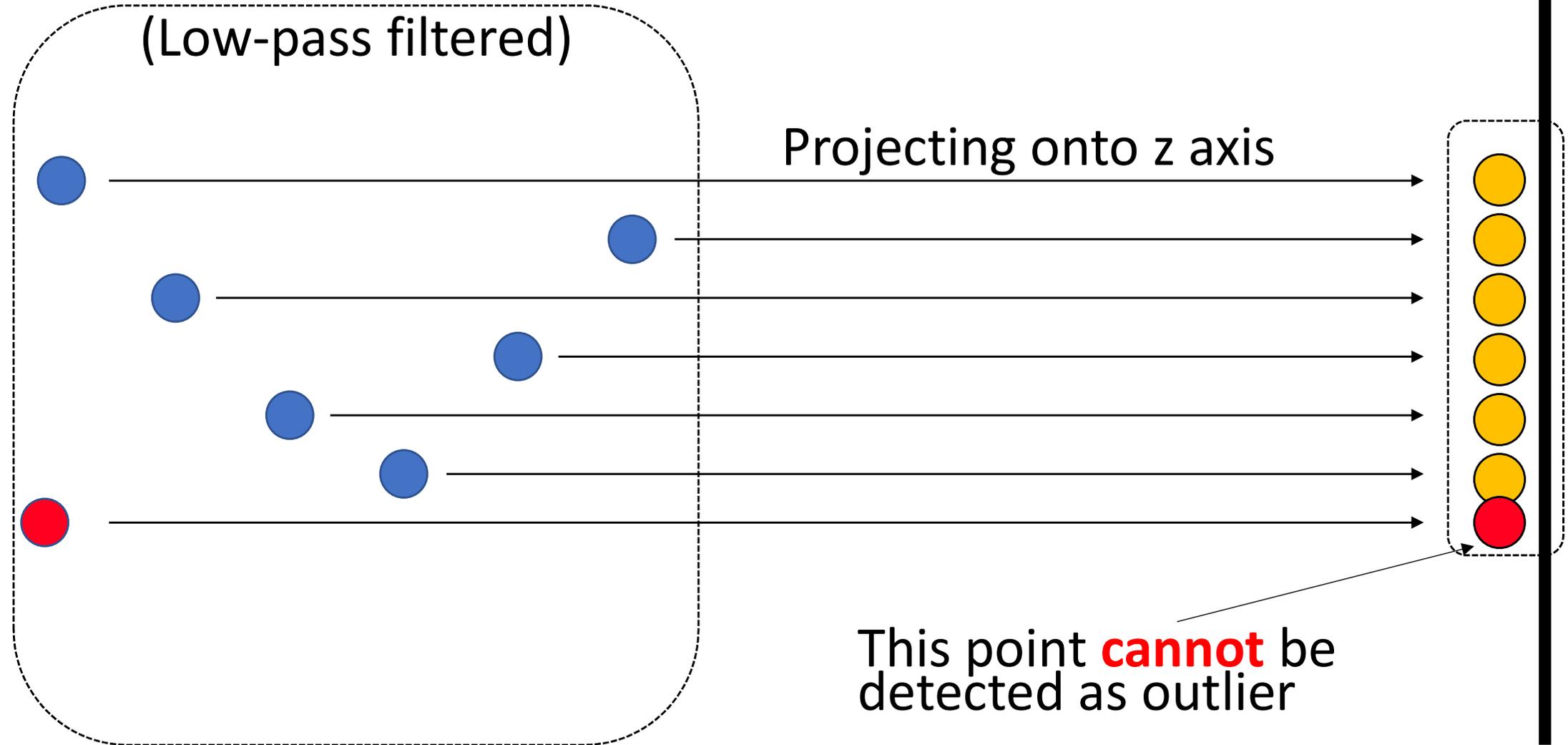


TECu

Original Data
(Low-pass filtered)

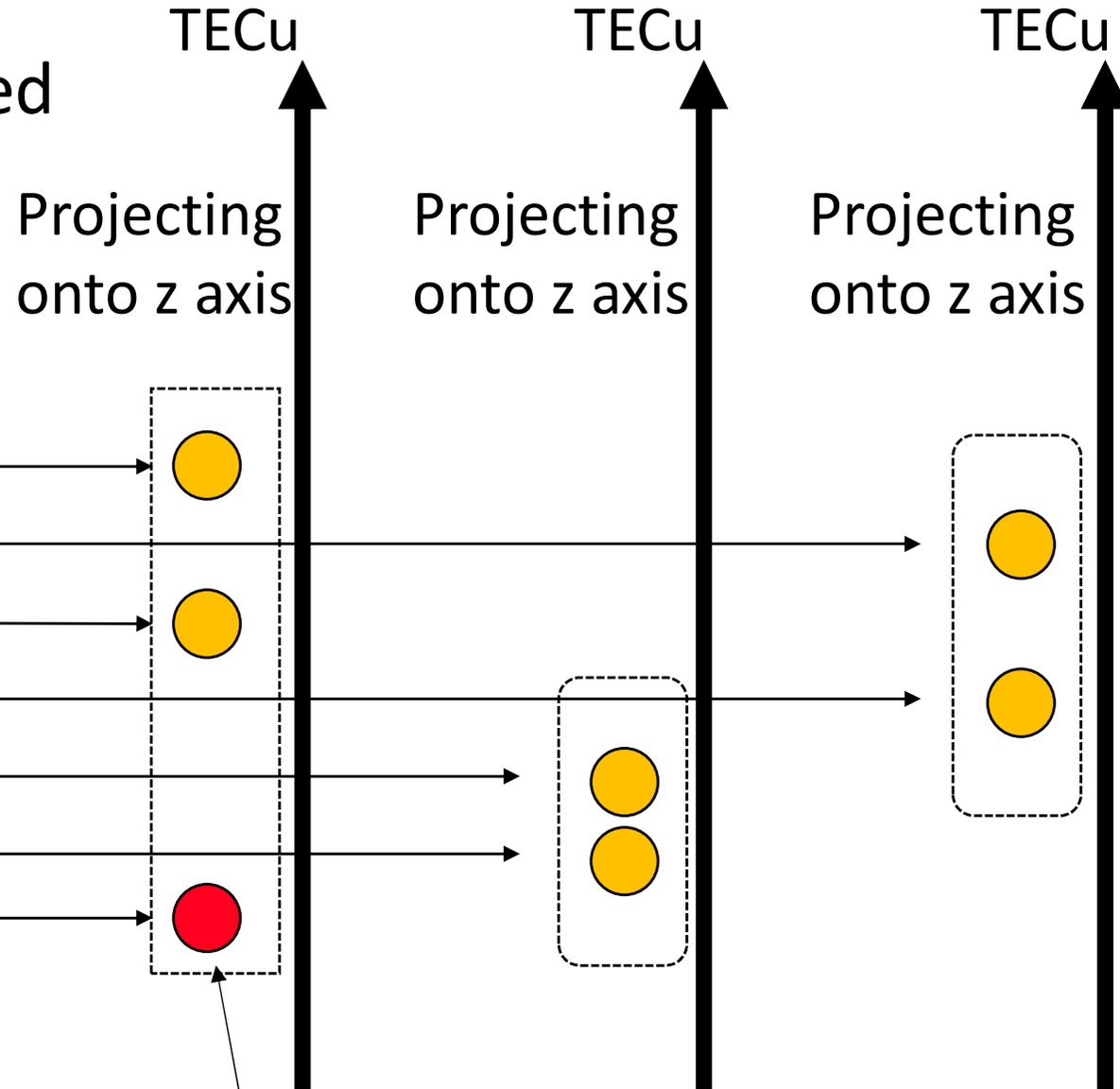
Projecting onto z axis

This point **cannot** be detected as outlier



New method

Dividing original data based on the distance



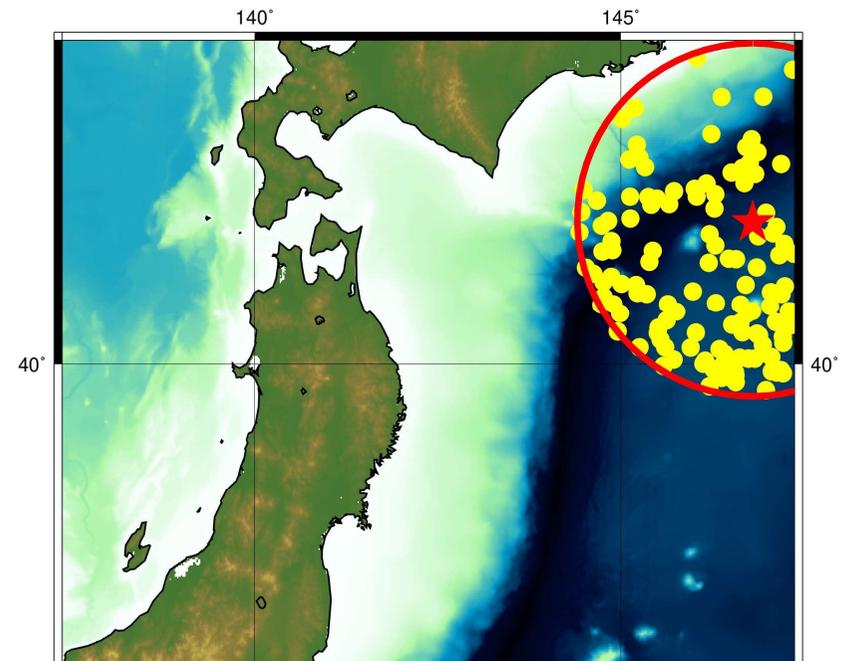
This point **can** be detected as an outlier

Outlier Detection

Input: Threshold = 3,
 n : The number of the observation points in the targeting area

```
for  $i = 1, \dots, n$  do
  Select a data point  $S_i$  from the entire targeting area
  Select all data points  $\{T\}$  located within 200km radius circle centered on  $S_i$ 
  Count the number,  $n_i$ , of  $\{T\}$  and set  $k_i = \sqrt{n_i}$ 
  for  $j = 1, \dots, n_i$  do
    Calculate the distance  $d_j$  between  $T_j$  and  $S_i$ 
  end for
  Sort  $\{T\}$  based on  $\{d\}$  in ascending order
  Calculate the TEC difference  $D_i$  between  $S_i$  and the  $k_i$ th nearest point
  if  $D_i > 3$  then
     $S_i$  is considered as an outlier
    Identify the receiver which detects  $S_i$ 
  end if
end for
```

The method can be implemented in real-time

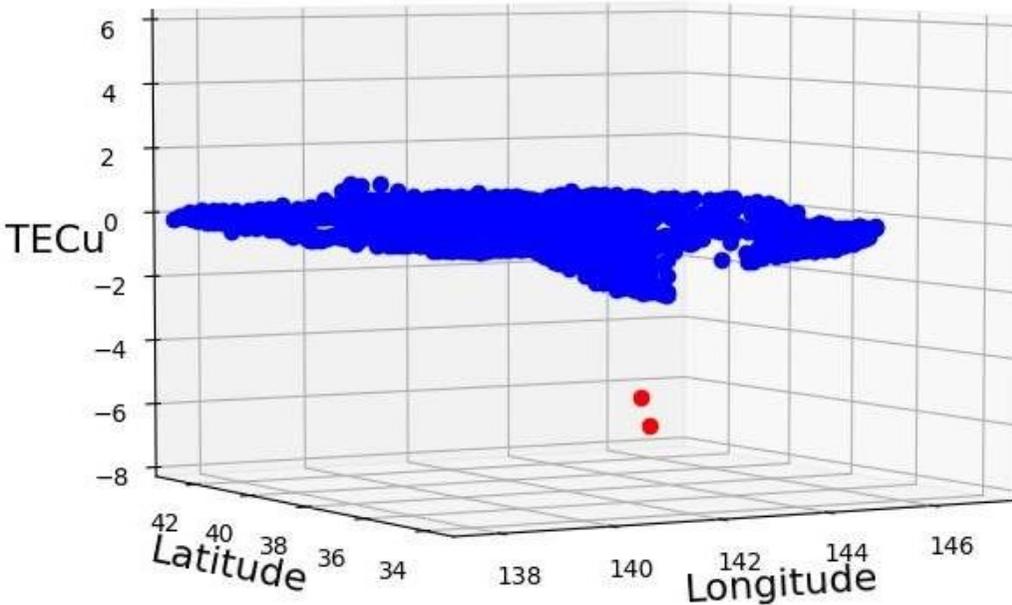


Variogram Cloud (at 6:08:00 UTC)

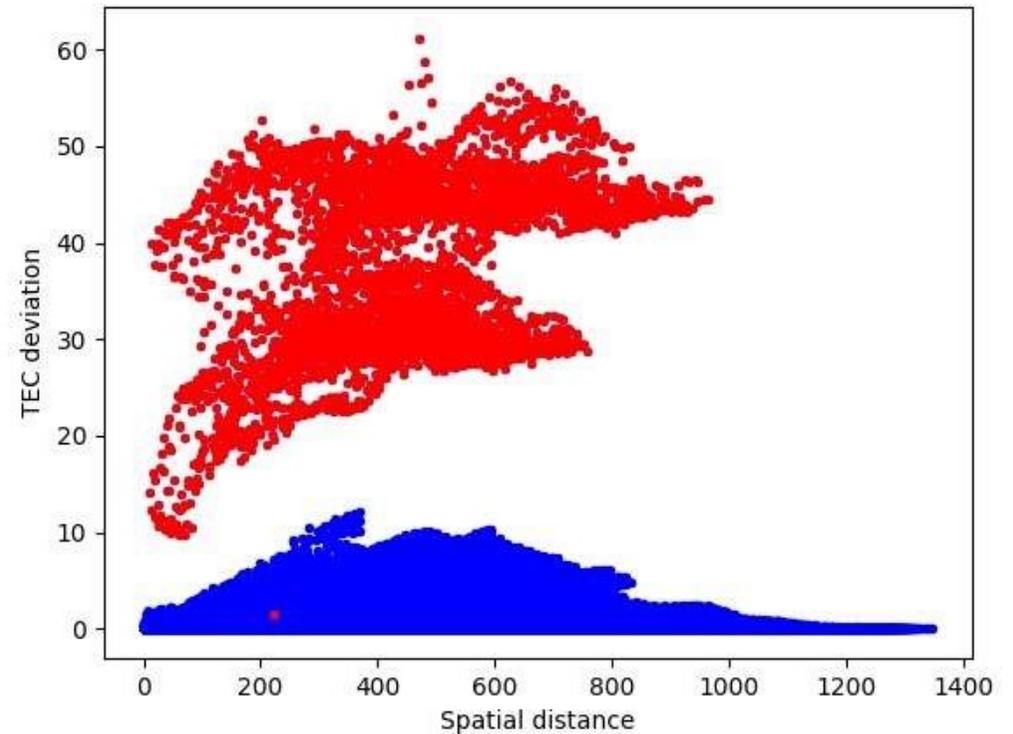
- ✓ Two outliers are detected by the method.
- ✓ These two points make the variogram cloud strange.

$$y = \frac{|\text{TEC}(x_1) - \text{TEC}(x_2)|^2}{2}$$
$$x = |x_1 - x_2|$$

3D plot

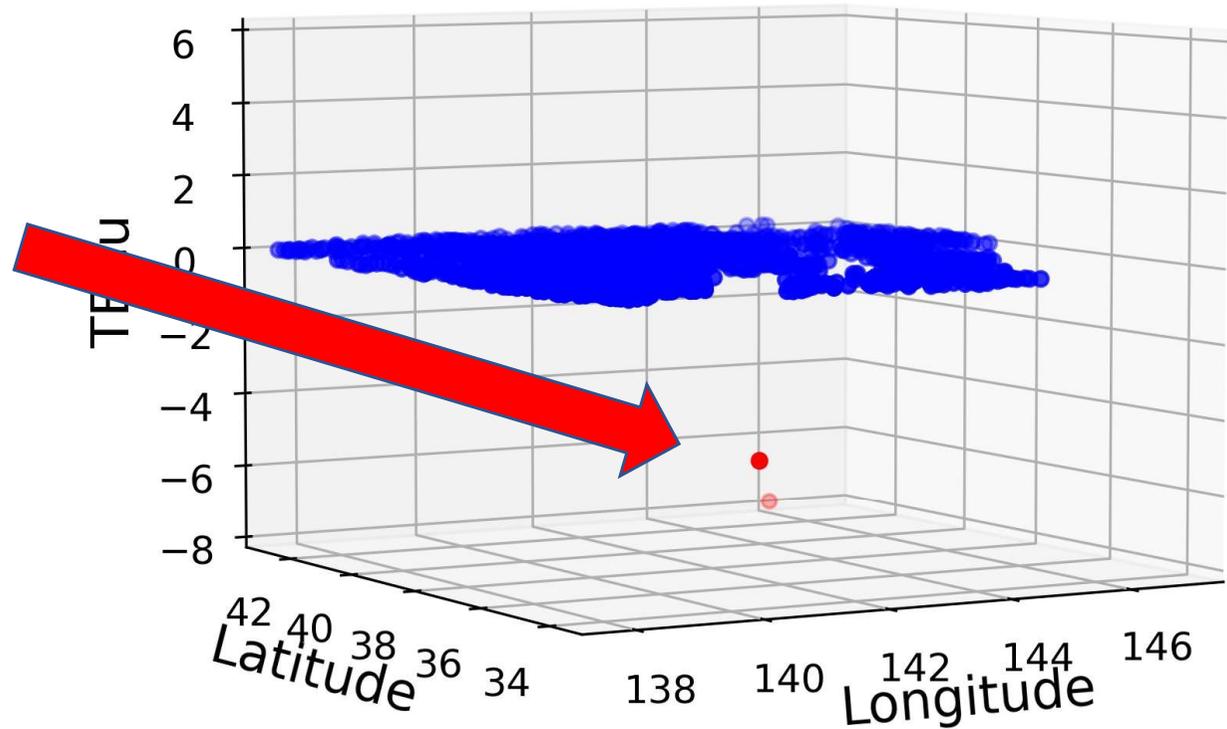


Variogram cloud

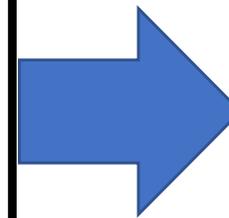


Outliers and Broken Receivers

These two outliers are detected by receiver **175** and **588**

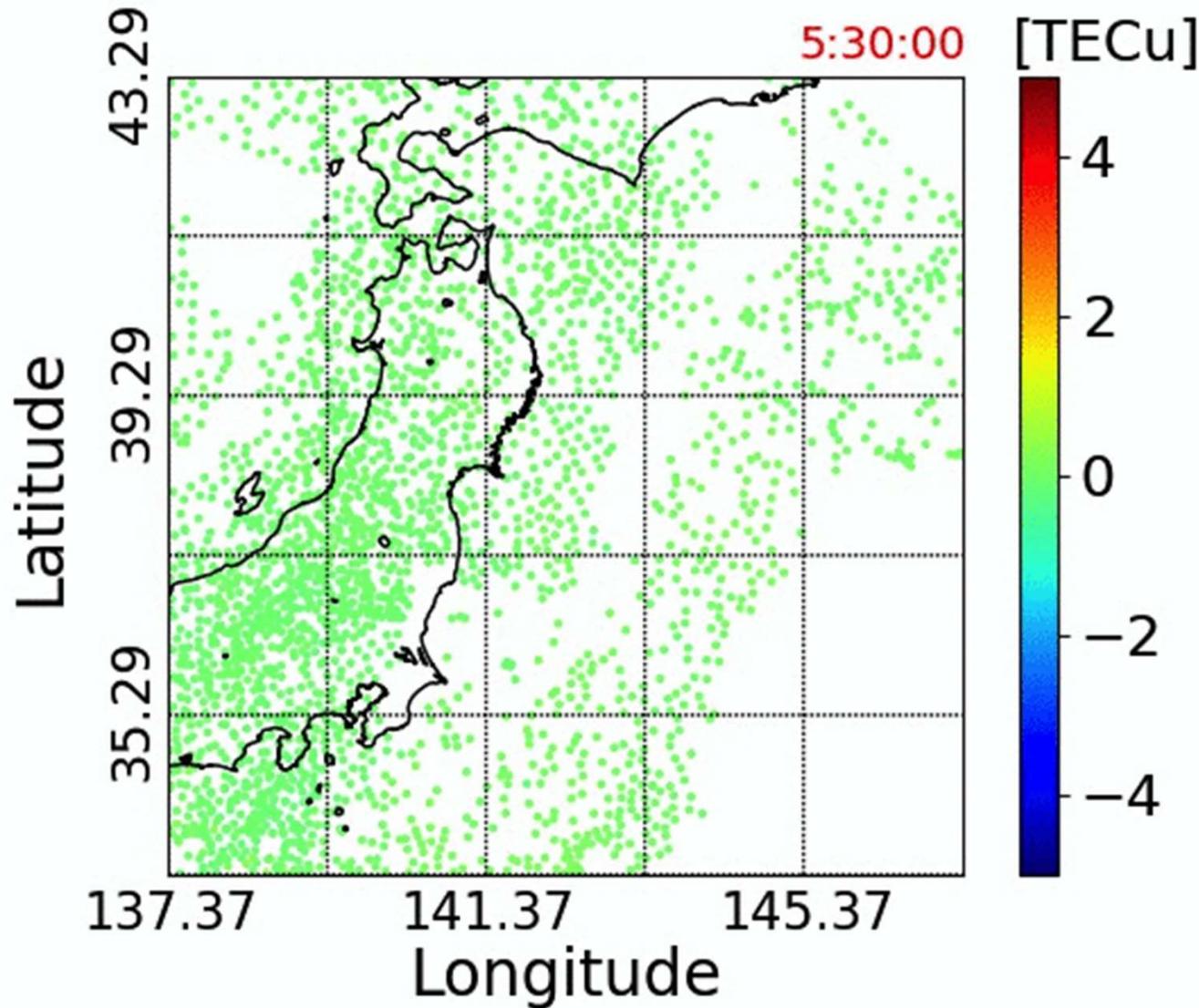


▪ GPS receiver **175**
▪ GPS receiver **588**
were broken and repaired by the Japanese government organization.



- Getting distorted can be avoided
- Broken receivers can be identified.

Original Data Captured by Satellites



Original data
from 5:30:00 to 6:16:30

The red star mark  is the epicenter

The earthquake occurred
at 6:46:30

Data points are **not uniformly** distributed.

The electron density depression is **anisotropic**.

Detected measurement points are **moving**.

Network is **sparse** in some areas (other than Japan) .

