

θ Model parameters

Network weights w

Fig. 1: Overview of a Masked Autoencoder for Density Estimation (MADE) [1], a class of NDE that has shown great success in modeling probability densities by ensuring its output is a normalised probability distribution.

3. Data Compression

- The very high dimensionality of the displacement data **D**, which represents full-waveforms at each station, makes training a NDE **infeasible**.
- Data **D** must be compressed to $\dim(\theta)$ summary statistics **t**, replacing **D** in Fig. 1.

This study investigates two compression methods ϕ for seismic data:

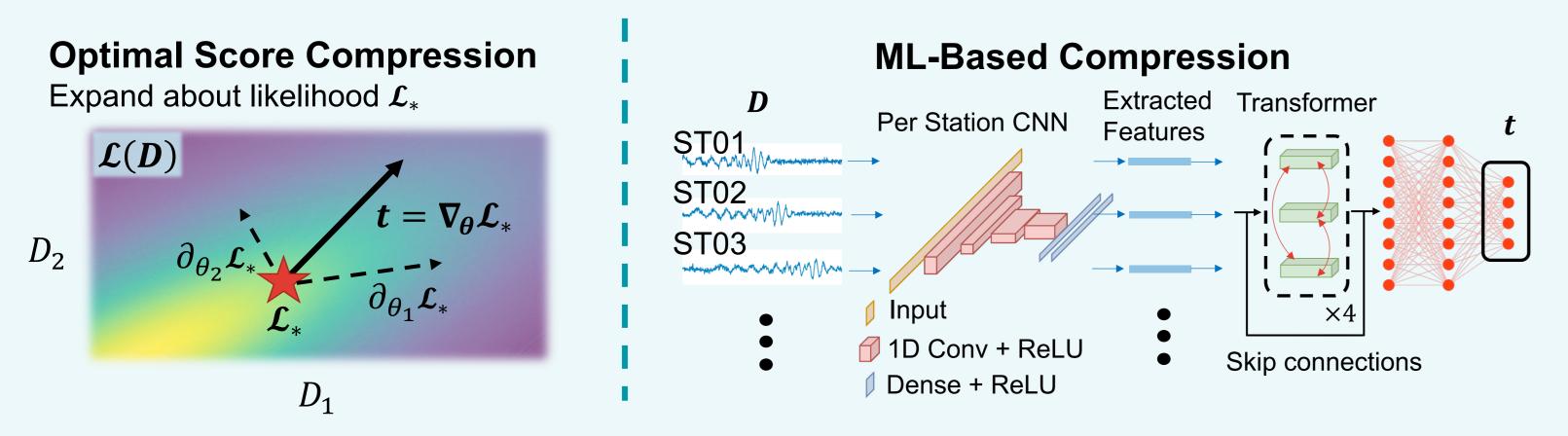


Fig. 2: Left: Classical technique for optimal compression using a first-order Taylor expansion, see [2]. Right: Neural network architecture, inspired by [3], trained to perform compression by learning to map $\phi: D \to \theta$.

References:

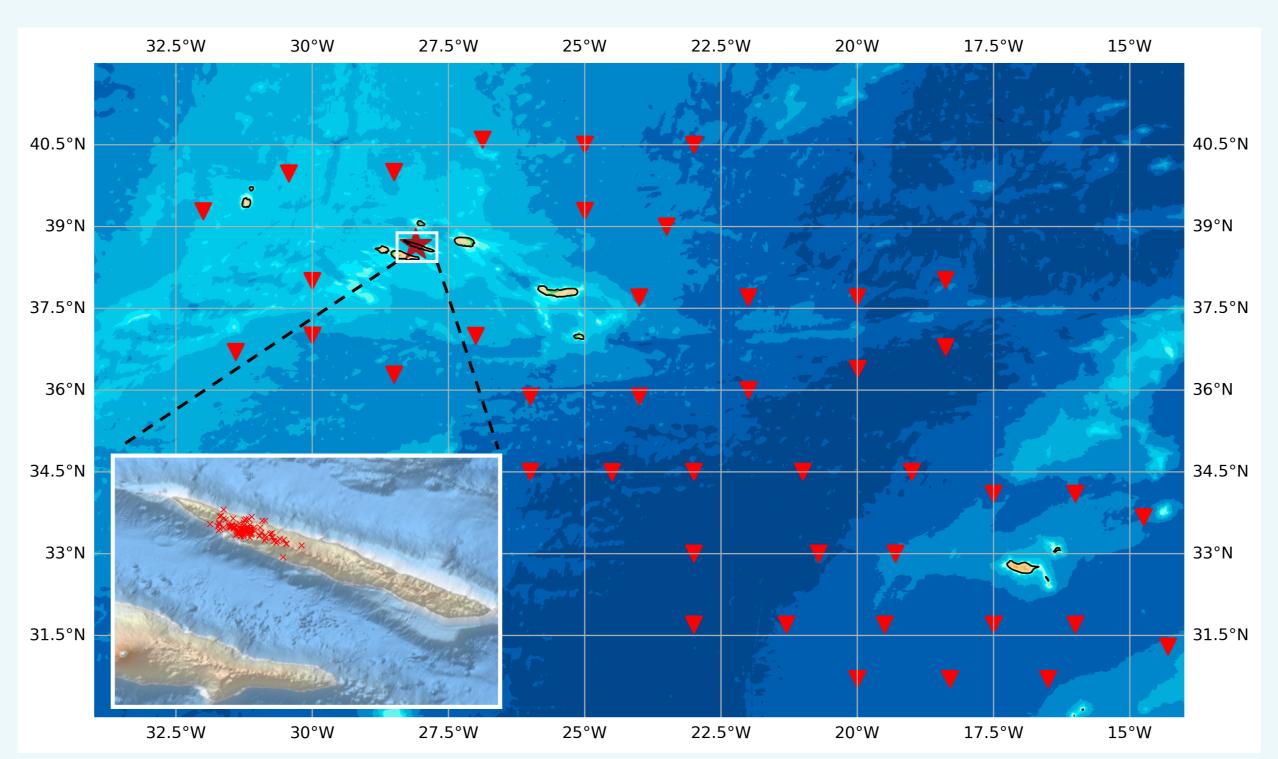
[5] Dziewonski & Anderson (1981), Phys. Earth Planet. Inter. - [6] IPMA, www.ipma.pt - [7] van Driel et al. (2015), Solid Earth.

Towards Bayesian Full-Waveform Source Inversion using Simulation-Based Inference

network:	$\psi(\boldsymbol{D}, \boldsymbol{ heta}; \boldsymbol{w})$
data:	$\mathcal{D} = \{\boldsymbol{\theta}_i, \boldsymbol{D}_i\}$
nction:	$\mathcal{L} = -\ln p(\mathbf{D} \boldsymbol{\theta})$

4. Example Simulation-Based Study

- We demonstrate SBI on a simplified problem of source-location inversion, i.e. $\theta = \{x, y, z, \Delta t\}$, a four parameter source inversion.
- Study region of the 2021 São Jorge crisis in the Azores, using the **UPFLOW 2021-22** [4] Ocean Bottom Seismometer (OBS) array.
- Entirely synthetic study, simulating events using the isotropic 1-D model PREM [5]. The simulations are accurate down to wave periods of $T \sim 2$ s.



 $p(\boldsymbol{D}|\boldsymbol{\theta}) = [p(D_i|D_{1:i-1}, \boldsymbol{\theta}; \boldsymbol{w})]$

Fig. 3: A schematic showing the experimental setup for this simulation-based study. We use OBS stations from the UPFLOW array (red triangular markers) to study the recent seismic swarm crisis located on São Jorge island (brown star). Bottom left: zoom in on the São Jorge crisis, with a small selection of events from the IPMA [6] catalogue.

6. Inversion Results

Samples from the posterior $p(\theta | t_{obs})$ are drawn using MCMC once the likelihood NDE is trained.