Our approach

Data preprocessing

- Low pass filter
- Outlier detection

Surface fitting

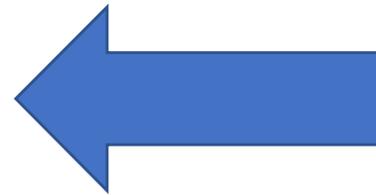
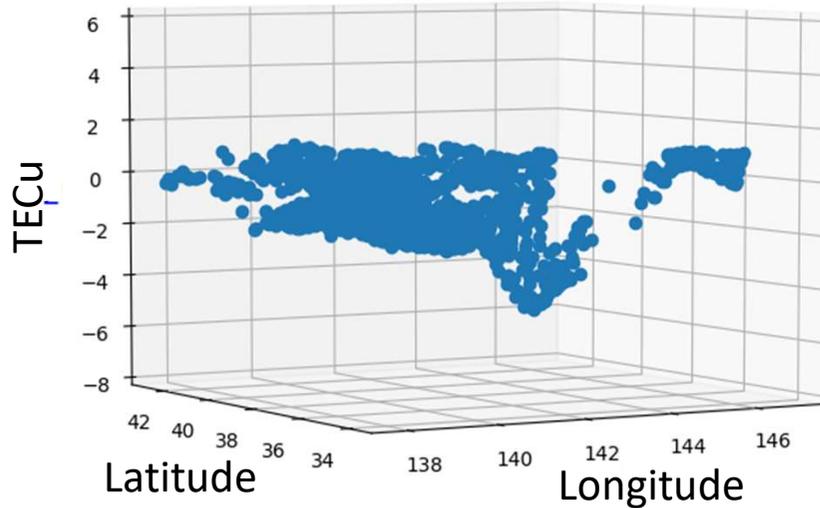
- **GP regression**
- Uncertainty evaluation

Analysis

- Depression propagation
- Overlapping with tsunami

Proposal for a new measure

- Volume of TIH
- Early warning system



Surface fitting

Matérn Kernel covariance function

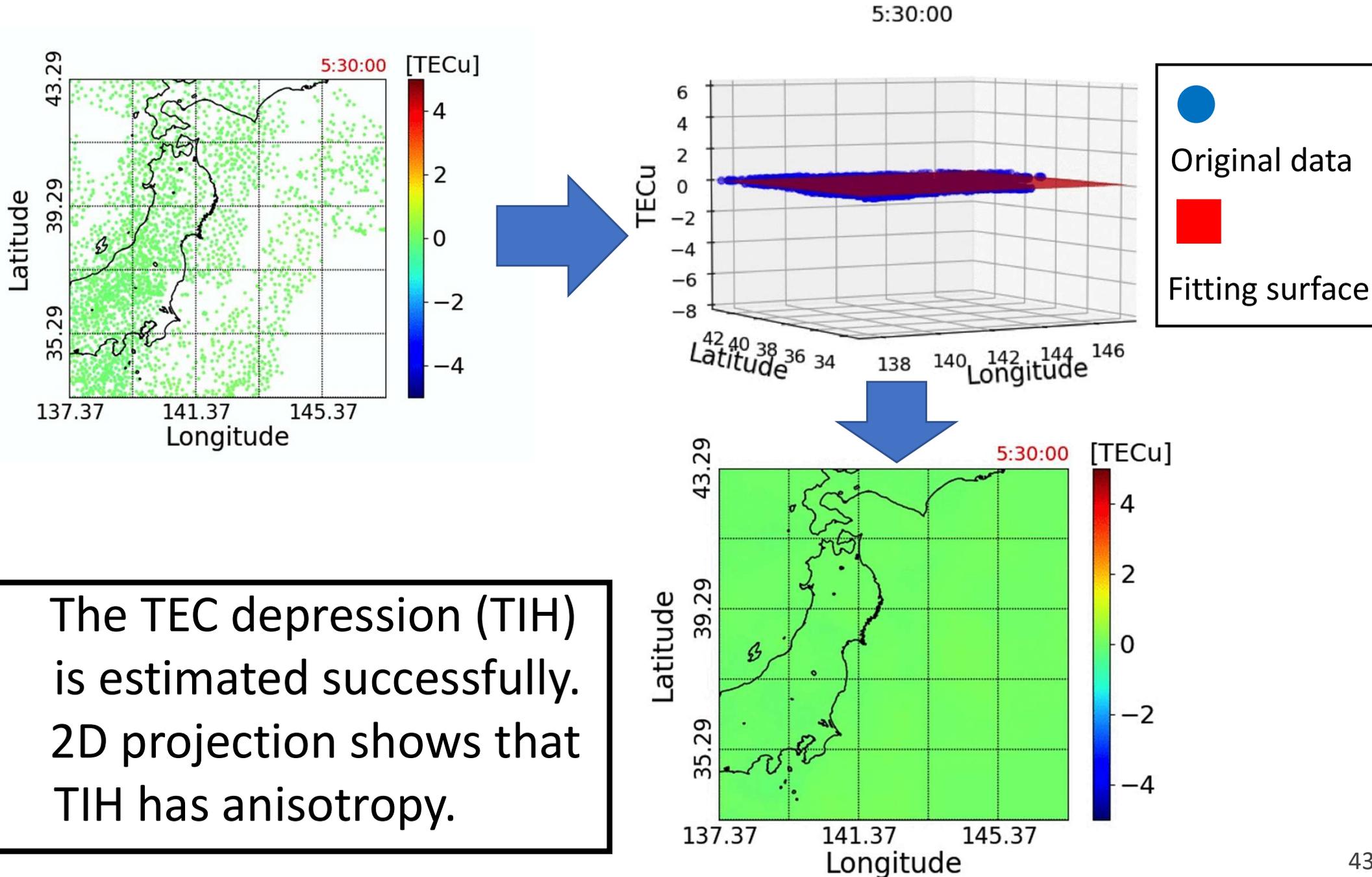
$$k_{\nu}(\mathbf{x}_p, \mathbf{x}_q) = \frac{2^{1-\nu}}{\Gamma(\nu)} \left(\frac{\sqrt{2\nu}r}{l} \right)^{\nu} K_{\nu} \left(\frac{\sqrt{2\nu}r}{l} \right), \quad \text{where } r = |\mathbf{x}_p - \mathbf{x}_q|$$

$$\text{cov}(y_p, y_q) = k_{\nu}(\mathbf{x}_p, \mathbf{x}_q) + \sigma^2 \delta_{p,q} \quad \text{The smoothness depends on } \nu.$$

Advantages

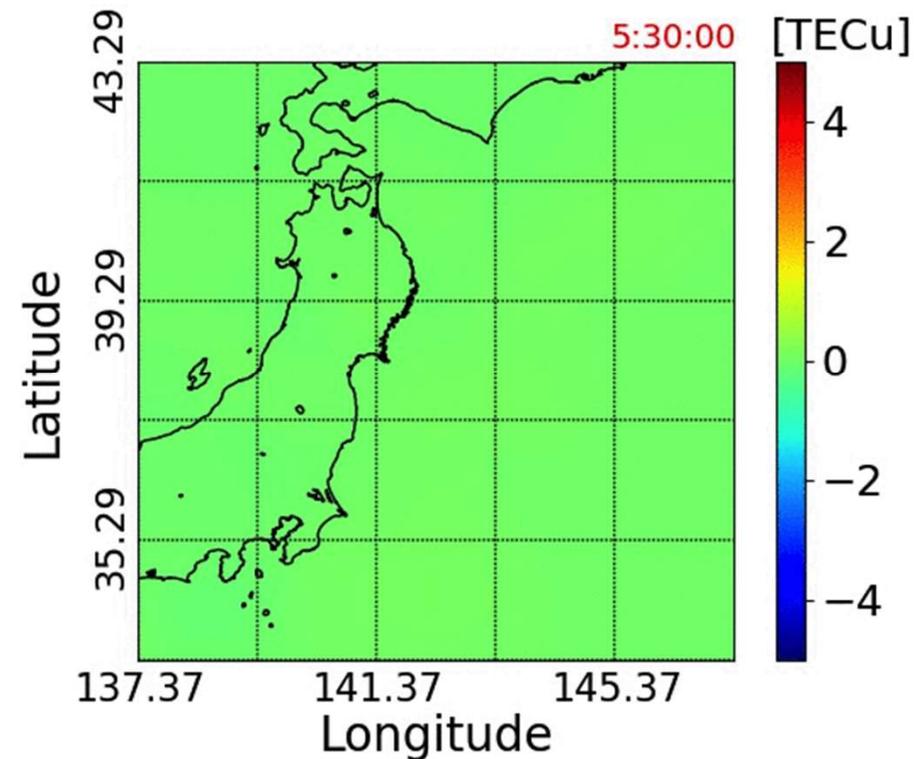
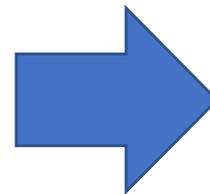
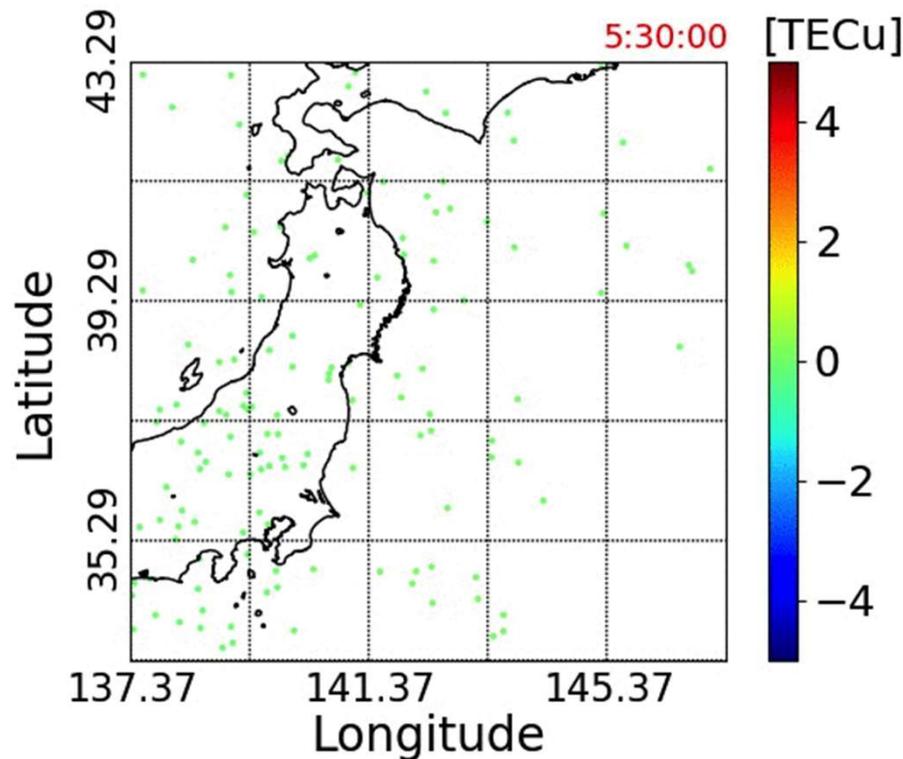
- ✓ Nonparametric method
- ✓ Evaluation of the uncertainty

Surface Fitting



Remove **95%** of GPS receivers.
Apply our method to the sparse data

TIH is estimated
successfully



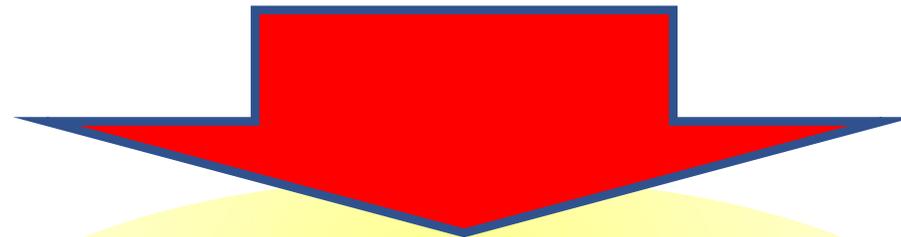
Our method can be applied to data in other countries where GPS receiver network is sparse.

Computation time

Time	5:46:30	6:00:00	6:08:00	6:16:00
Computation time (sec)	411.5	627.6	653.6	736.7

Data is detected by satellite every 30 seconds

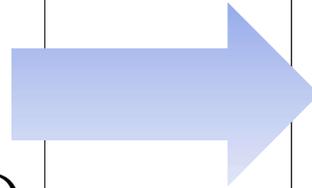
Computation time is More than 10 minutes



Acceleration is necessary to apply the method in real-time

Ordinary Gaussian Process

- Cholesky decomposition
- > Computation burden $O(n^3)$



INLA-SPDE approach

- inference with a GMRF
- > Computation burden $O(n^{3/2})$

GF

SPDE

GMRF

INLA

Fitting

Gaussian Field (GF) with Matérn covariance function represents the Electron density depression.

A certain stochastic partial differential equation (SPDE) can construct Gaussian Markov Random Field (GMRF).

GMRF is defined by sparse matrices that allow for computationally effective numerical methods.

Integrated Nested Laplace Approximation (INLA) algorithm can deal with Bayesian inference for GMRF.

Computation for the surface fitting can be accelerated.

Stochastic Partial Differential Equation

$$\frac{\sigma^2}{2^{\nu-1}\Gamma(\nu)} (\kappa\|v - u\|)^\nu K_\nu(\kappa\|v - u\|)$$

Matérn covariance function

$K_\nu(\cdot)$: modified Bessel function of the second kind

$$(\kappa^2 - \Delta)^{\alpha/2} \mathbf{x}_{GF}(\mathbf{u}) = \mathcal{w}(\mathbf{u}) \quad (1)$$

$$\begin{aligned} \mathbf{u} &\in D \subset \mathbb{R}^d \\ \alpha &= \nu + d/2 \end{aligned}$$

$$\begin{aligned} \kappa &> 0, \nu > 0 \\ \Delta &= \sum_{i=1}^n \frac{\partial^2}{\partial x_i^2} \end{aligned}$$

$$(\kappa^2 - \Delta)^{\alpha/2} \mathbf{x}_{GF}(\mathbf{u}) = \mathcal{w}(\mathbf{u})$$

solution

$$\mathbf{x}_{GF}(\mathbf{u})$$

The Gaussian Field (GF) with a Matérn function

This approach, that is the SPDE approach, uses a finite element representation to define the GF with Matérn function as a linear combination of basis functions defined on a triangulation of the domain D .

The basis function representation

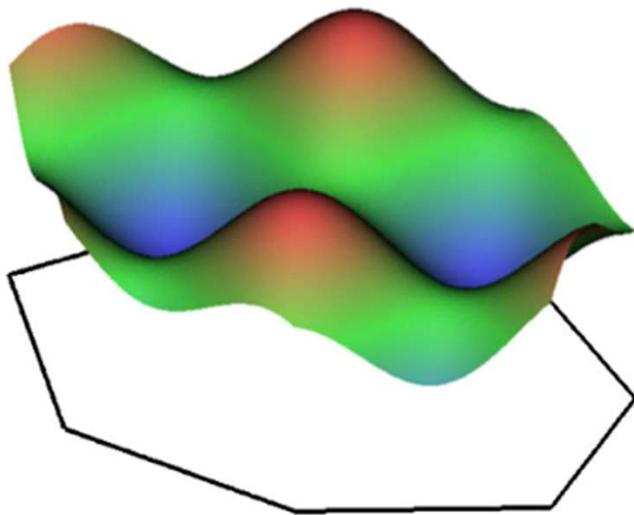
To construct a discrete approximation

$$\mathbf{x}_{GF}(\mathbf{u}) = \sum_{i=1}^n \psi_i(\mathbf{u}) \omega_i \quad (2)$$

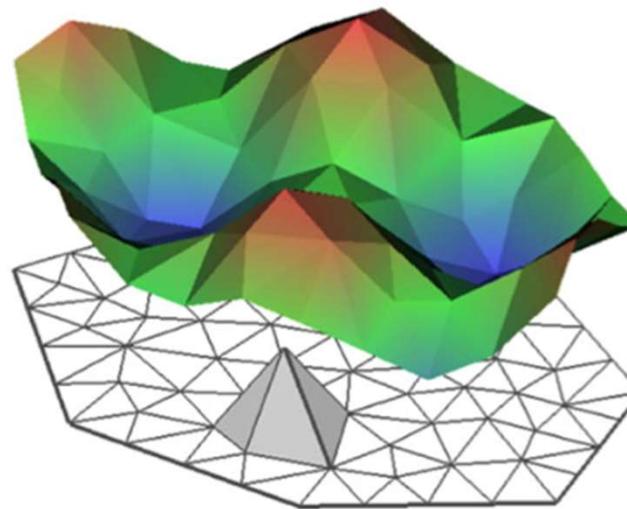
$\psi_l(\mathbf{u})$: basis functions

ω_l : weights

n : the total number of vertices,



(a) Spatial random field



(b) corresponding finite element representation of the spatial random field

- The height of each triangle is given by the weight.
- The values in the interior of the triangle are determined by linear interpolation.

Gaussian Markov Random Fields (GMRF)

A spatial process that models the spatial dependence of data observed on areal units, such as regular grid, lattice structure or geographic regions and has Markovian property (See the next slide sheet).

- An n-dimensional GMRF with mean μ and symmetric and positive definite precision matrix Q , is expressed as

$$\mathbf{x}_{GMRF} \sim \mathcal{N}(\mu, Q^{-1})$$

- The density of \mathbf{x}_{GMRF} is

$$\begin{aligned} \pi(\mathbf{x}_{GMRF}) \\ = (2\pi)^{-n/2} |Q|^{1/2} \exp \left[-\frac{1}{2} (\mathbf{x}_{GMRF} - \mu)^T Q (\mathbf{x}_{GMRF} - \mu) \right] \end{aligned}$$

Markovian property

The Markovian property is related to the definition of a neighbourhood structure, in that the full conditional distribution of \mathbf{x}_i ($i = 1, \dots, n$) depends only on a few of the components of \mathbf{x} .

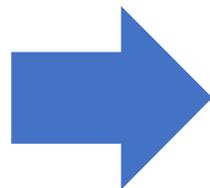
The conditional distribution of \mathbf{x}_i

$$\boldsymbol{\pi}(\mathbf{x}_i | \mathbf{x}_{-i}) = \boldsymbol{\pi}(\mathbf{x}_i | \mathbf{x}_{\delta_i})$$

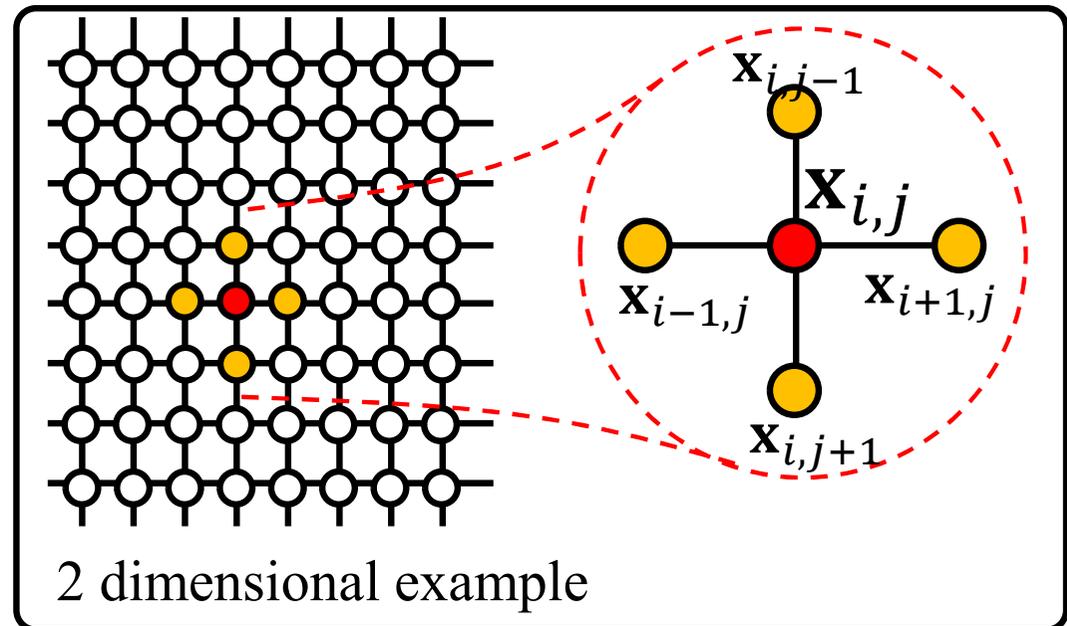
The notation \mathbf{x}_{-i} denotes all elements in \mathbf{x} except for \mathbf{x}_i

This conditional independence relation can be written as

$$\mathbf{x}_i \perp \mathbf{x}_{-\{i, \delta_i\}} | \mathbf{x}_{\delta_i}$$



The key point is that this conditional independence property is strictly related to the precision matrix \mathbf{Q} .



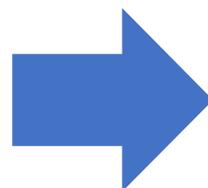
the precision matrix \mathbf{Q} .

Sparsity $j \notin \{i, \delta_i\} \Leftrightarrow \mathbf{Q}_{i,j} = \mathbf{0}$

The nonzero pattern of \mathbf{Q} is given by the neighbourhood structure of the process.

In other words, $\mathbf{Q}_{i,j} \neq \mathbf{0}$ if $j \in \{i, \delta_i\}$

Ordinary matrix
factorization
for a dense matrix
 $\mathcal{O}(n^3)$



Temporal GMRF with the sparse matrix
 $\mathcal{O}(n)$
Spatial GMRF with the sparse matrix
 $\mathcal{O}(n^{3/2})$
Spatio-temporal GMRF with the sparse
matrix
 $\mathcal{O}(n^2)$

Stochastic Partial Differential Equation

$$(\kappa^2 - \Delta)^{\alpha/2} \mathbf{x}_{GF}(u) = w(u) \quad (1) \quad \mathbf{x}_{GF}(u) = \sum_{i=1}^n \psi_i(u) \omega_i \quad (2)$$

A stochastic weak solution to the SPDE is given by requiring that

$$[\langle \varphi_i, (\kappa^2 - \Delta)^{\alpha/2} \mathbf{x}_{GF} \rangle]_{i=1, \dots, n} \stackrel{D}{=} [\langle \varphi_i, w \rangle]_{i=1, \dots, n}$$

$$[\langle \varphi_i, (\kappa^2 - \Delta)^{\alpha/2} \psi_l \rangle]_{i,l} \omega \stackrel{D}{=} [\langle \varphi_j, w \rangle]_{j=1, \dots, n} \quad \stackrel{D}{=} \text{denotes equality in distribution.}$$

for each set of **test functions** φ

When $\alpha = 2$, $\varphi = \psi$

$$[\langle \psi_i, (\kappa^2 - \Delta) \psi_l \rangle]_{i,l} \omega \stackrel{D}{=} [\langle \psi_j, w \rangle]_{j=1, \dots, n}$$

$$\underbrace{(\kappa^2 [\langle \psi_i, \psi_l \rangle])}_C + \underbrace{[\langle \psi_i, -\Delta \psi_l \rangle]}_G \omega \stackrel{D}{=} \underbrace{[\langle \psi_j, w \rangle]}_{\mathcal{N}(0, C)}$$

Stochastic Partial Differential Equation $(\kappa^2 - \Delta)^{\alpha/2} \mathbf{x}_{GF}(\mathbf{u}) = \boldsymbol{w}(\mathbf{u})$ (1)

A stochastic weak solution to the SPDE is given by

$$\mathbf{x}_{GF}(\mathbf{u}) = \sum_{i=1}^n \psi_i(\mathbf{u}) \omega_i \quad (2) \quad , \text{ where } (\kappa^2 C + G) \boldsymbol{\omega} \sim \mathcal{N}(0, C)$$

The precision of the weight, $\boldsymbol{\omega}$, is $(\kappa^2 C + G)^T C^{-1} (\kappa^2 C + G)$

As you know, C and G are sparse, but C^{-1} is not.