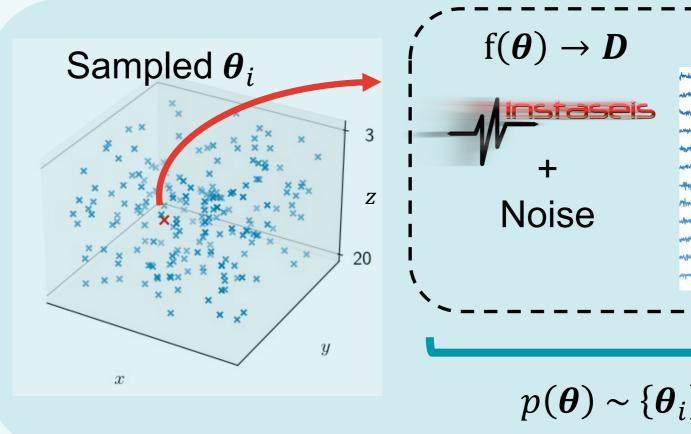
	Compression	
	Score	ML
Compression MSE \downarrow	2.6×10^{-2}	8.2×10 ⁻⁴
Posterior CRPS* ↓	2.6×10^{-2}	1.3×10^{-2}
Calibration Error ↓	0.35	0.31

Table improves ML-based method compression and yields tighter posteriors (lower CRPS). Both methods are relatively well calibrated.