To obtain sparse precision matrix we replace the C^{-1} matrix with a diagonal matrix \tilde{C}^{-1} with elements

$$\tilde{C}_{i,i} = \int \psi_i(s) ds$$

The solution has the Markovian property

D. Simpson, F. Lindgren, H. and Rue, "In order to make spatial statistics computationally feasible, we need to forget about the covariance function." *Environmetrics* 23.1 65-74 (2012)

M. Cameletti, F. Lindgren, D. Simpson, H. Rue "Spatio-temporal modeling of particulate matter concentration through the SPDE approach." *ASTA Adv Stat Anal* 97:109–131 (2013)

Bayesian inference

It is possible to make use of the Integrated Nested Laplace Approximation (INLA) algorithm. It produces almost immediately accurate approximations to posterior distributions

$$\boldsymbol{\pi}(\mathbf{x}|\mathbf{y}) \propto \boldsymbol{\pi}(\mathbf{y}|\mathbf{x})\boldsymbol{\pi}(\mathbf{x})$$

$$\boldsymbol{\pi}(\mathbf{x}, \boldsymbol{\theta} | \mathbf{y}) \propto \boldsymbol{\pi}(\boldsymbol{\theta})\boldsymbol{\pi}(\mathbf{x}|\boldsymbol{\theta})\boldsymbol{\pi}(\mathbf{x}) \prod \boldsymbol{\pi}(y_i | \mathbf{x}_i, \boldsymbol{\theta})$$

$$\tilde{\boldsymbol{\pi}}_{LA}(\mathbf{x}_i | \boldsymbol{\theta}, \mathbf{y}) \propto \frac{\boldsymbol{\pi}(\mathbf{x}, \boldsymbol{\theta}, \mathbf{y})}{\tilde{\boldsymbol{\pi}}_{GC}(\mathbf{x}_{-i} | \mathbf{x}_i, \boldsymbol{\theta}, \mathbf{y})} \Big|_{\mathbf{x}_{-i} = \mathbf{x}_{-i}^*(\mathbf{x}_i, \boldsymbol{\theta})}$$

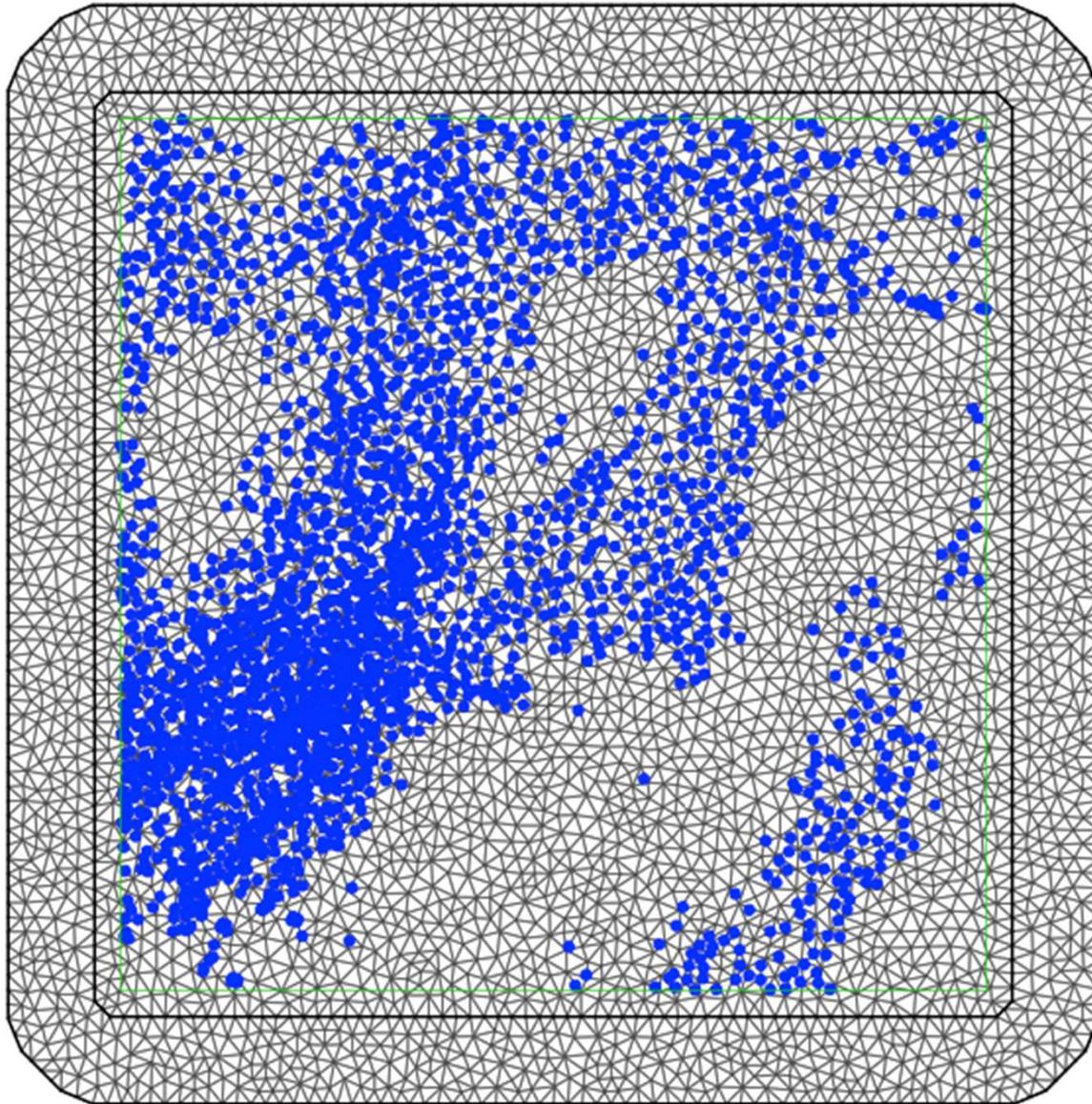
$\tilde{\boldsymbol{\pi}}_{GC}(\mathbf{x}_{-i} | \mathbf{x}_i, \boldsymbol{\theta}, \mathbf{y})$ is a Gaussian approximation to $\mathbf{x}_{-i} | \mathbf{x}_i, \boldsymbol{\theta}, \mathbf{y}$ around its mode $\mathbf{x}_{-i}^*(\mathbf{x}_i, \boldsymbol{\theta})$

\mathbf{x} is assumed to be a Gaussian Markov random field

$Q(\boldsymbol{\theta})$ is variance-covariance matrix
observations y_i are independent of each other

Discretization with triangulation

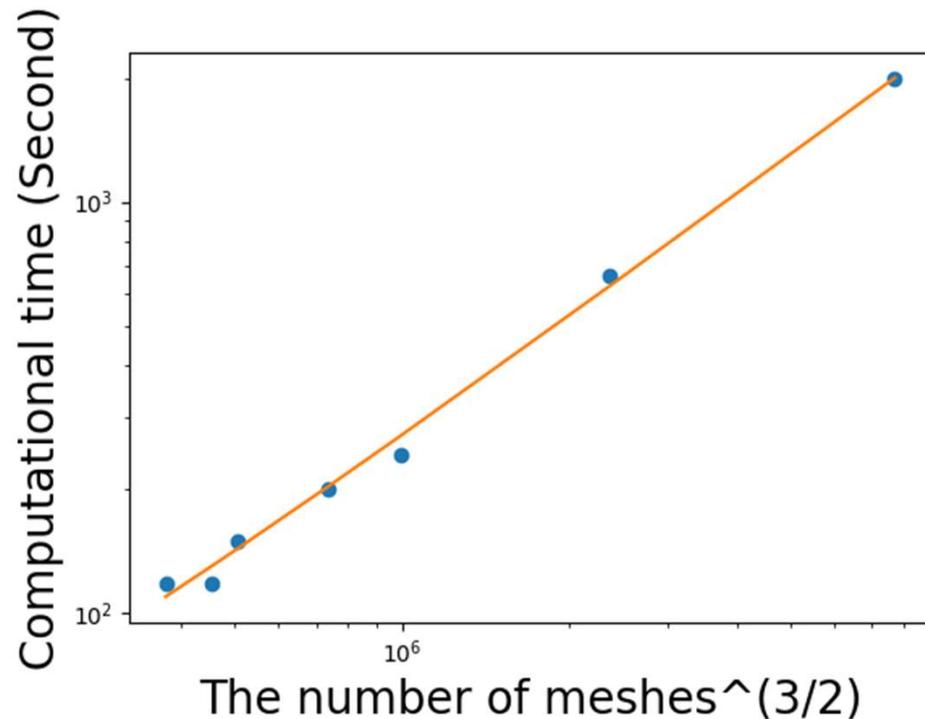
The number of mesh: 4525



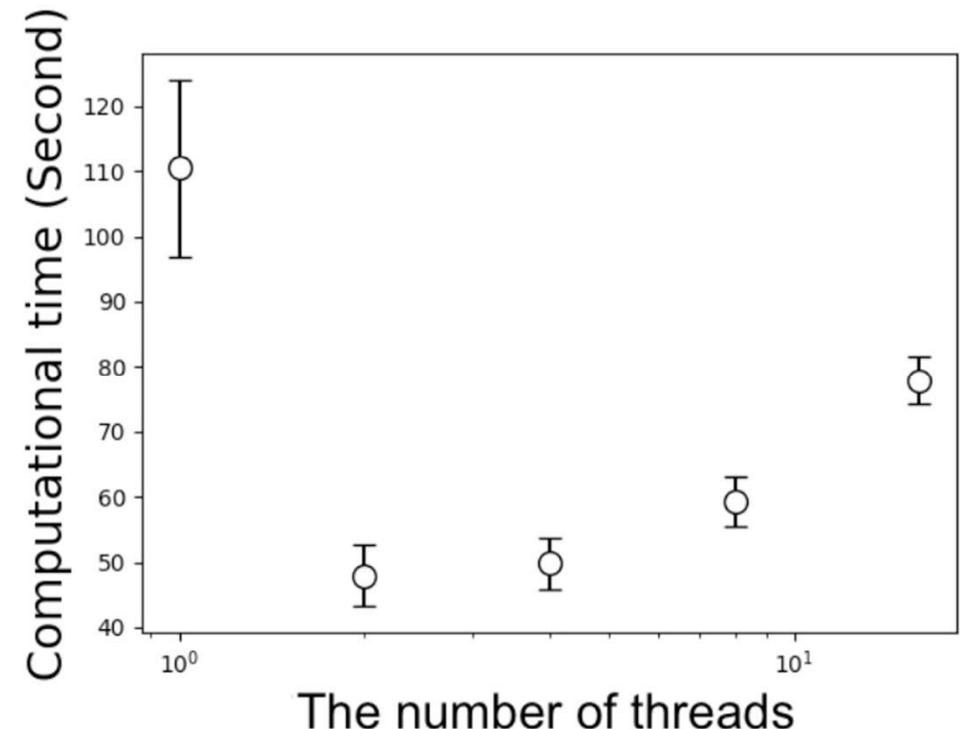
Need to find

1. the number mesh elements which allows faster computation and appropriate fitting.
2. the number threads to parallelise for faster computation.

The number of mesh elements



The number of threads



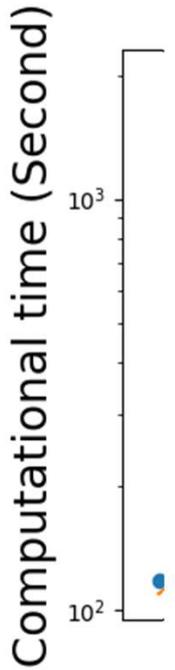
If the number of mesh elements is smaller than around 4500, the surface fitting fails. Also, the computation burden is $\mathcal{O}(n^{3/2})$.

Parallelization efficiency depends on the number of mesh elements. The most efficient number cannot be known until you do the actual calculations.

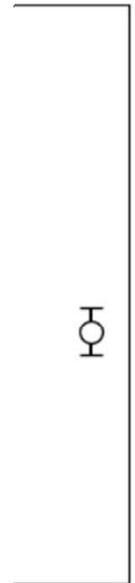
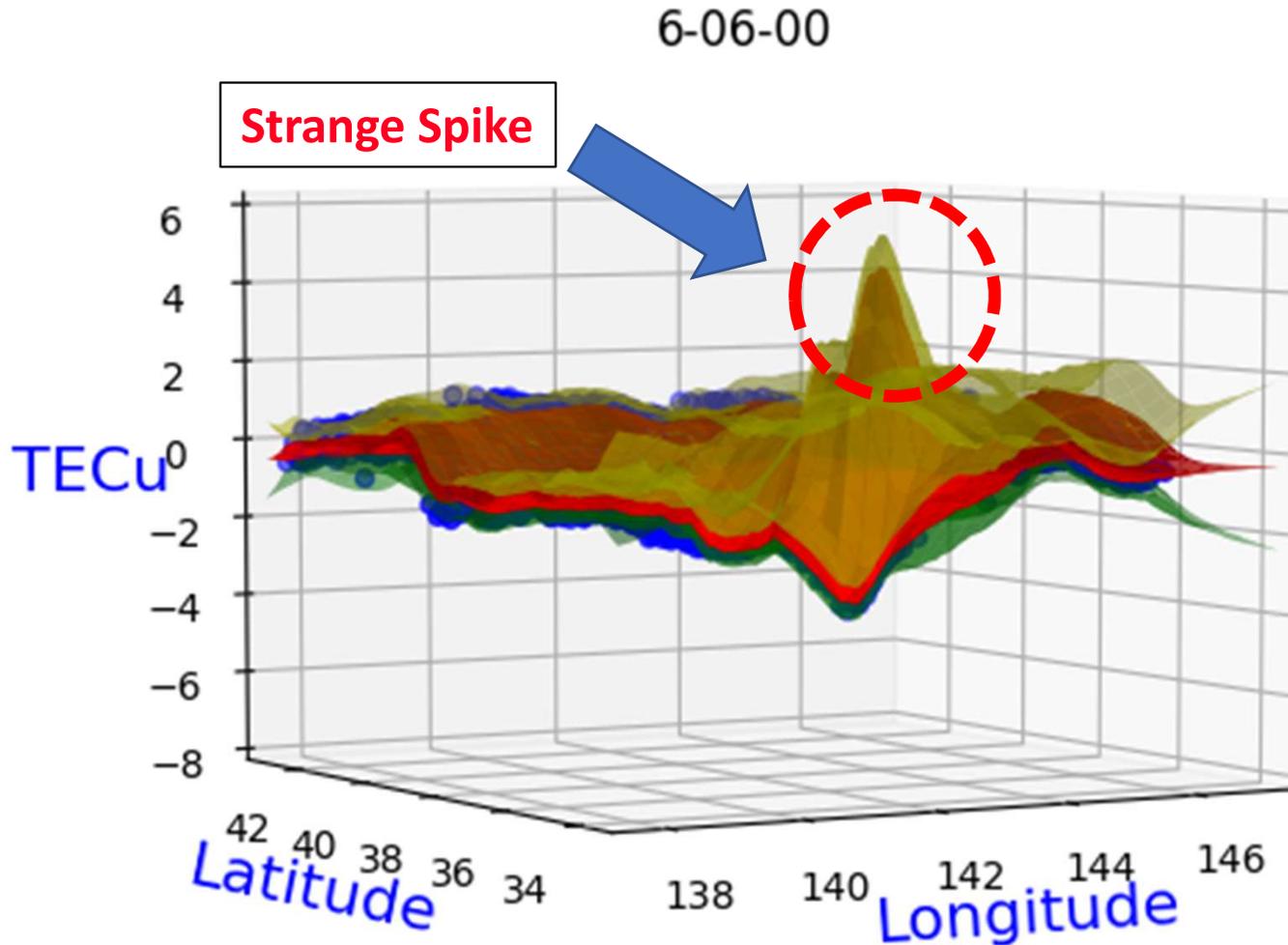
Parallelization



The



If the
smaller
surface
comp

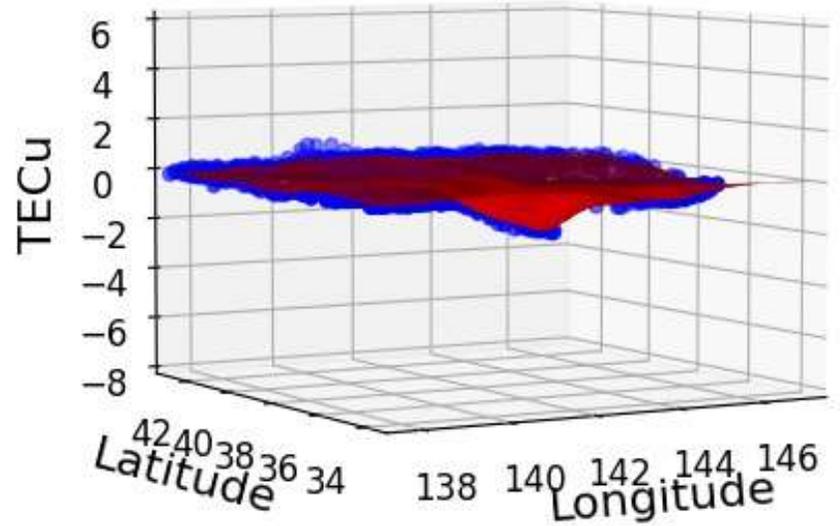


ds on
The
be

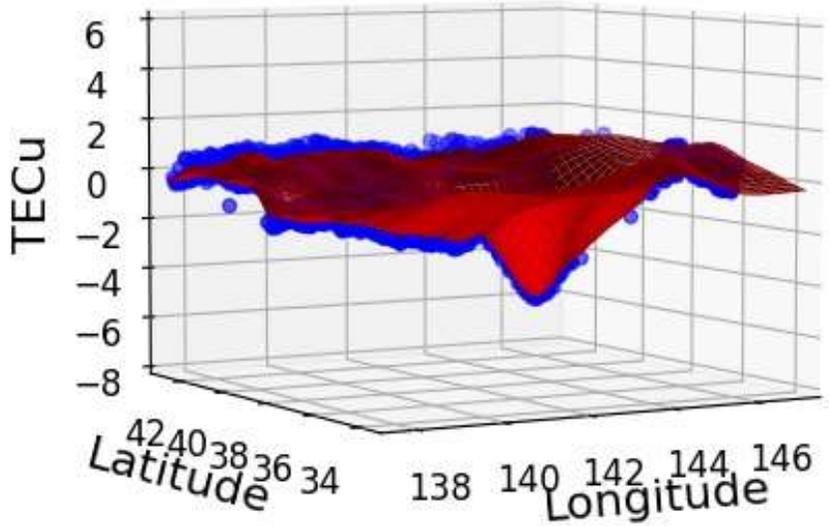
calculations.

Accelerated computation

6:00:00



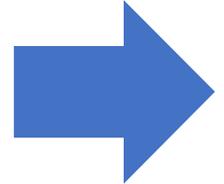
6:08:00



(sec)

Time	5:46:30	6:00:00	6:08:00	6:16:00
Ordinary Gaussian Process	411.5	627.6	653.6	736.7
INLA-SPDE	51.9	50.0	48.2	49.9

Computation time is less than 1 minute



Real-time analysis

1 Background

2 Application of statistics

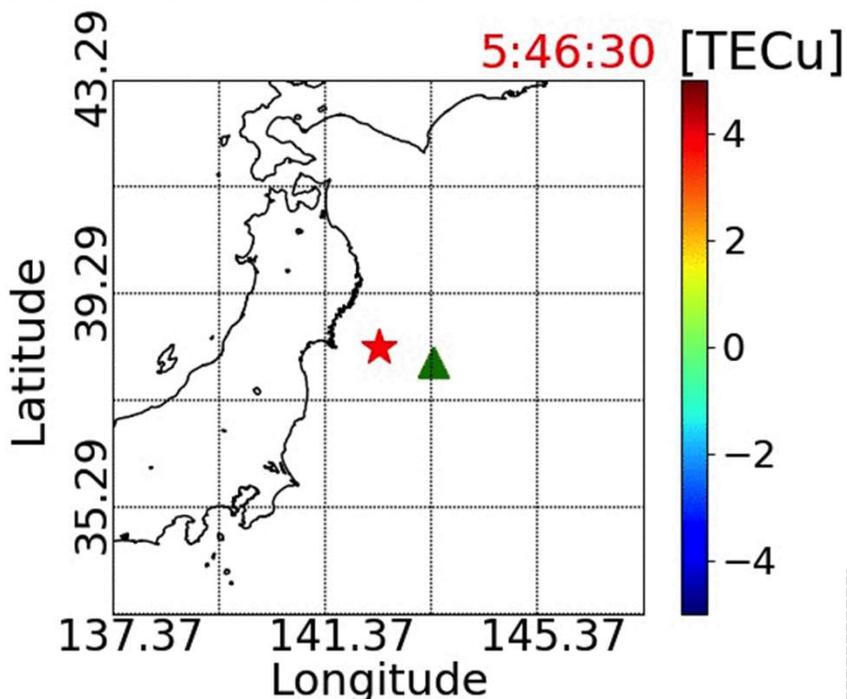
3 Results

4 Future step

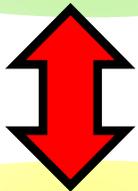


Anisotropic property

Our new method



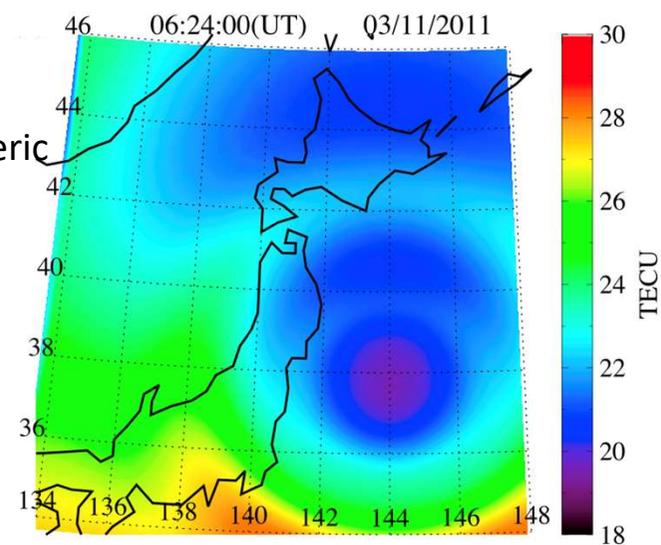
Anisotropy
Observation/our method



Symmetry
Simulation output

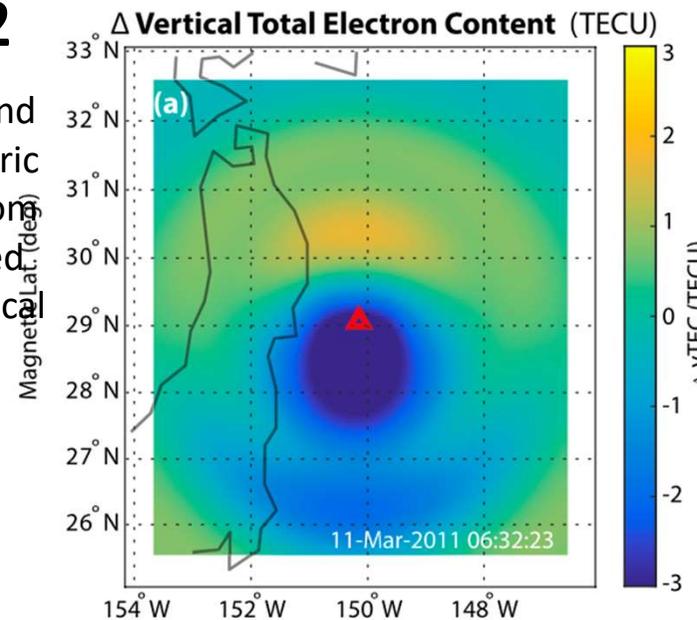
Previous study 1

H. Shinagawa, et. al., "Two-dimensional simulation of ionospheric variations in the vicinity of the epicenter of the Tohoku-oki earthquake on 11 March 2011", Geophysical Research Letters, 40, (2013)

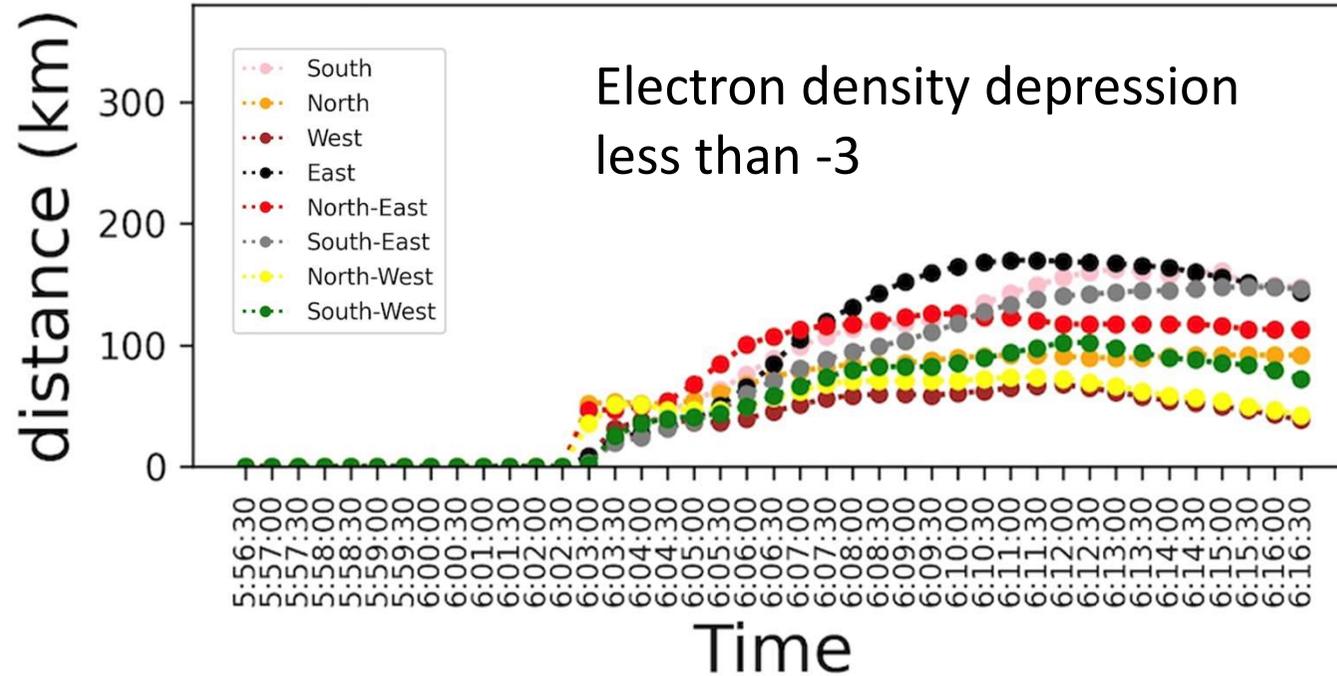


Previous study 2

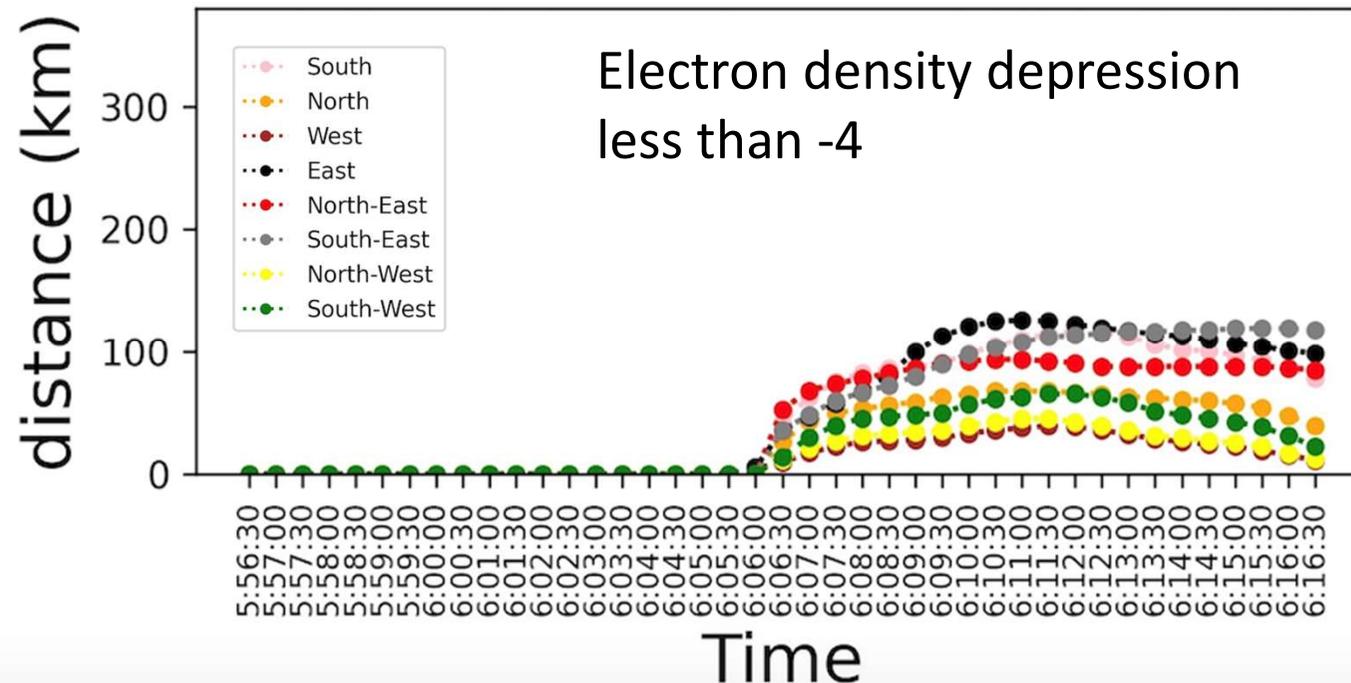
M. D. Zettergren, et. al., "Latitude and Longitude Dependence of Ionospheric TEC and Magnetic Perturbations From Infrasonic-Acoustic Waves Generated by Strong Seismic Events", Geophysical Research Letters, 46, (2019)



Expansion in each direction

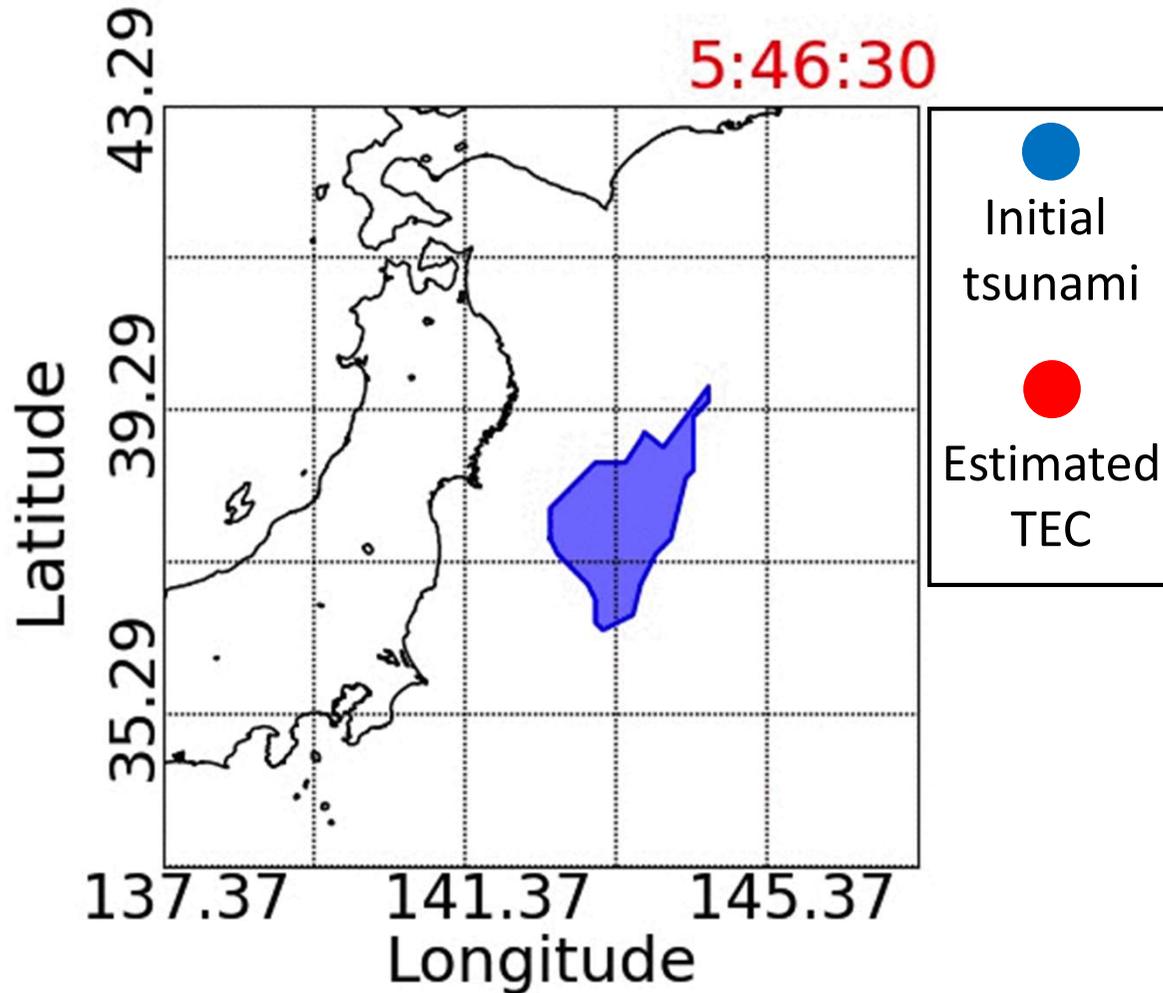


Original data **can't** show the expansion speed in each direction because of the data limitation.



Our method **succeeds** in estimating expansion speed in 8 direction for the first time.

Initial tsunami information



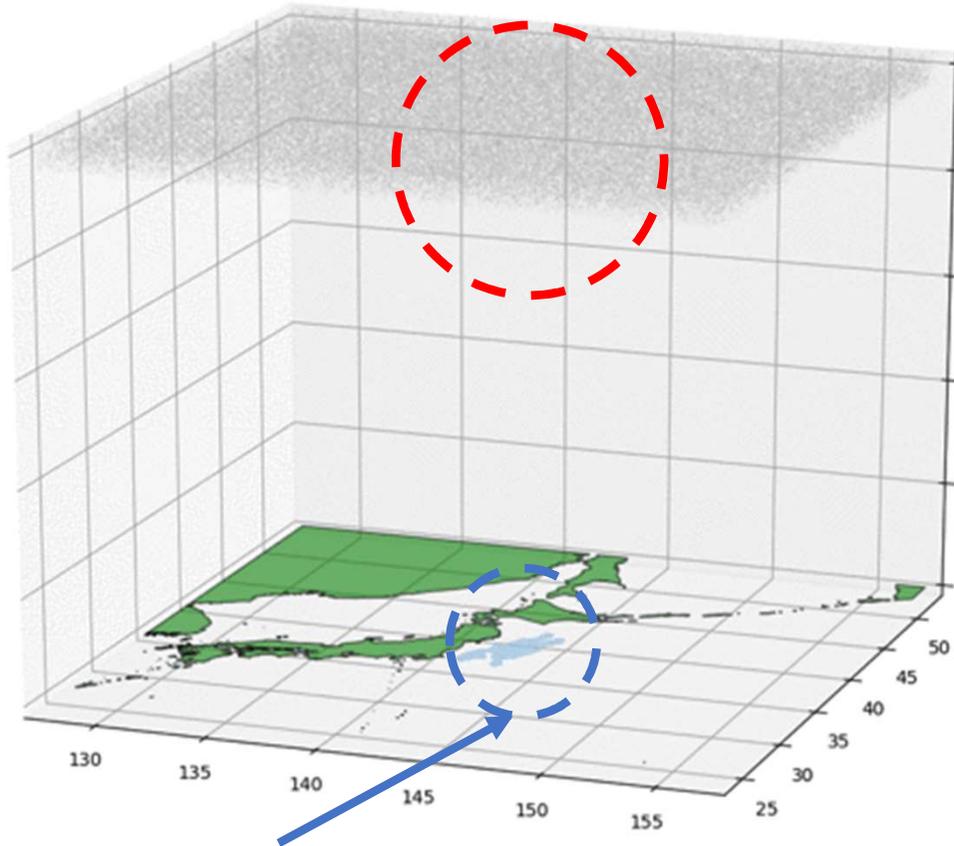
Initial tsunami area estimated by another research group **overlaps** the **electron density depression area**.



It indicates that the information obtained from the electron density depression contains the information about the initial tsunami.

Saito, T. et al., Tsunami source of the 2011 Tohoku-Oki earthquake, Japan: Inversion analysis based on dispersive tsunami simulations, Geophysical Research Letters, 38, (2011)

TEC depression



Initial tsunami

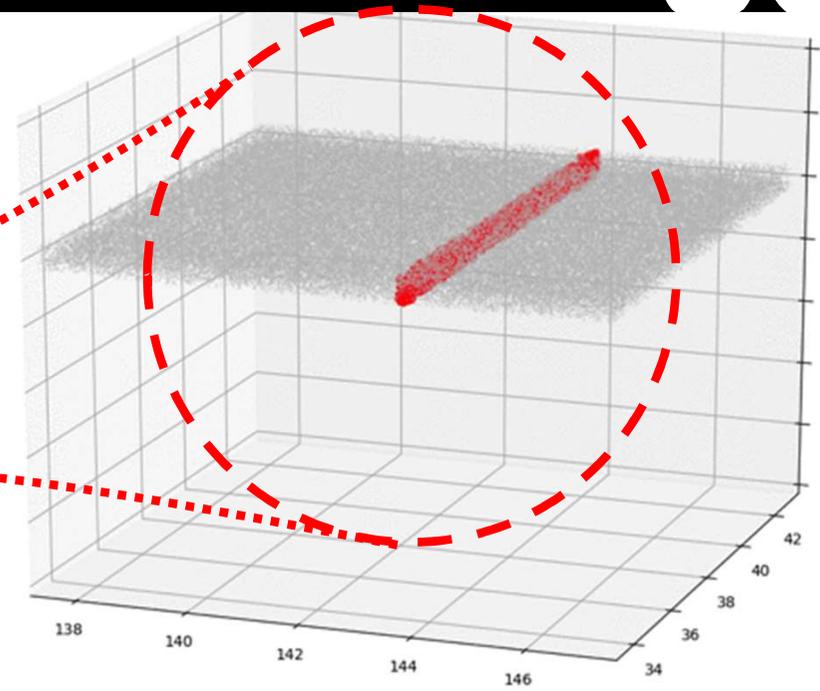
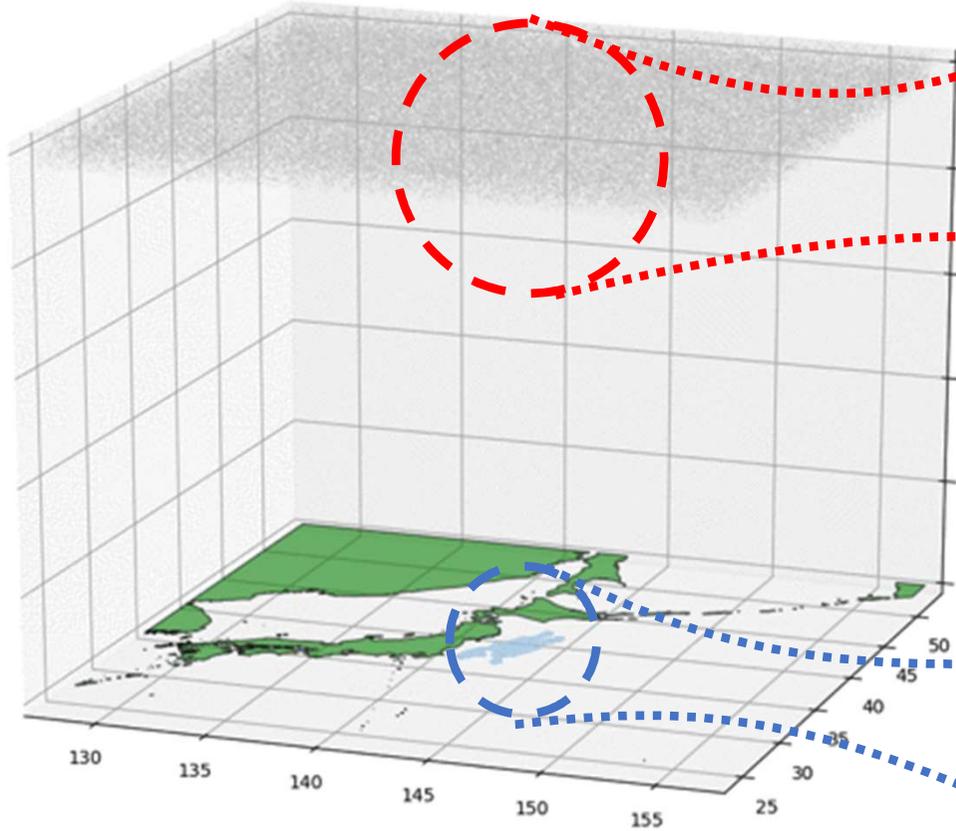
- ✓ Initial tsunami region
- Initial tsunami height?



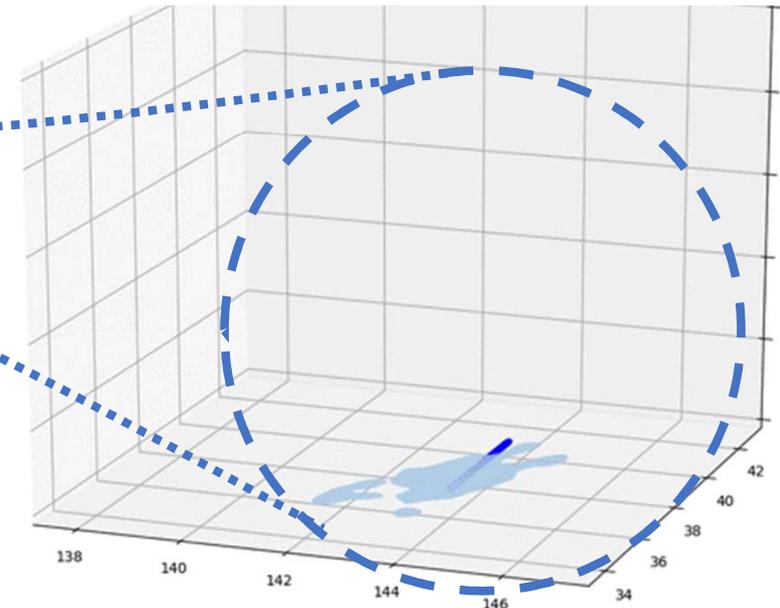
Input data for tsunami propagation model.
E.g. Jugrus, VOLNA....

Initial tsunami information

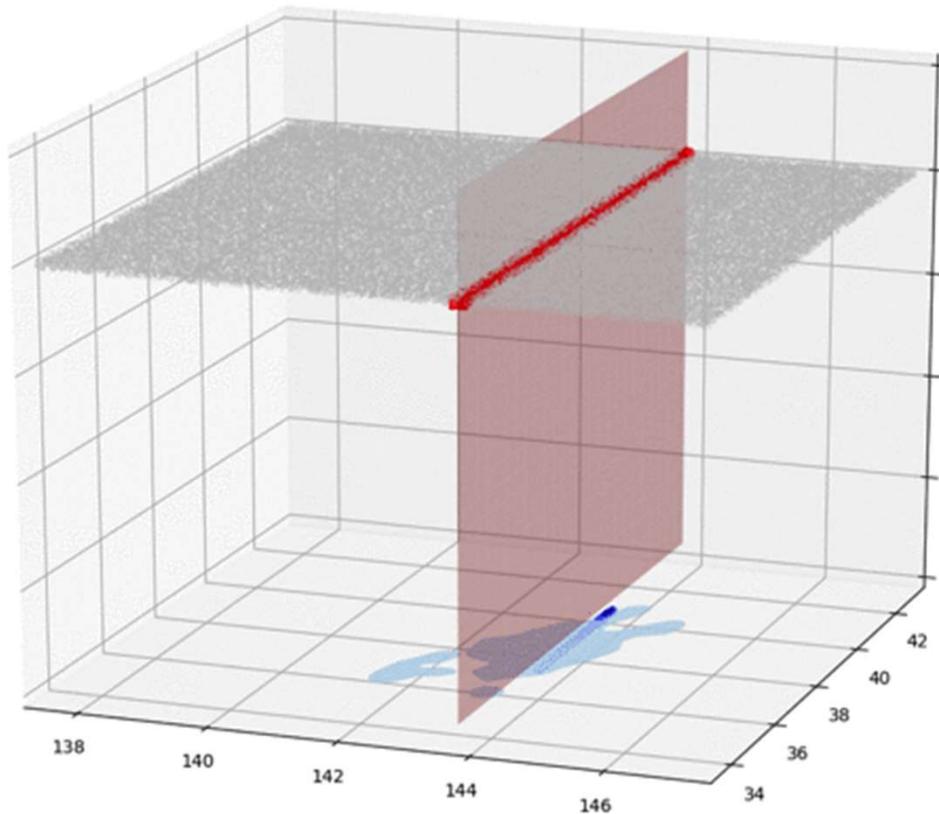
TEC depression



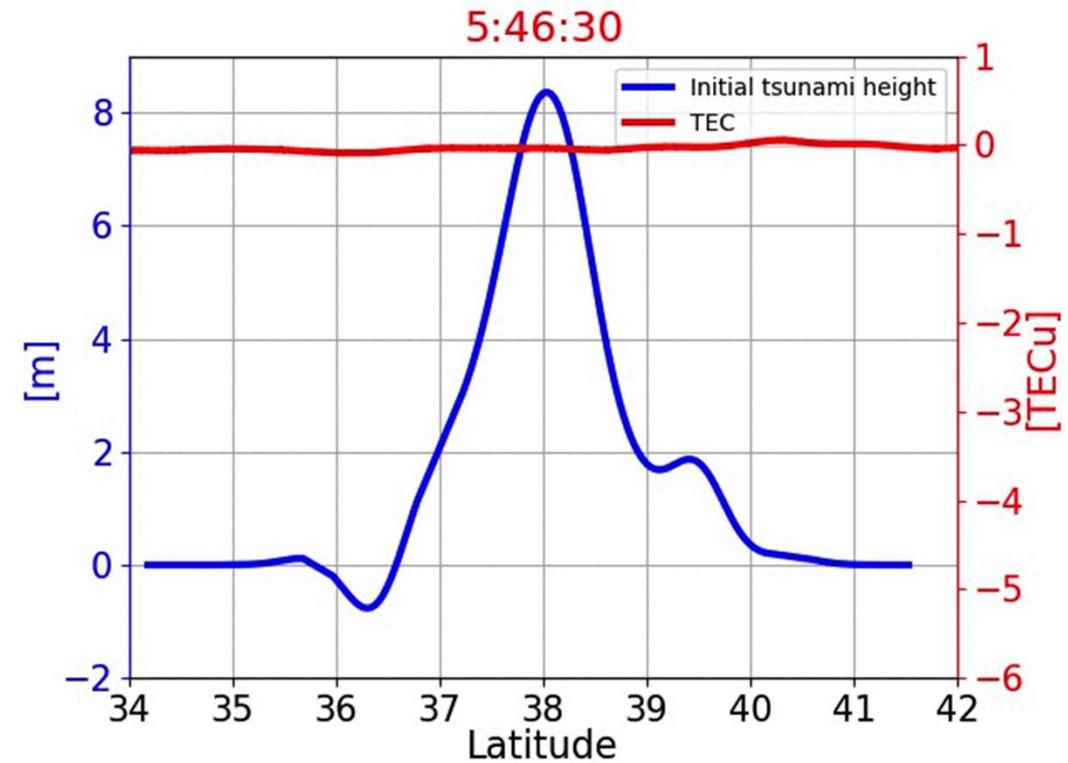
Initial tsunami



TEC depression

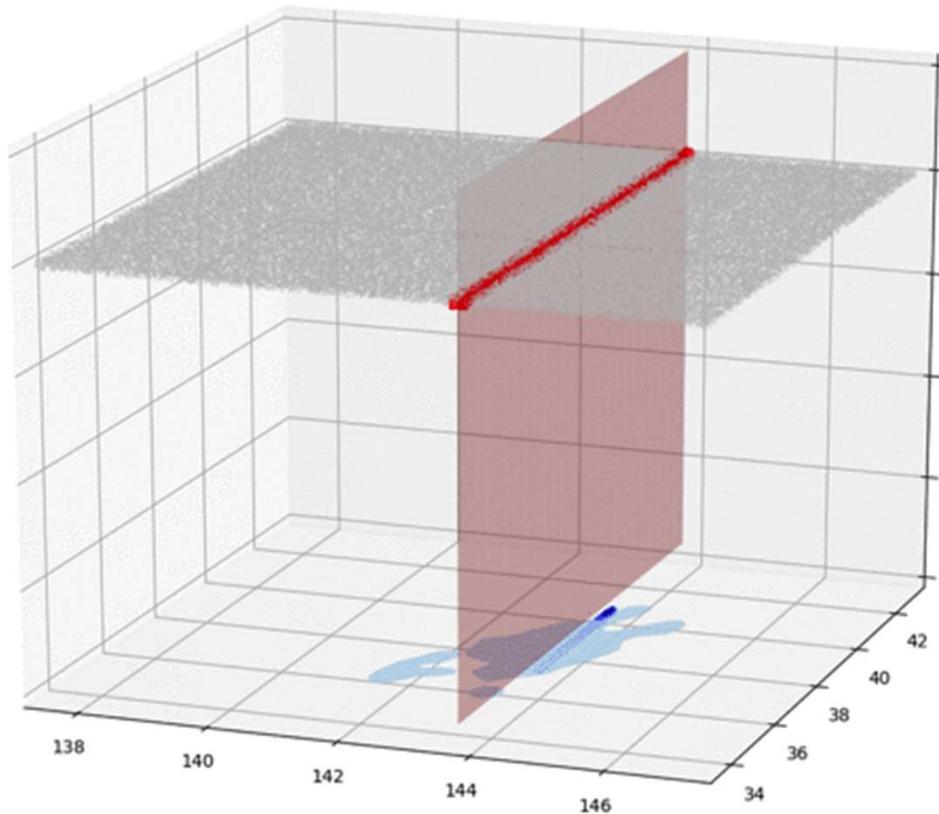


Initial tsunami



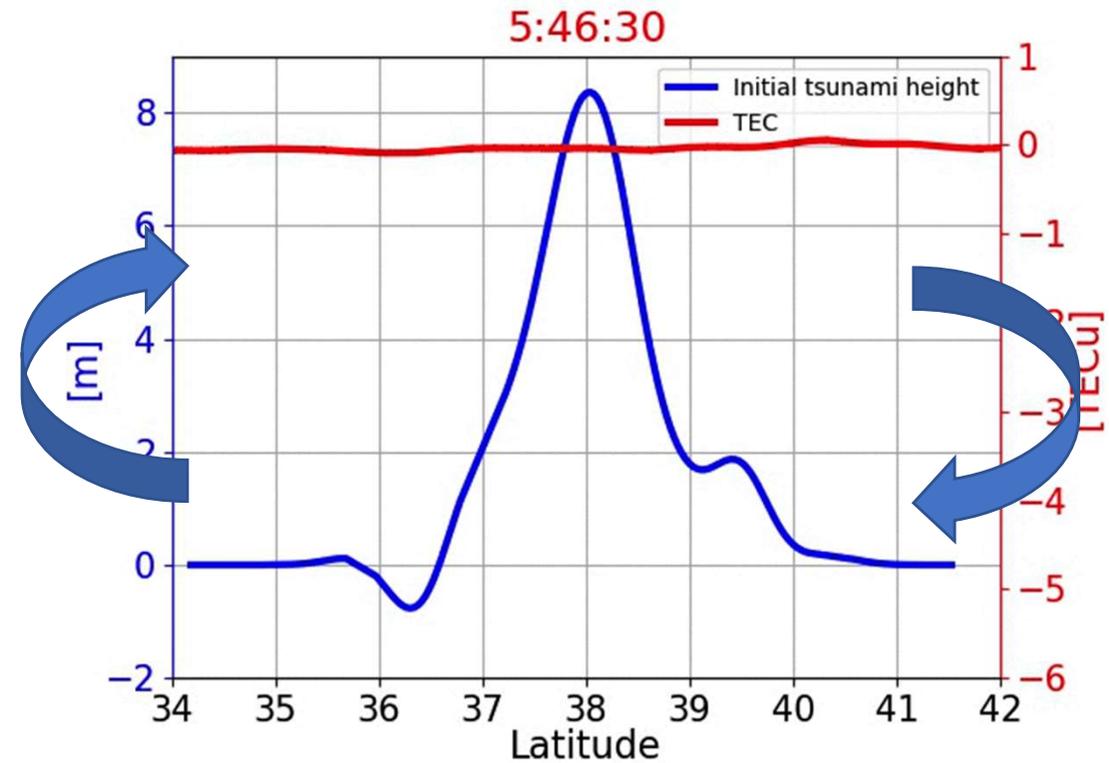
Initial tsunami shape and TEC depression shape are correlated.

TEC depression



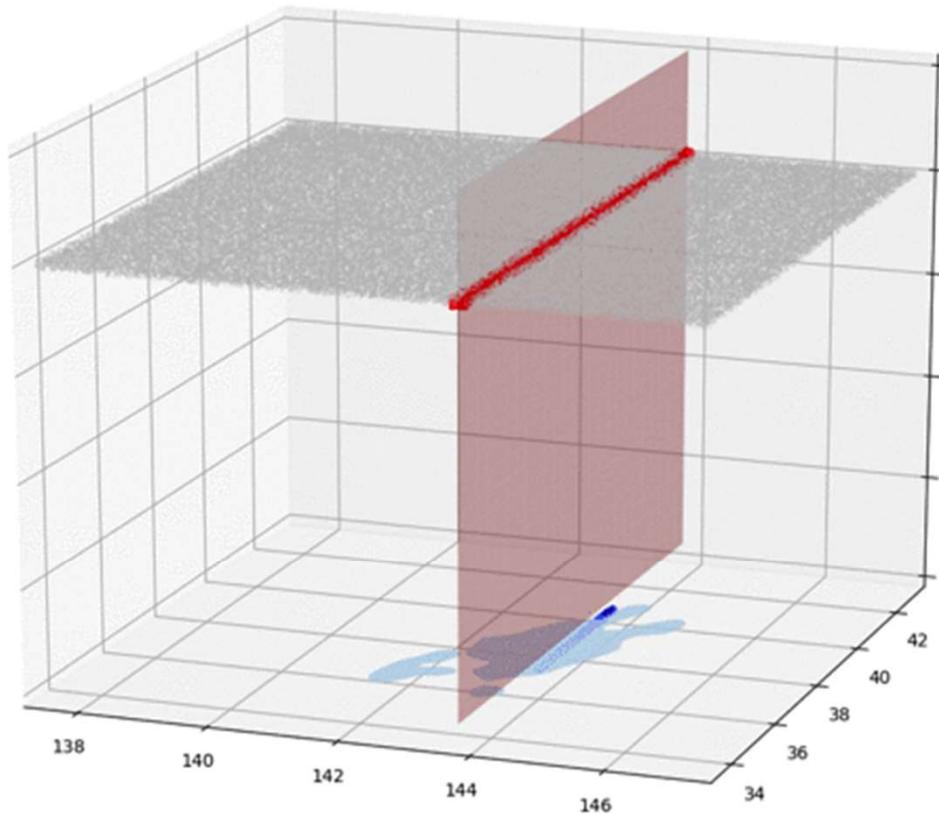
Initial tsunami

Turn it upside down

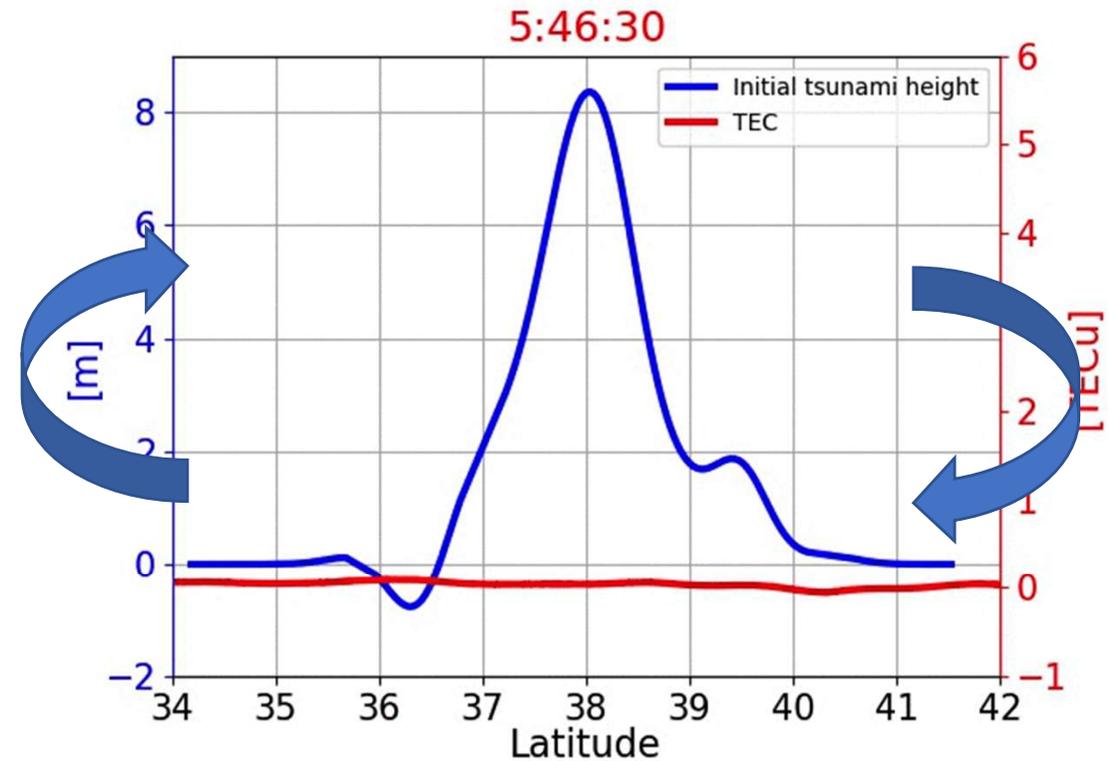


Initial tsunami shape and TEC depression shape are correlated.

TEC depression

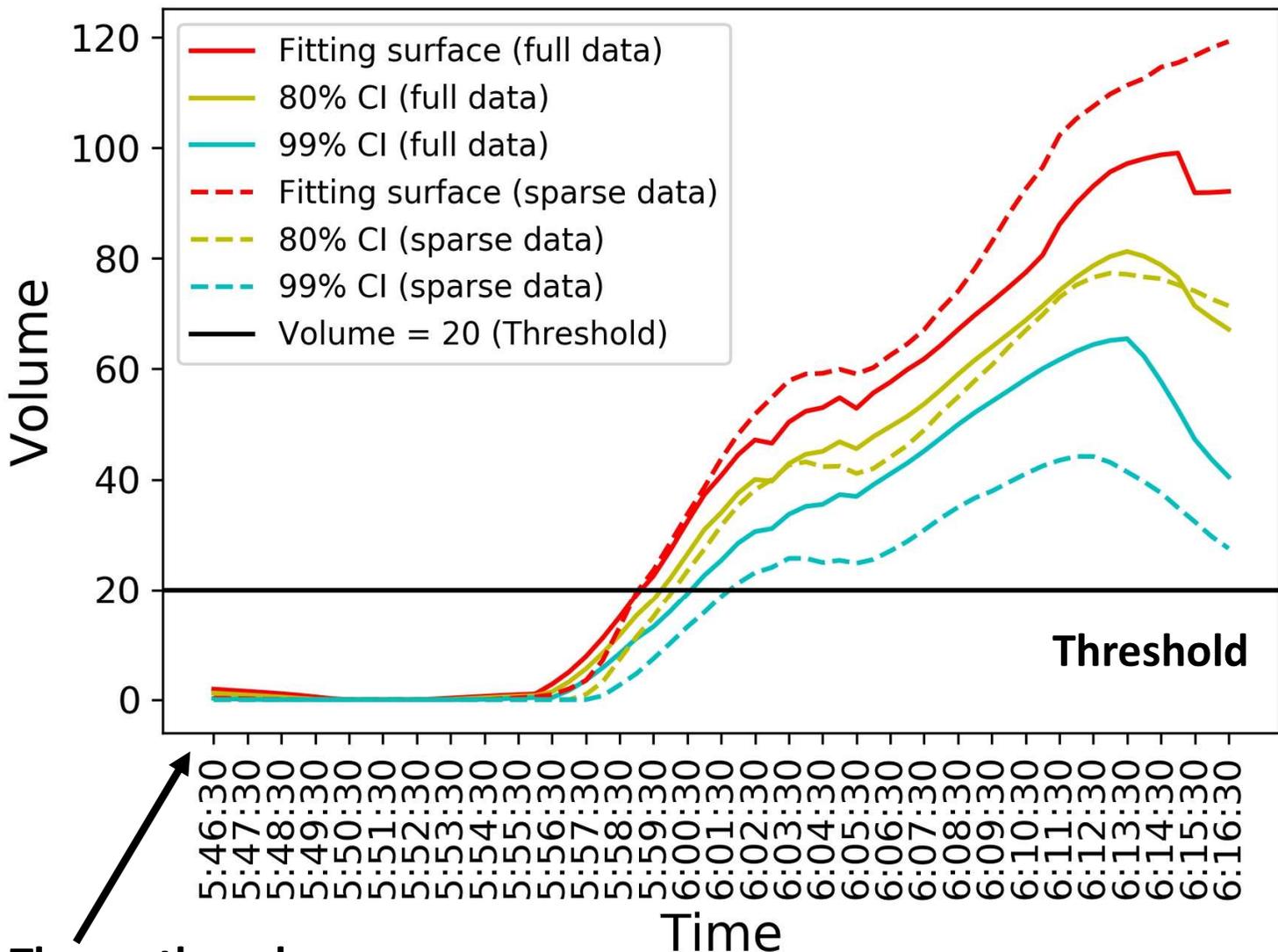


Initial tsunami

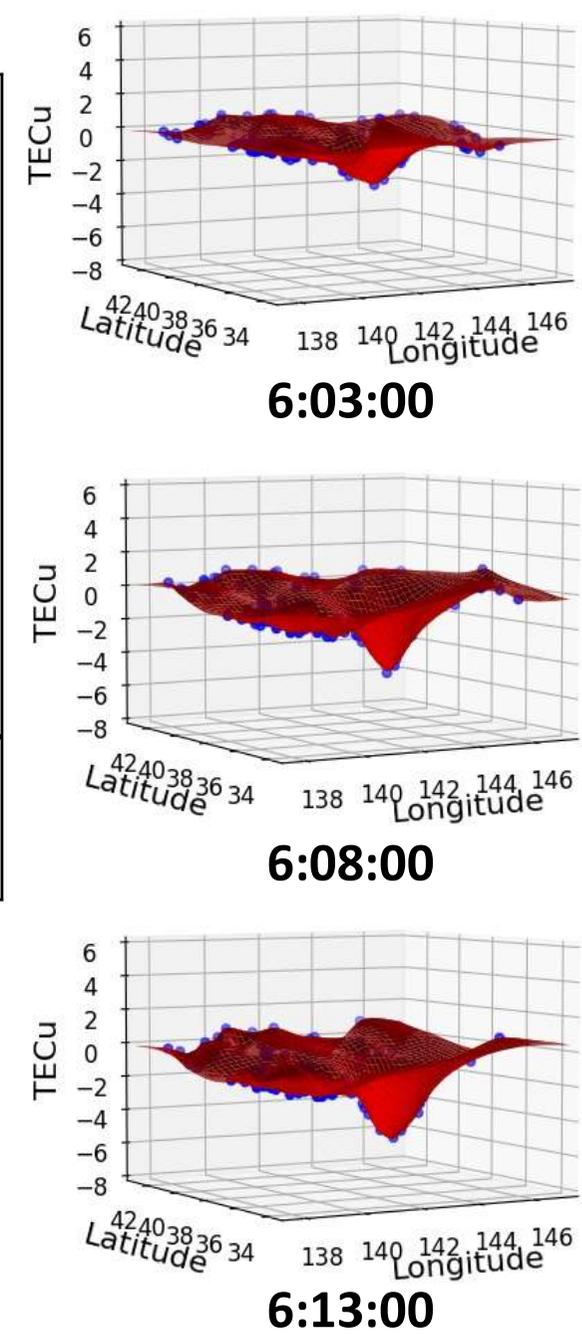


Explicitly TEC depression data has information about initial tsunami shape

Volume computation and warning system



The earthquake occurrence



The depression time series

**Frequency
Filter**

- 1 -

**Outlier
detection**

- 2 -

**Surface
fitting**

- 3 -

**Computation
Acceleration**

- 4 -

- ✓ **Dense estimation with uncertainty**
- ✓ **Propagation speed estimation**
- ✓ **Tsunami early warning system**
- ✓ **Initial tsunami information possibility**

1 Background

2 Application of statistics

3 Results

4 Future step



Ultimate goal

Ionosphere

TEC depression

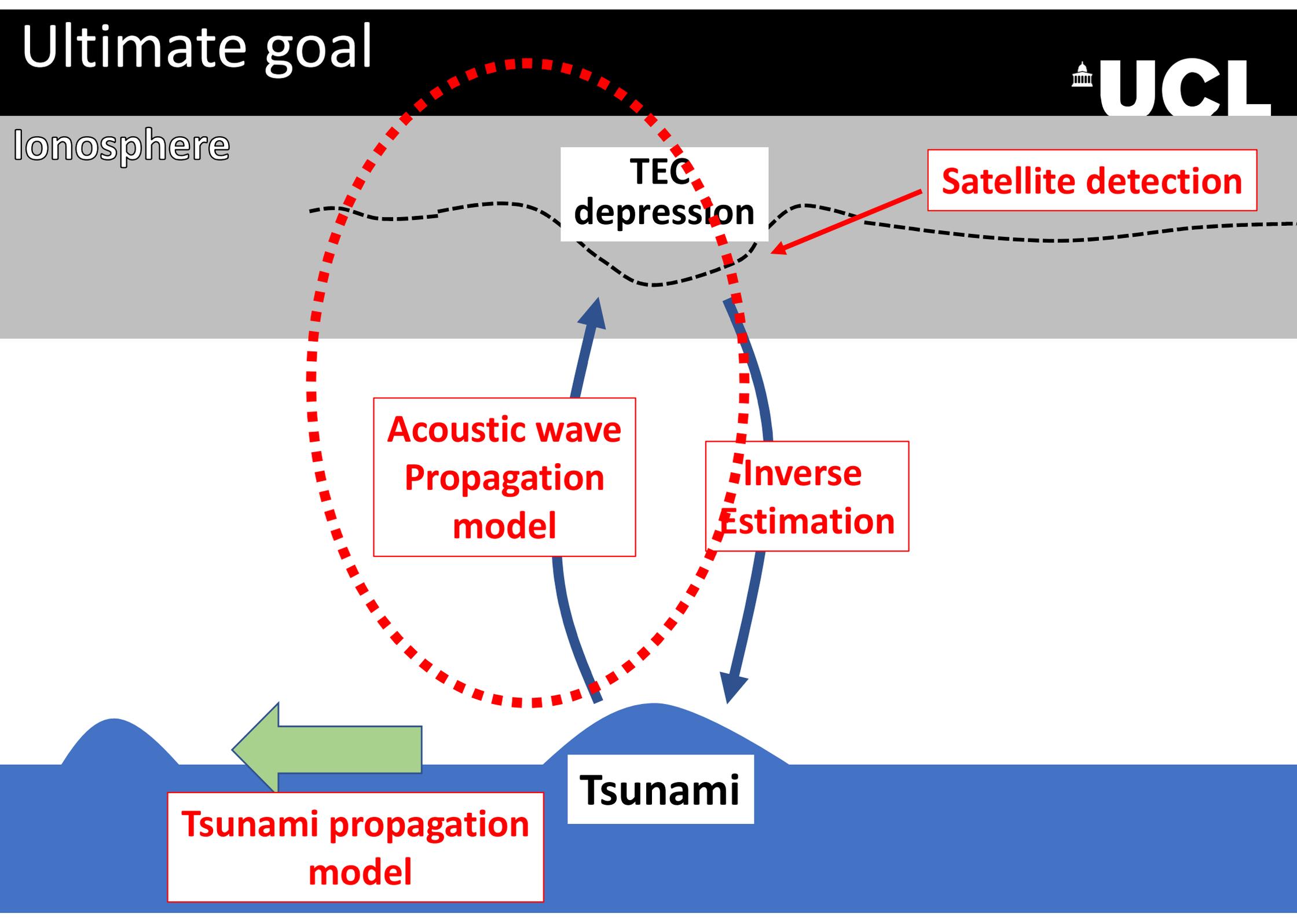
Satellite detection

Acoustic wave Propagation model

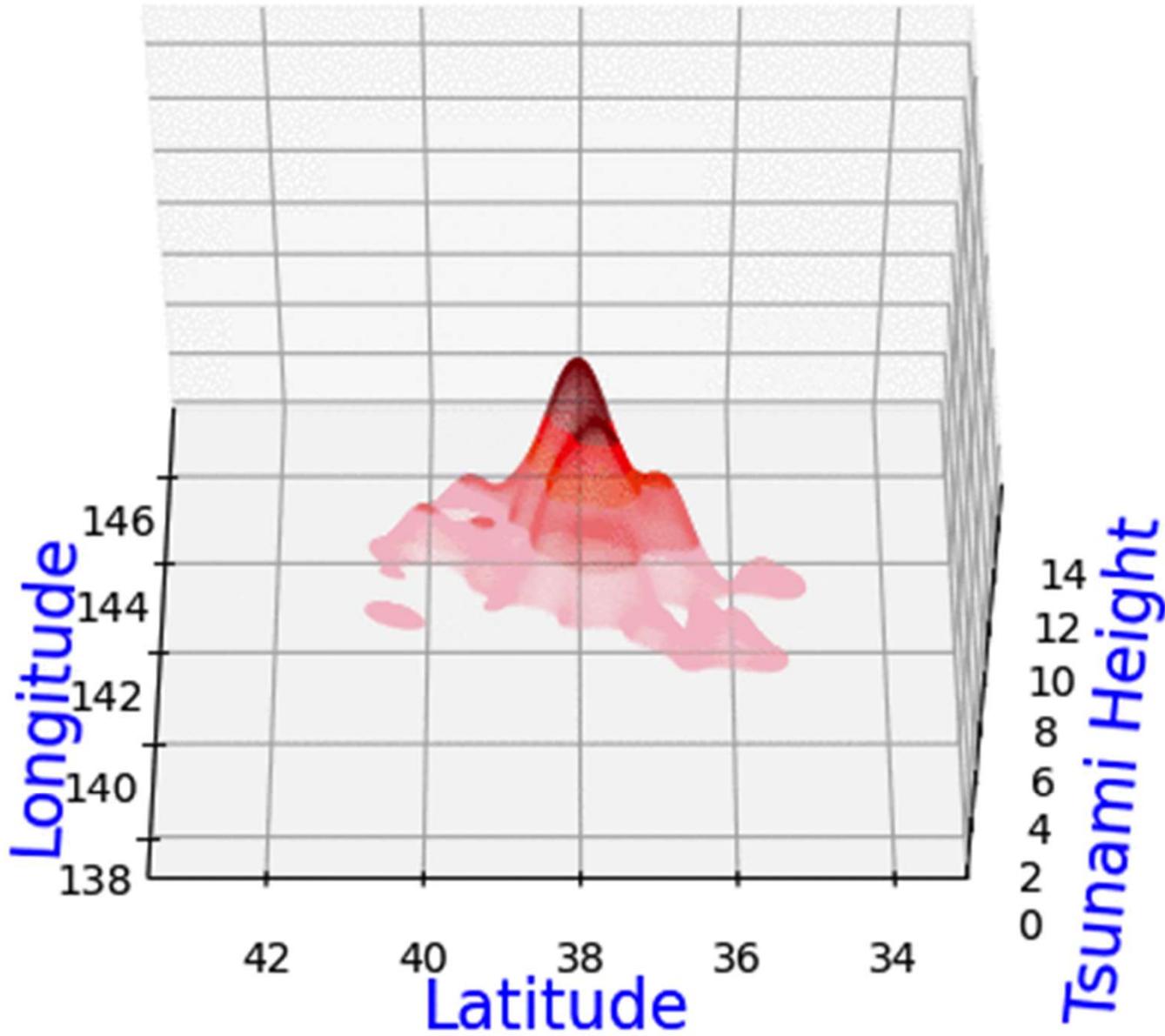
Inverse Estimation

Tsunami propagation model

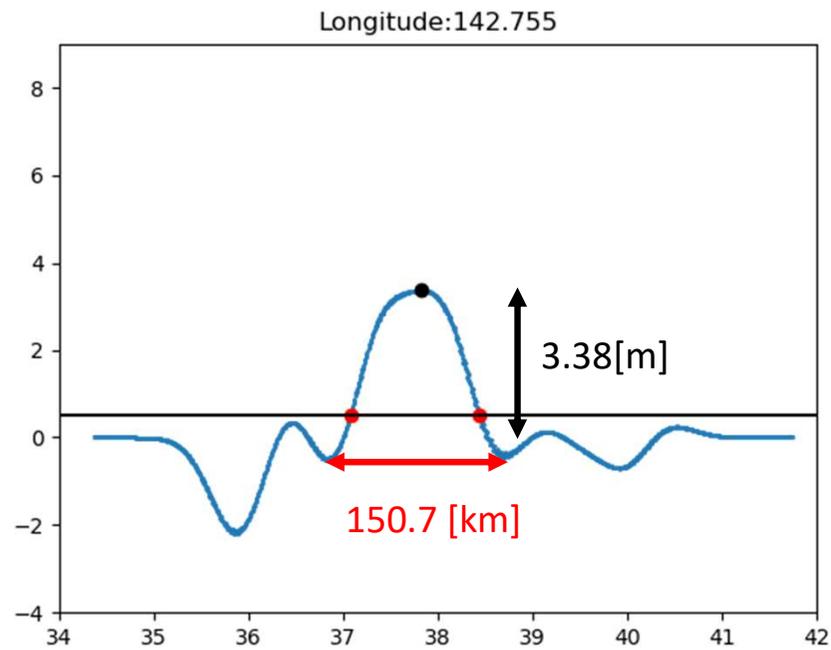
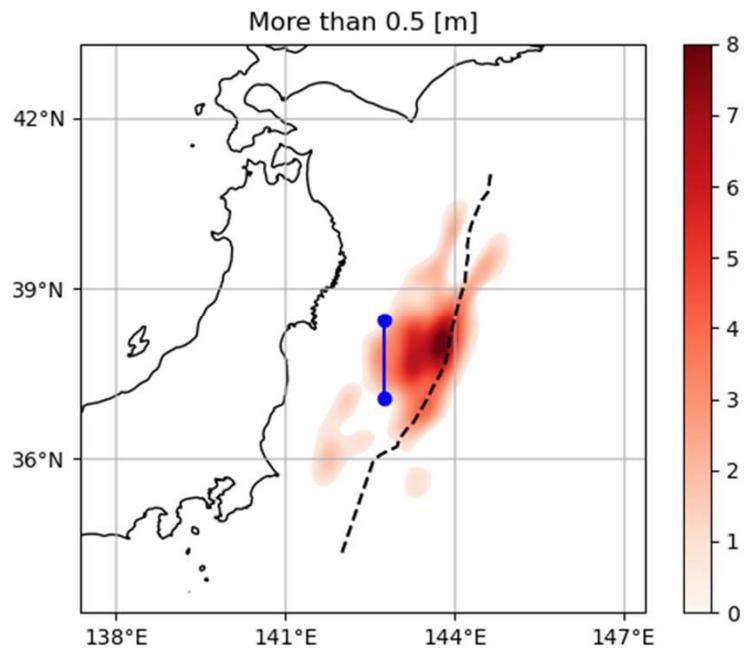
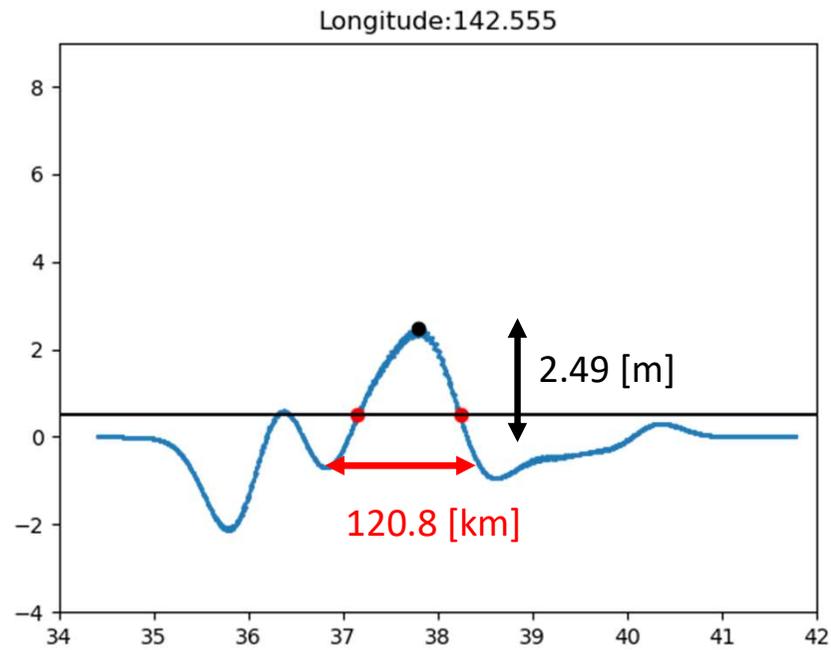
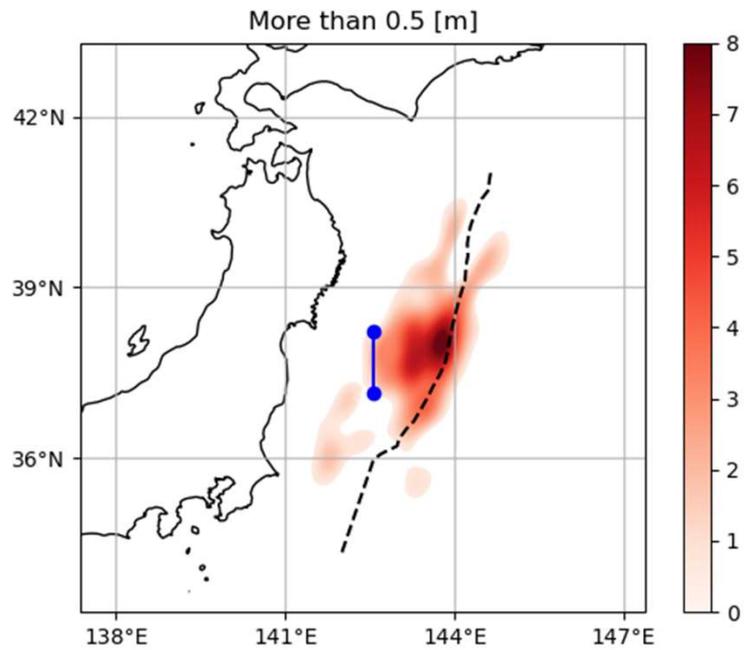
Tsunami



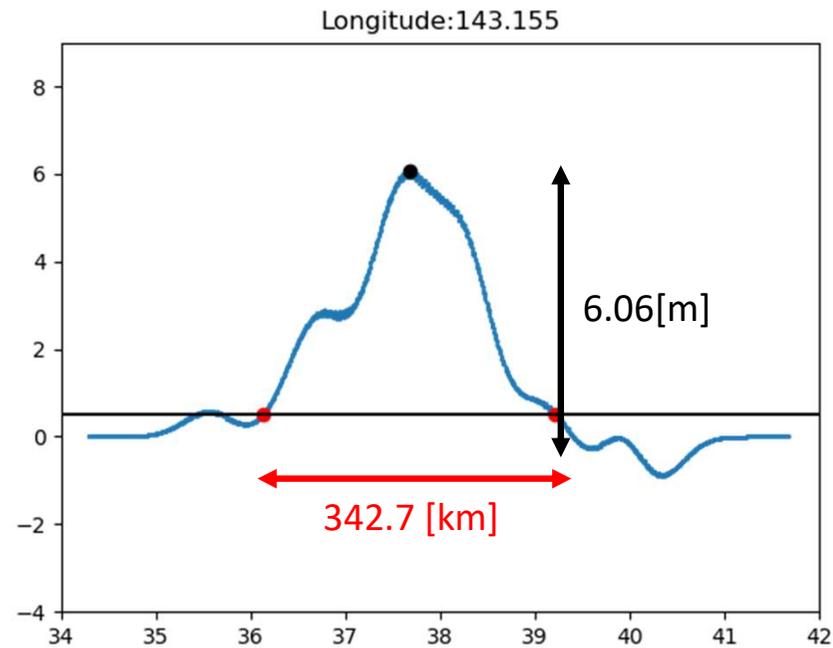
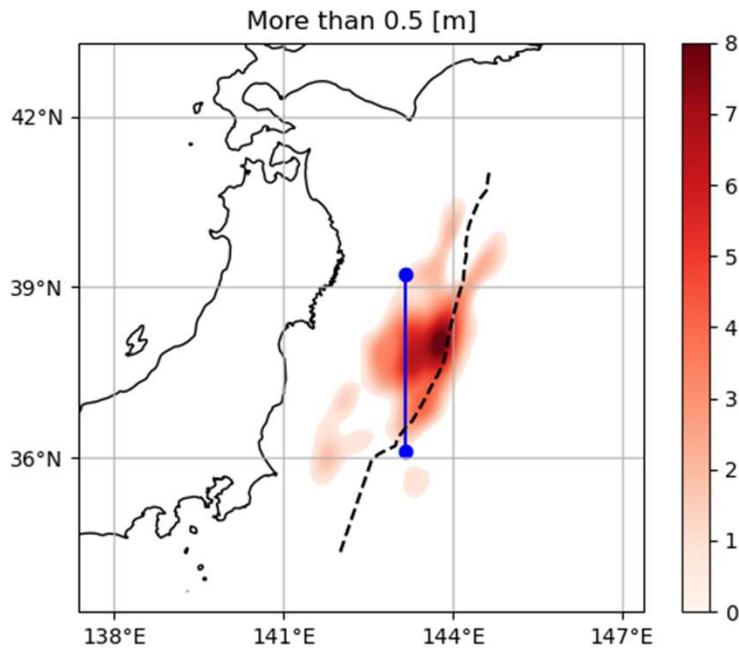
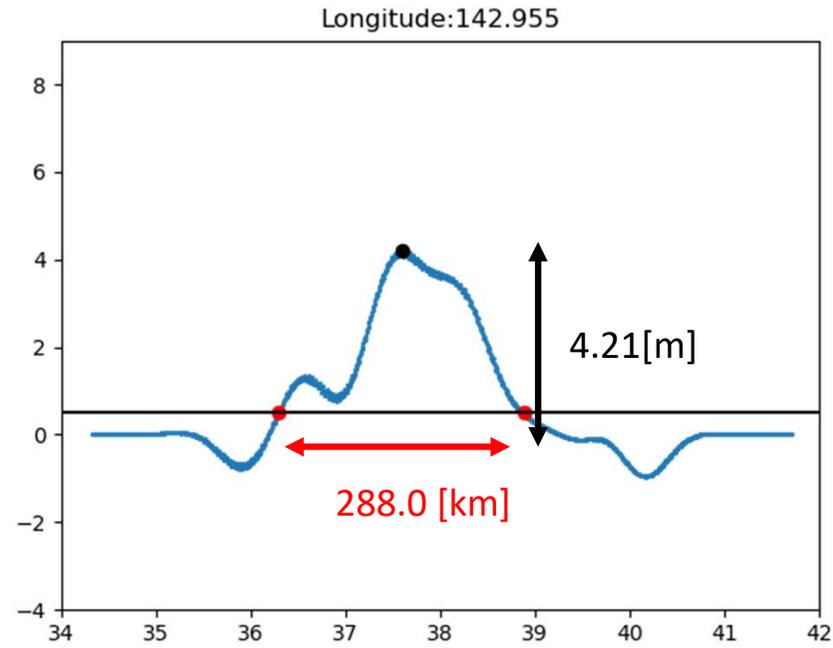
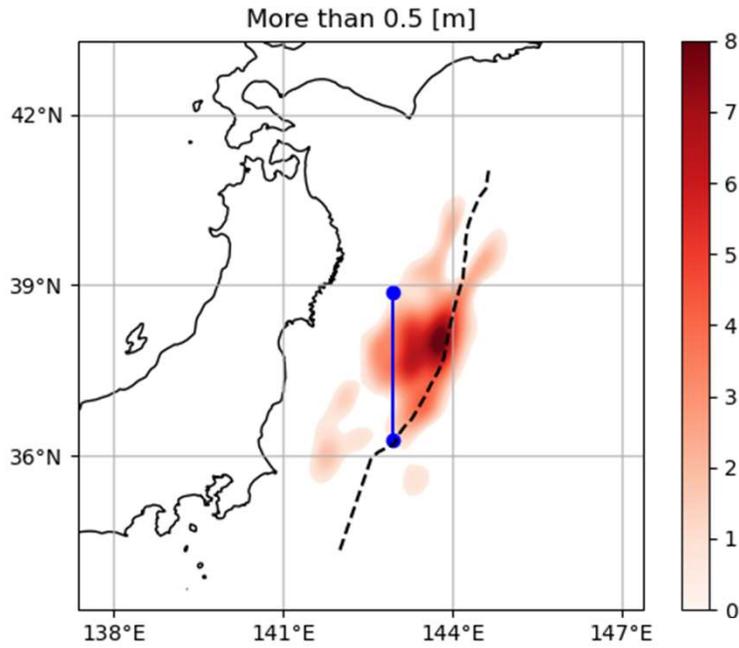
Initial tsunami



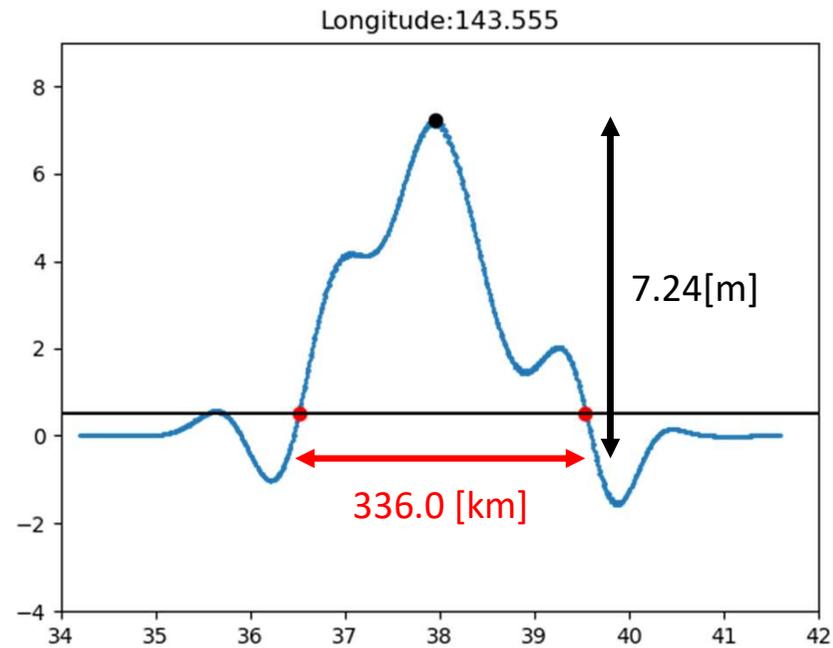
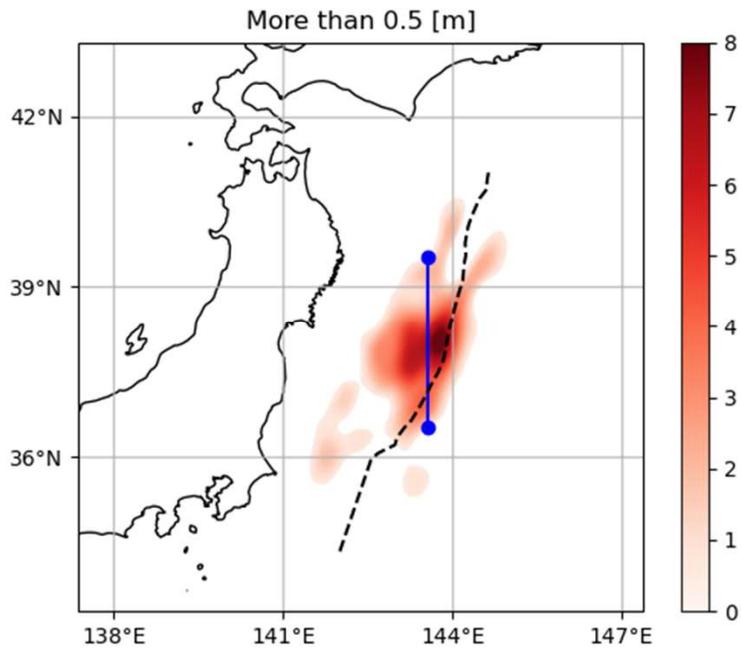
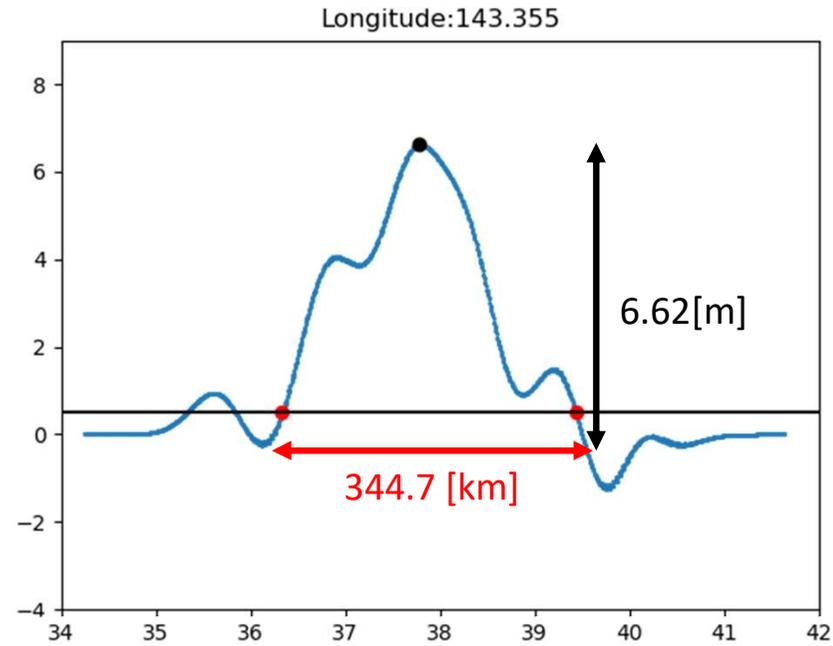
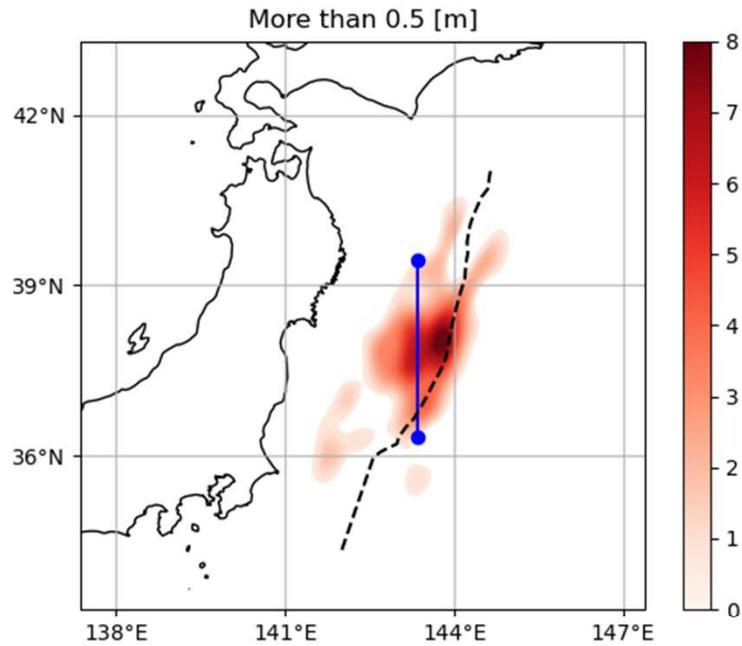
Initial Tsunami



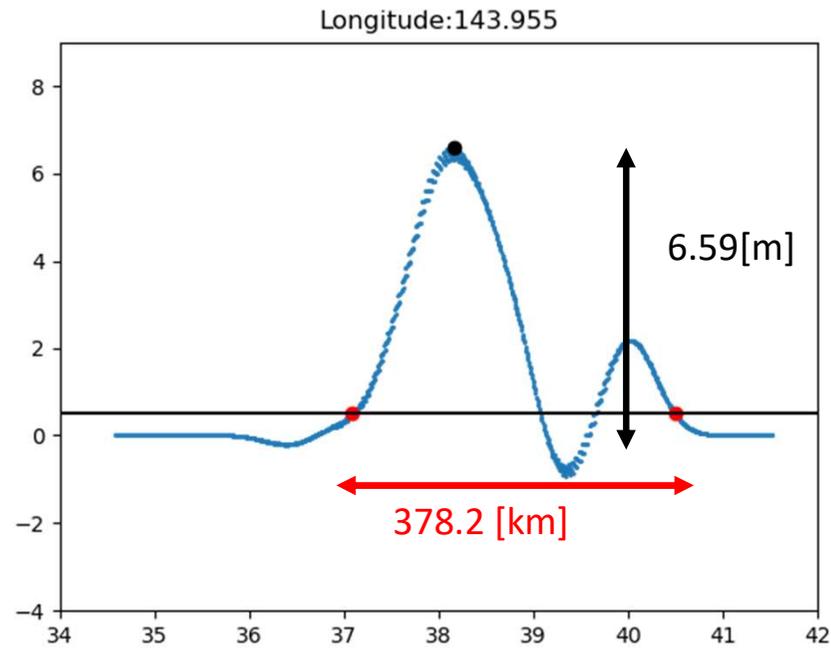
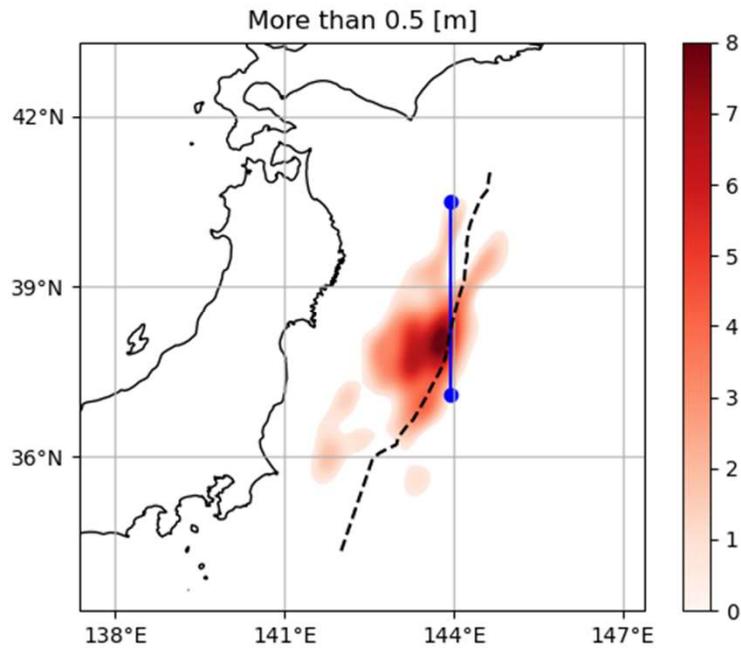
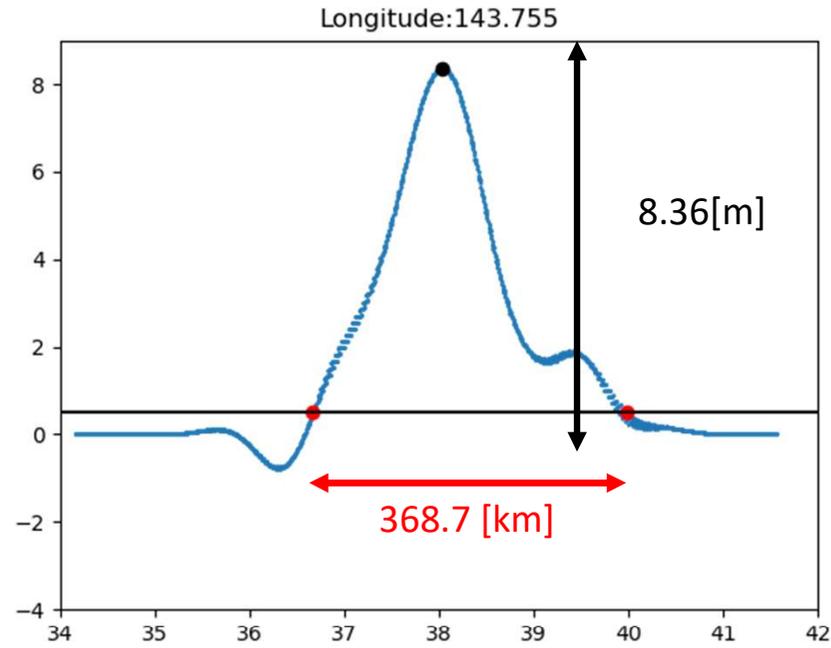
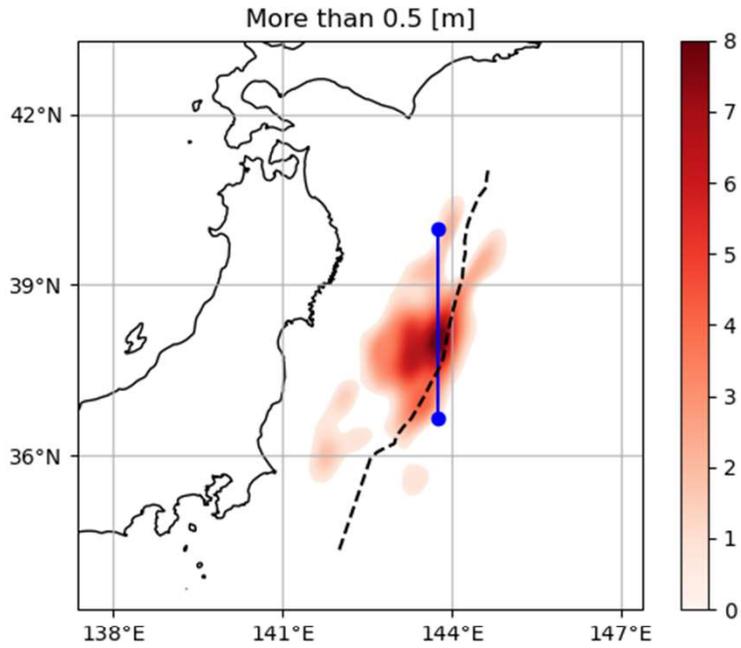
Initial Tsunami



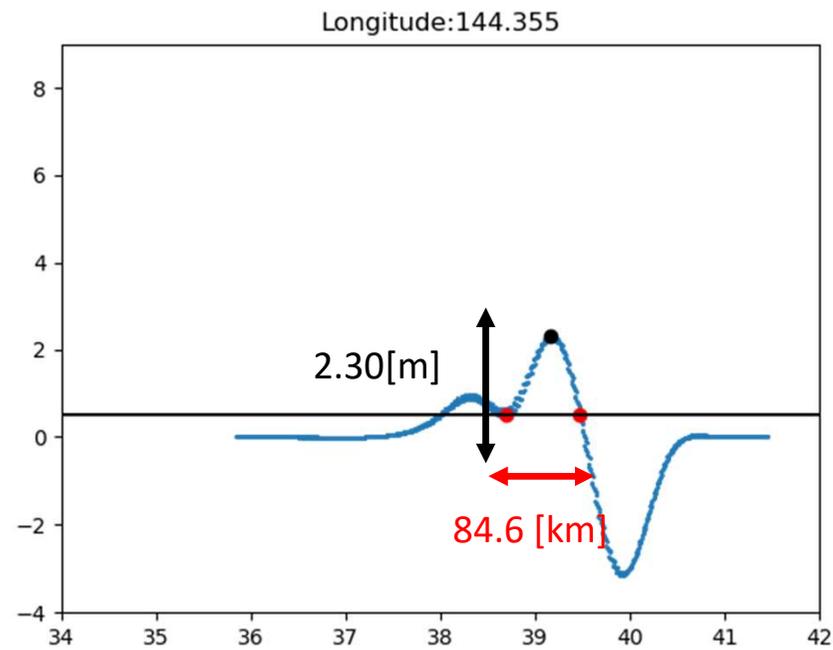
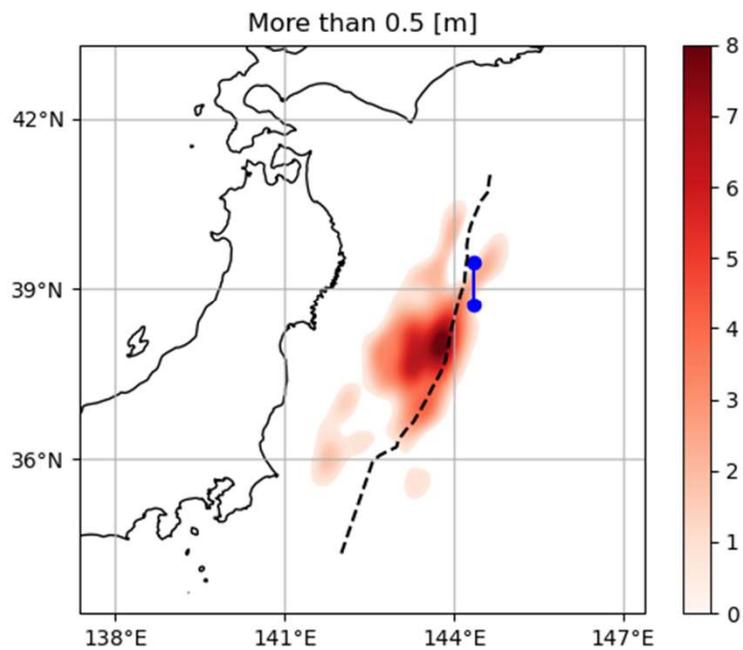
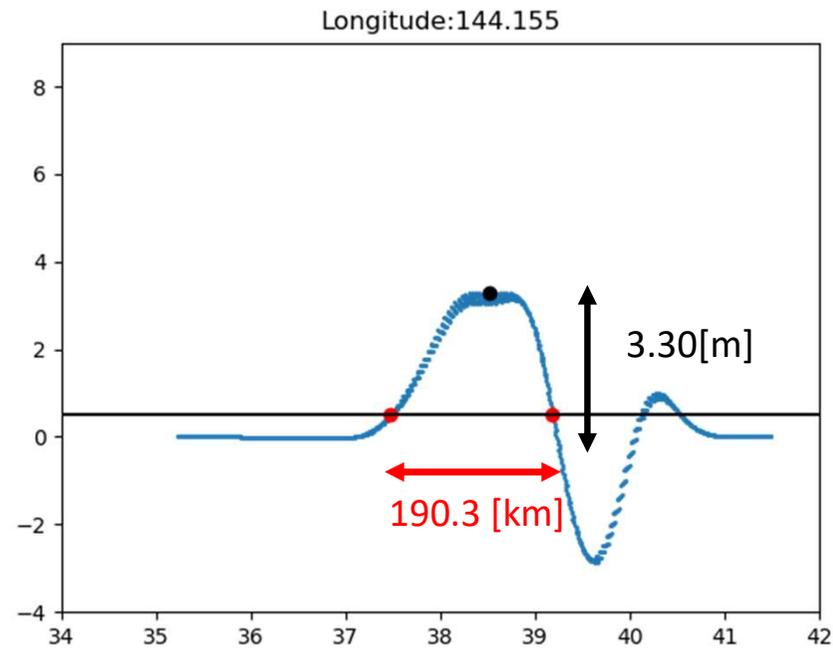
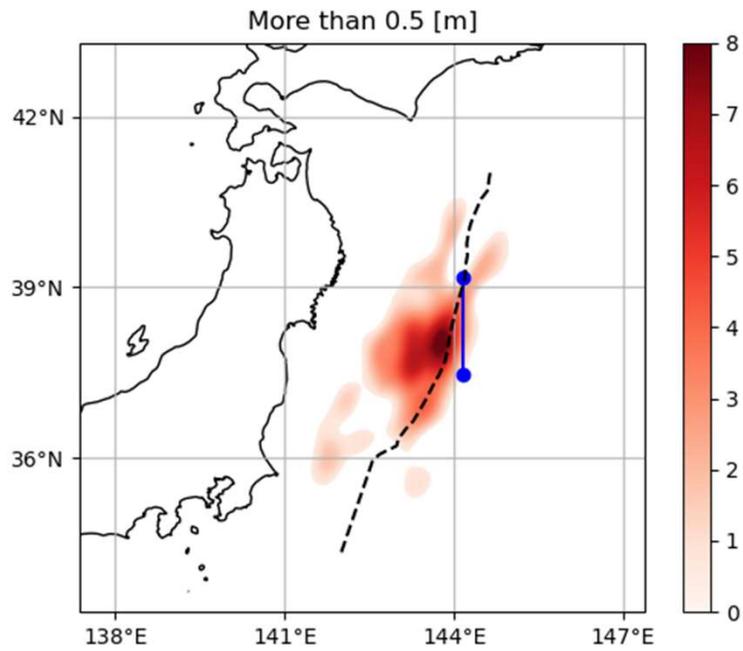
Initial Tsunami



Initial Tsunami

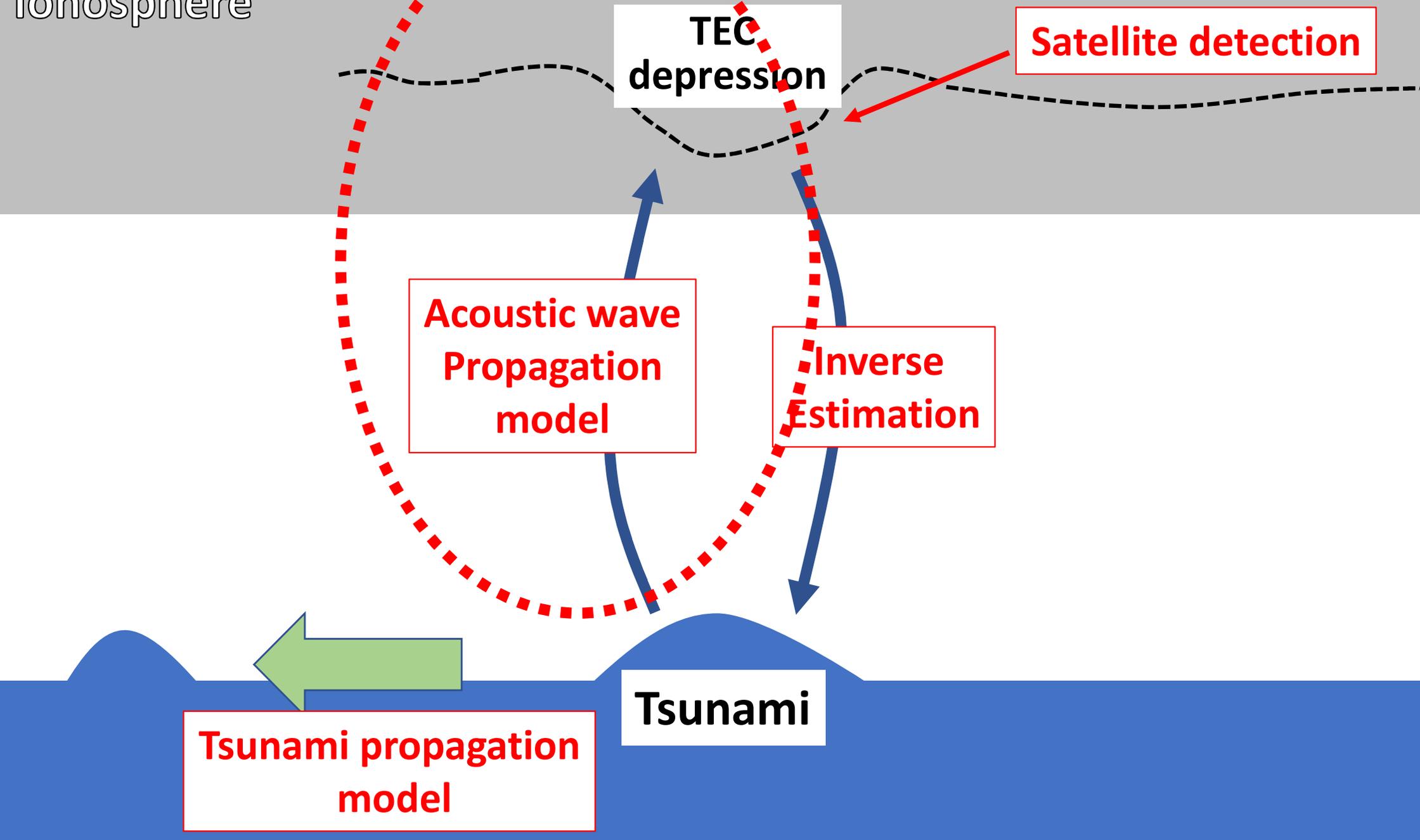


Initial Tsunami



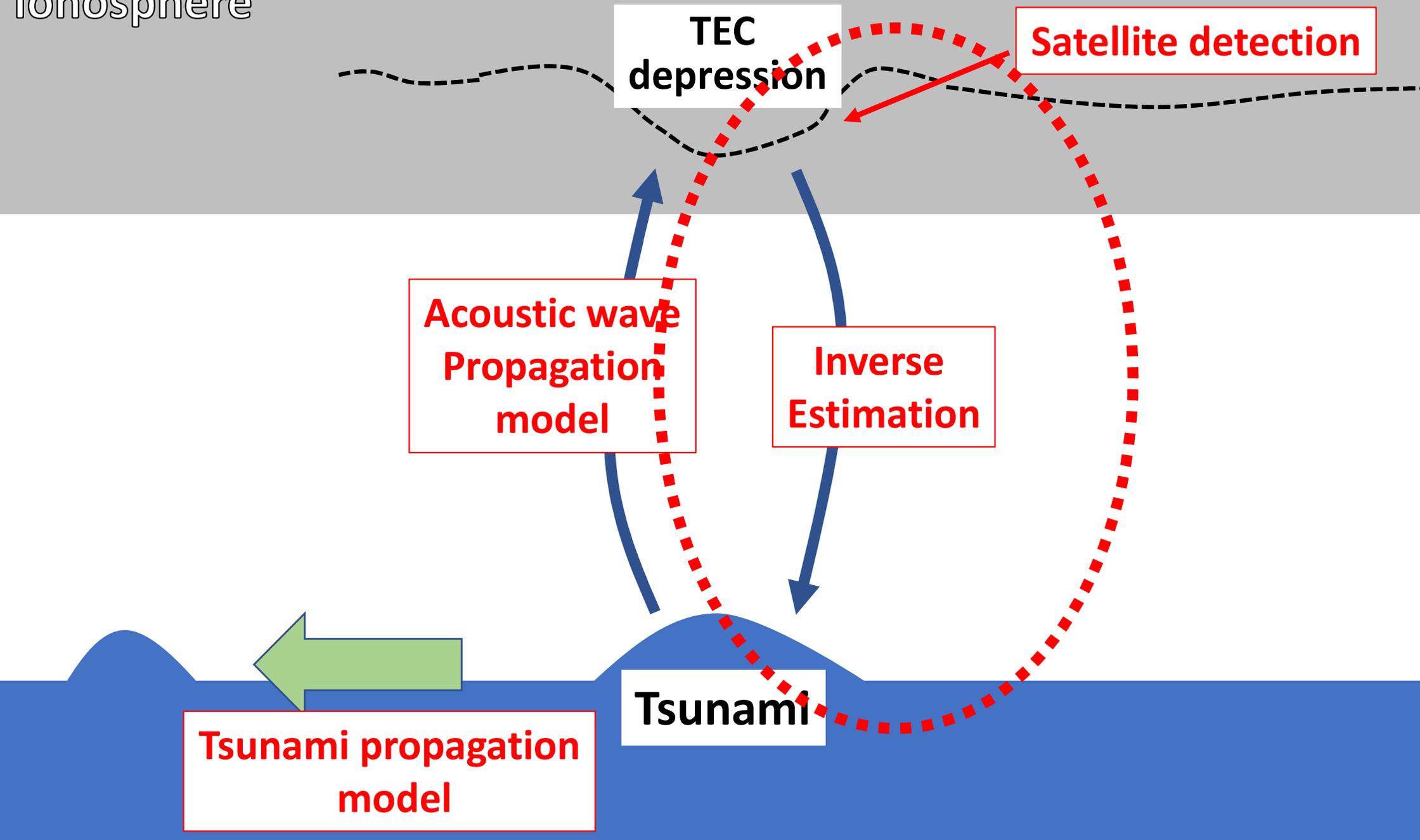
Ultimate goal

Ionosphere



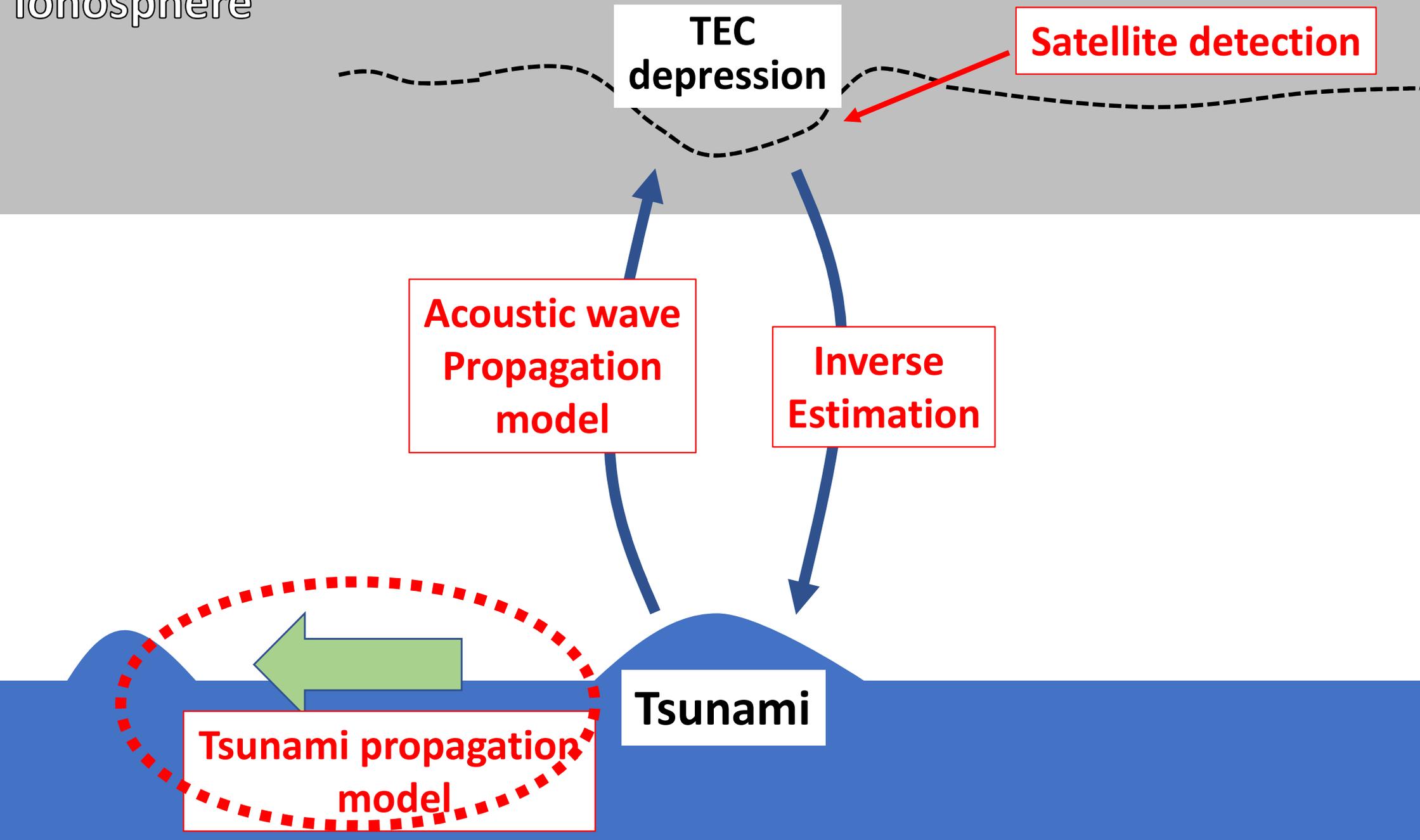
Ultimate goal

Ionosphere



Ultimate goal

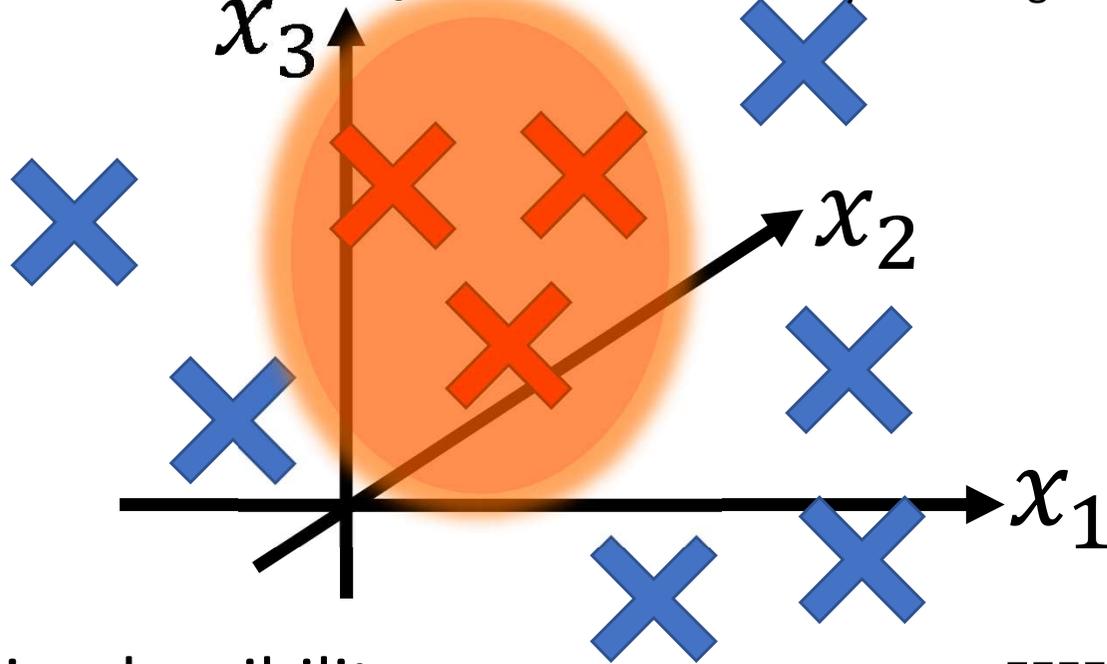
Ionosphere



Concept of History Matching

I. Vernon, Michael Goldstein, et. al., , *et al.* : "Galaxy Formation: Bayesian History Matching for the Observable Universe." *statistical science*, 29. (2014).

Parameter space



-  : implausible
-  : plausible
-  : plausible parameter subspace

I : implausibility

$$I(\theta) = \frac{\|z - E(\eta(\theta))\|}{\sqrt{\text{Var}(z - E(\eta(\theta)))}}$$

z : Observation

$E(\eta(\theta))$: Expectation of the model output

If the implausibility exceeds a certain threshold



The input parameter θ must be implausible

1 dimensional example

Observation = -0.8

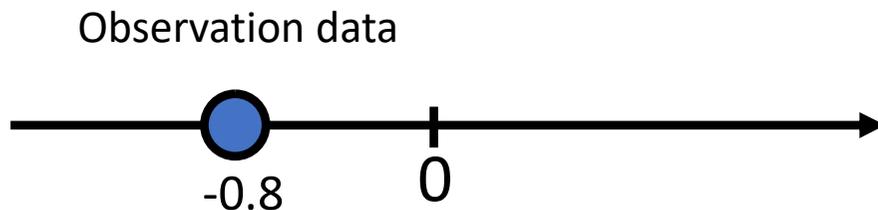
Observation error = 0.0004

Simulation model

$$f_{sim}(\theta)$$

This simulation model has
high computational cost.

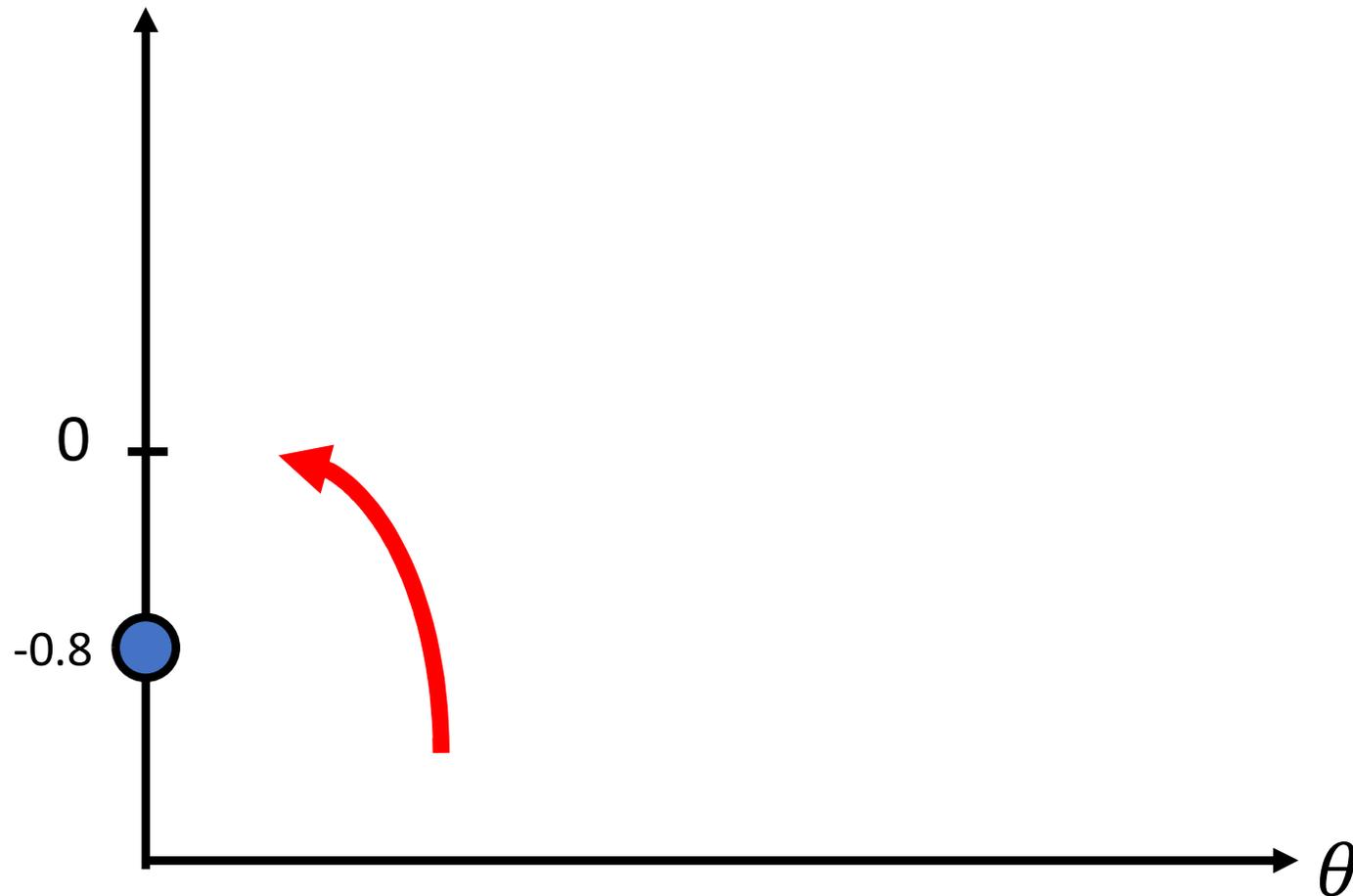
Only 6 points are computed.



1 dimensional example

Observation = -0.8

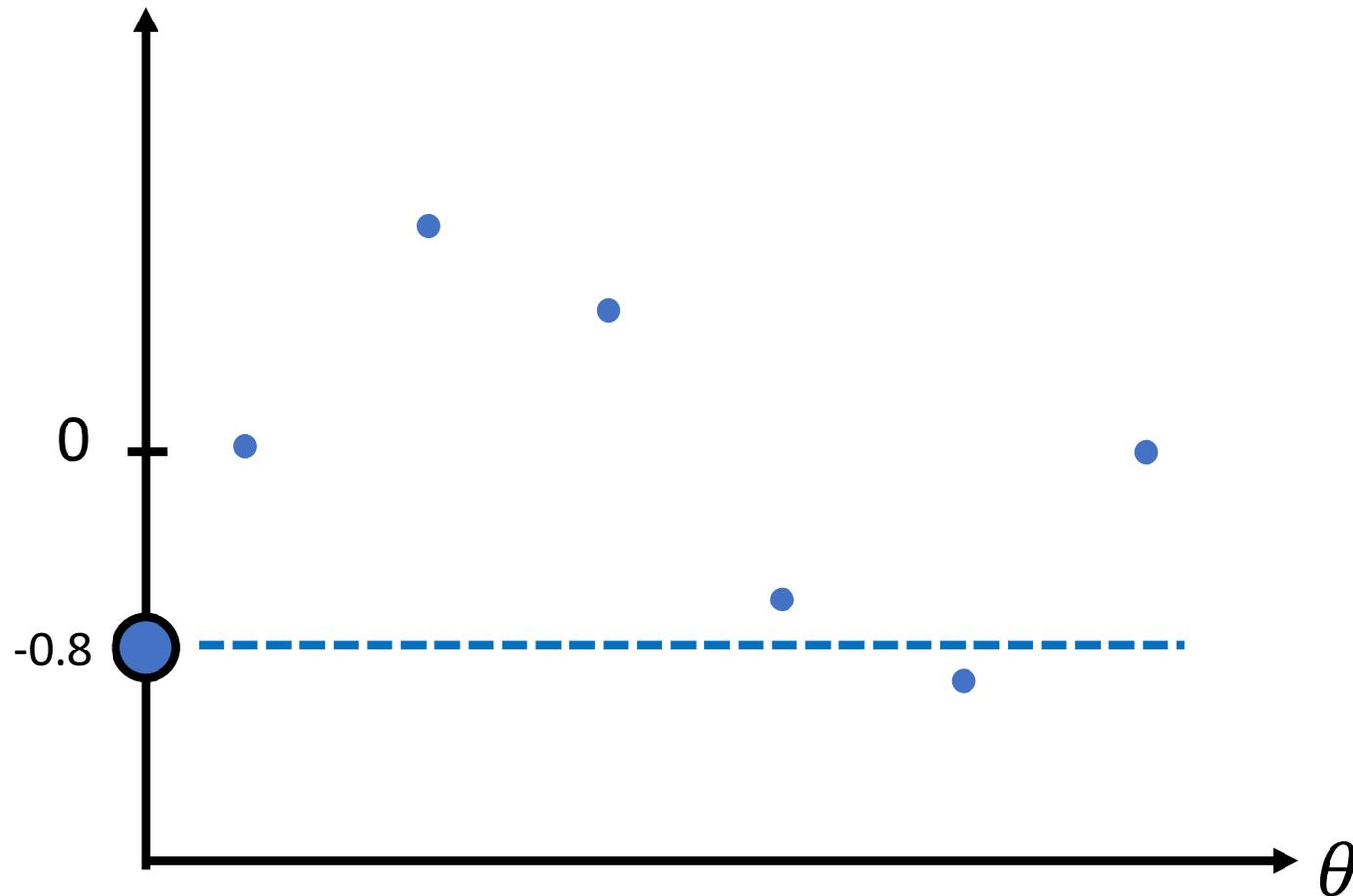
Observation error = 0.0004



1 dimensional example

Observation = -0.8

Observation error = 0.0004



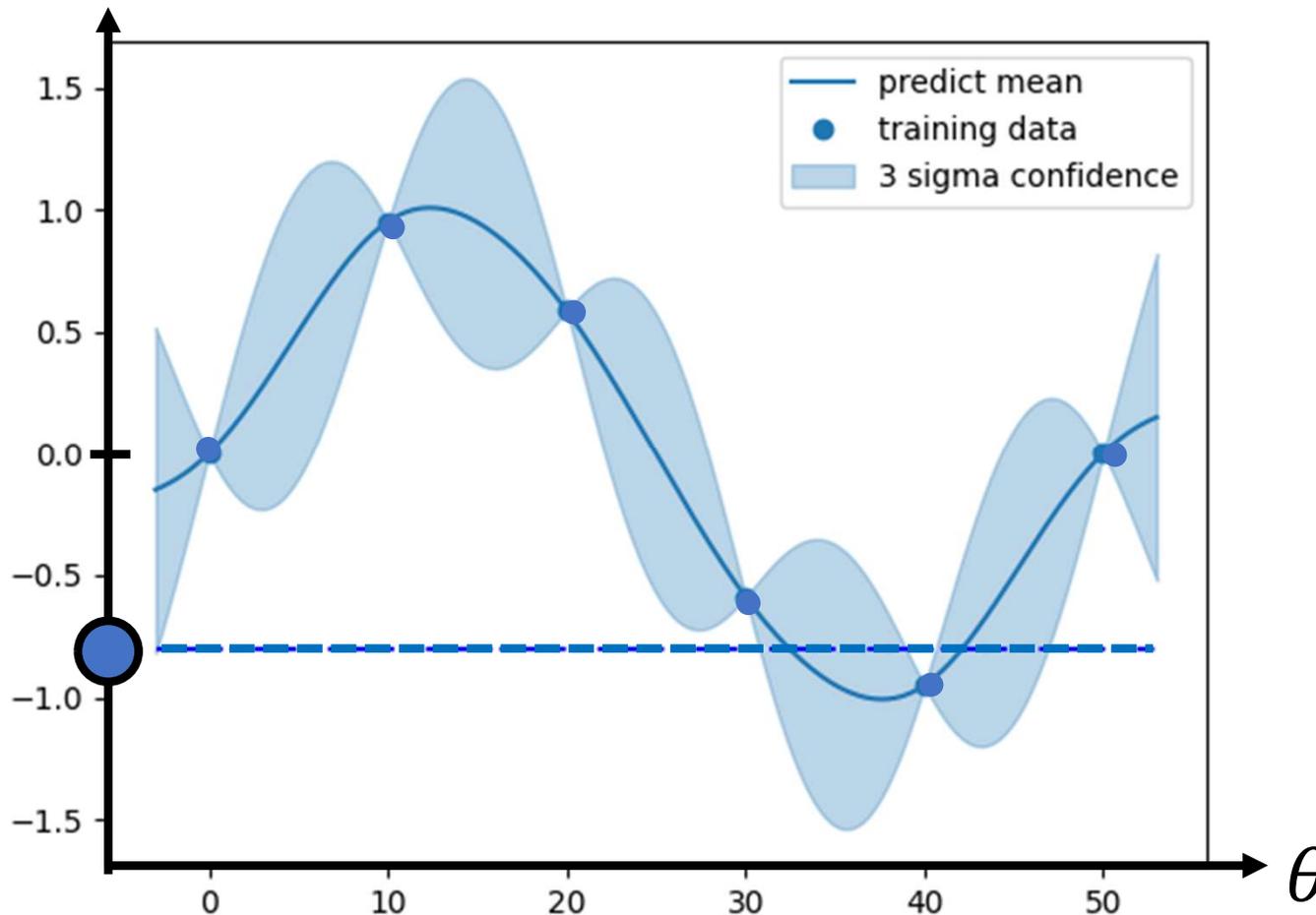
Toy Example (run an emulator)

1 dimensional example

Observation = -0.8
Observation error = 0.0004

Emulator

Gaussian process
Matern Kernel



Implausibility

$$I(\theta) = \frac{\|z - E(\eta(\theta))\|}{\sqrt{\text{Var}(z - E(\eta(\theta)))}}$$

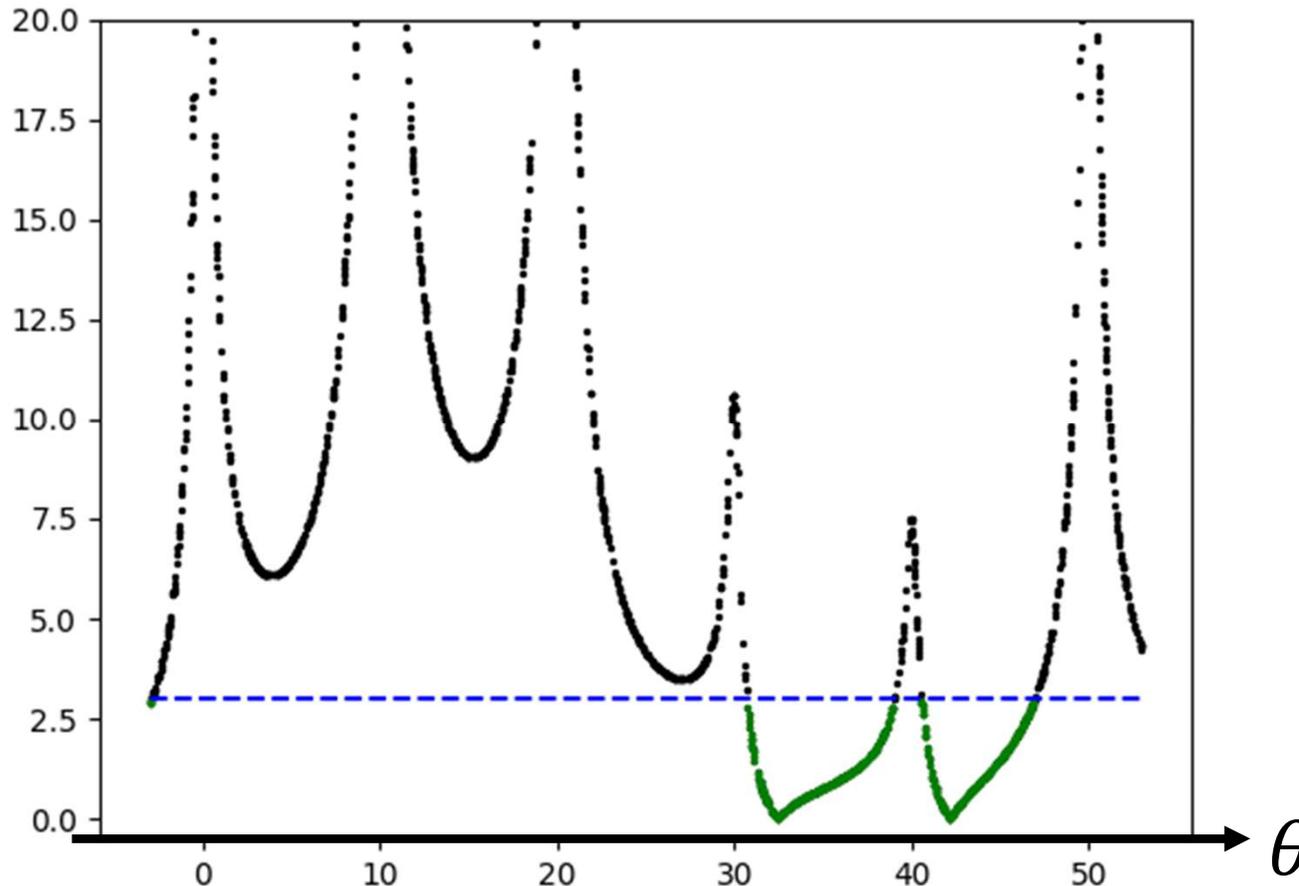
Toy Example (compute the implausibility)

1 dimensional example

Observation = -0.8
Observation error = 0.0004

Emulator

Gaussian process
Matern Kernel



Implausibility

$$\mathbf{I}(\boldsymbol{\theta}) = \frac{\|\mathbf{z} - \mathbf{E}(\boldsymbol{\eta}(\boldsymbol{\theta}))\|}{\sqrt{\mathbf{Var}(\mathbf{z} - \mathbf{E}(\boldsymbol{\eta}(\boldsymbol{\theta})))}}$$

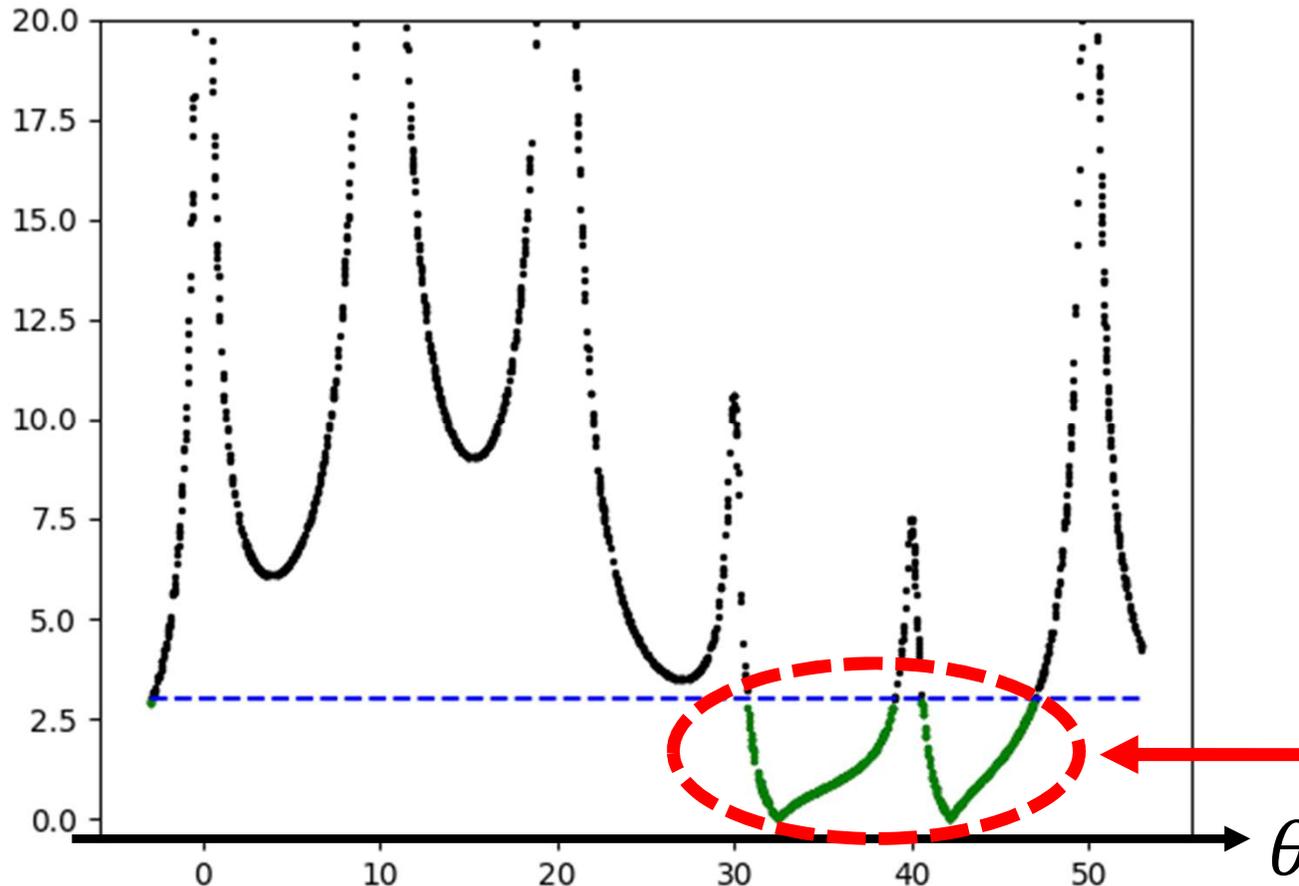
Toy Example (compute the implausibility)

1 dimensional example

Observation = -0.8
Observation error = 0.0004

Emulator

Gaussian process
Matern Kernel



Implausibility

$$I(\theta) = \frac{\|z - E(\eta(\theta))\|}{\sqrt{\text{Var}(z - E(\eta(\theta)))}}$$

Randomly choose an additional point to simulate

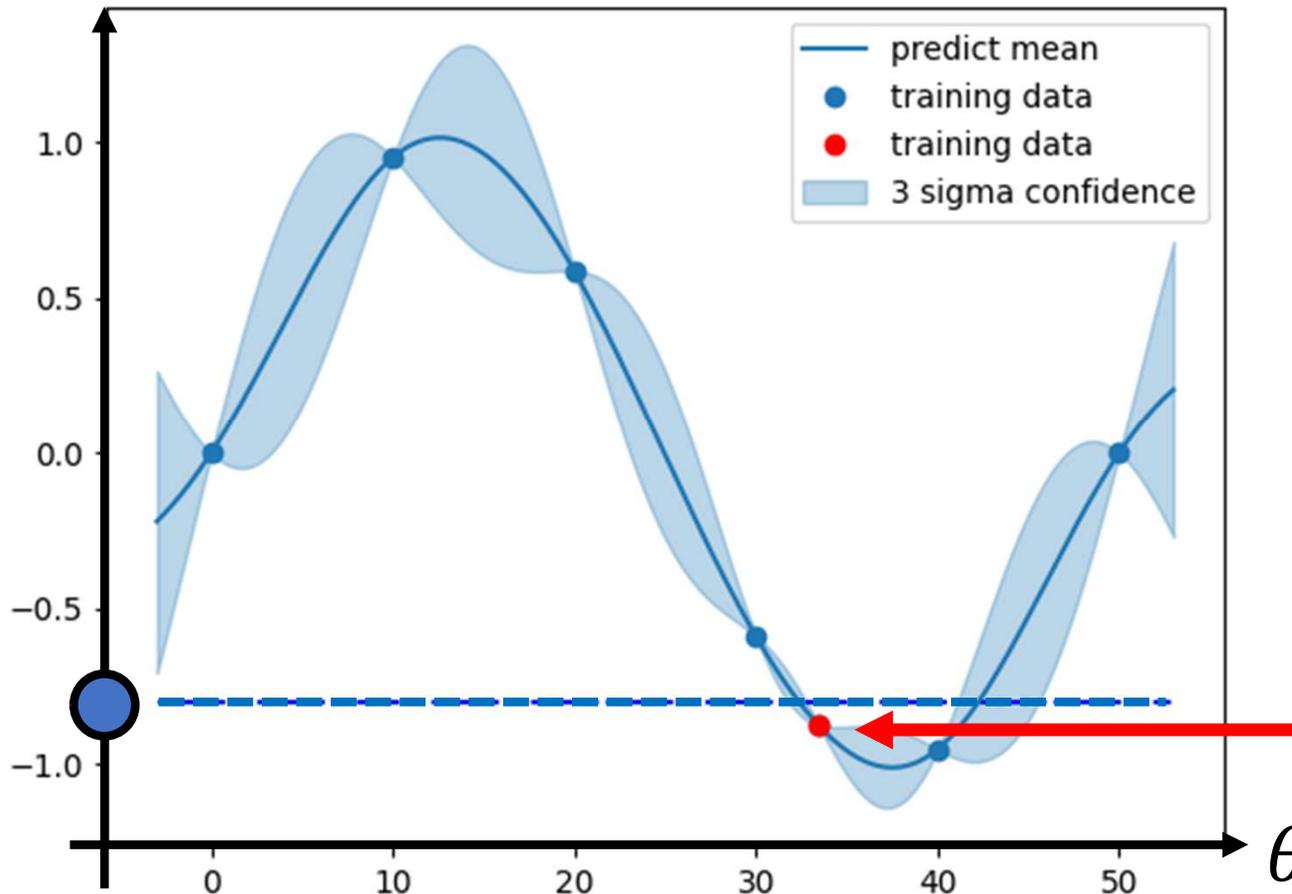
Toy Example (an additional point)

1 dimensional example

Emulator

Observation = -0.8
Observation error = 0.0004

Gaussian process
Matern Kernel



Implausibility

$$I(\theta) = \frac{\|z - E(\eta(\theta))\|}{\sqrt{\text{Var}(z - E(\eta(\theta)))}}$$

Randomly choose an additional point to simulate

Toy Example (compute the implausibility)

1 dimensional example

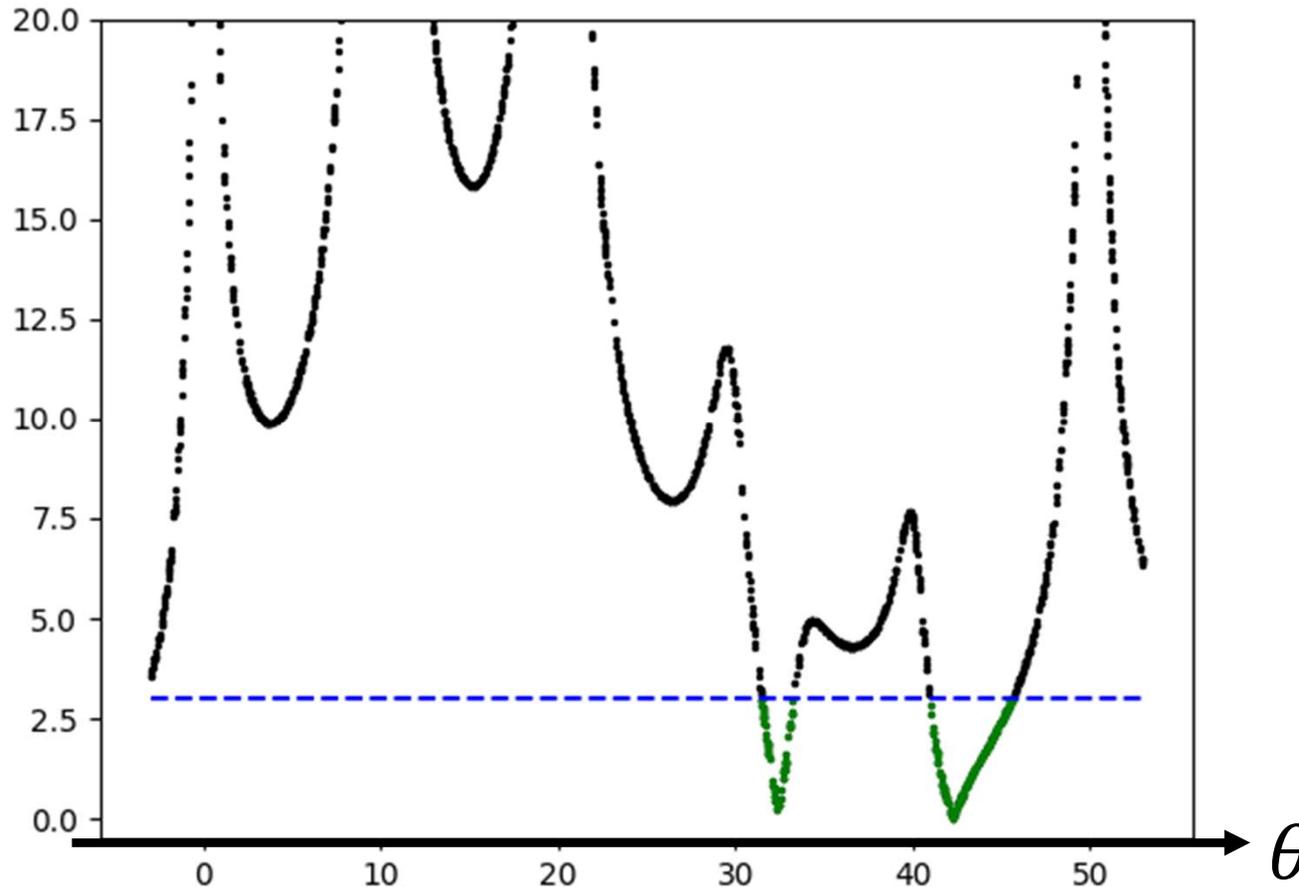
Observation = -0.8

Observation error = 0.0004

Emulator

Gaussian process

Matern Kernel



Implausibility

$$I(\boldsymbol{\theta}) = \frac{\|\mathbf{z} - \mathbf{E}(\boldsymbol{\eta}(\boldsymbol{\theta}))\|}{\sqrt{\text{Var}(\mathbf{z} - \mathbf{E}(\boldsymbol{\eta}(\boldsymbol{\theta})))}}$$

Thank you for your listening!

[Our result]

R. Kanai, M. Kamogawa, T. Nagao, A. Smith, and S. Guillas "Robust uncertainty quantification of the volume of tsunami ionospheric holes for the 2011 Tohoku-Oki Earthquake: towards low-cost satellite-based tsunami warning systems" (preprint)

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[Acoustic wave propagation]

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- M. D. Zettergren, et. al., "Latitude and Longitude Dependence of Ionospheric TEC and Magnetic Perturbations From Infrasonic-Acoustic Waves Generated by Strong Seismic Events", *Geophysical Research Letters*, 46, (2019)

[SPDE approach]

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