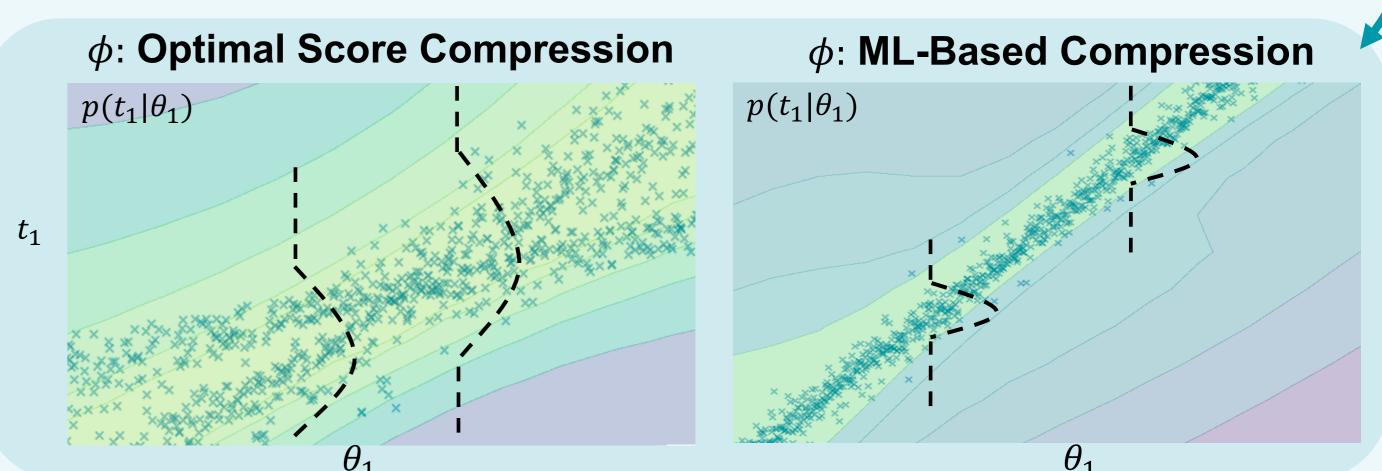
*Continuous Ranked Probability Score

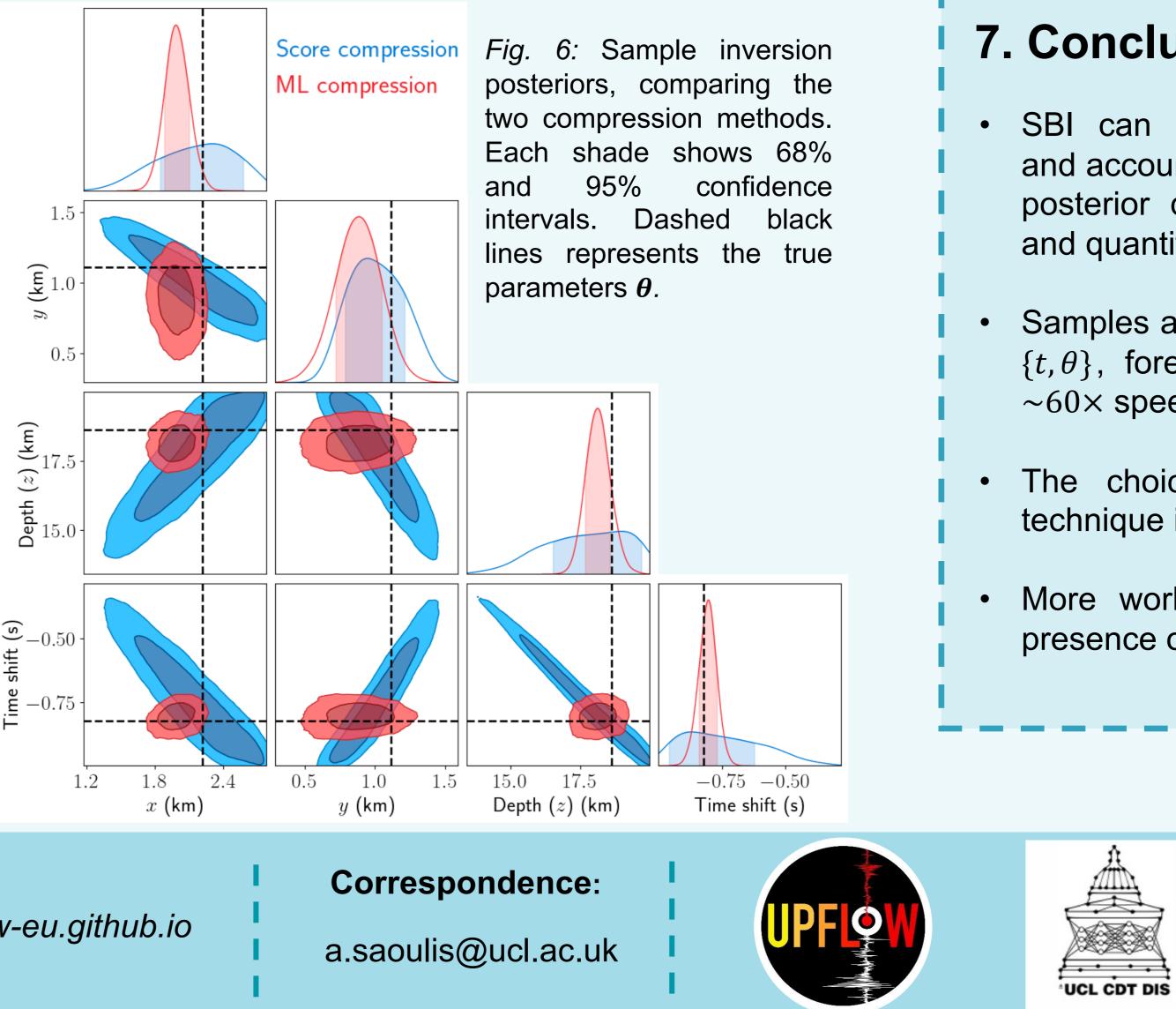
[1] Germain et al. (2015), CoRR - [2] Alsing & Wandelt (2018), MNRAS - [3] Münchmeyer et al. (2021), GJI - [4] UPFLOW project, upflow-eu.github.io

5. Simulation-Based Inference Workflow









Generate a Compressed Dataset

→ D → D → ise	<section-header></section-header>	Compressor $\phi(D) \rightarrow t$	Summary Statistic $\begin{pmatrix} t_1 \\ t_2 \\ t_3 \\ t_4 \end{pmatrix}$
$(\boldsymbol{\theta}) \sim \{\boldsymbol{\theta}\}$	$\boldsymbol{\theta}_i$	$\Rightarrow \qquad \mathcal{D} = \{ \boldsymbol{\theta} \}$	$\boldsymbol{\theta}_i, \boldsymbol{t}_i \}$

Fig 4. Events simulated with *Instaseis* [7]. Events are centred on the seismic swarm, sampled within a cube $(x, y, z) = 3 \text{ km} \times 3 \text{ km} \times 17 \text{ km}$, between depths 3 - 20 km. The time shift Δt is sampled between [-1,1] s. Synthetic noise is added to each waveform before compression.

Train NDE to Estimate Likelihood

Fig 5. NDE likelihood contours for each compression method. Score compression performs poorly due to displacement D non-linearity with respect to location parameters θ . ML-based compression gives sharper likelihood contours, corresponding to less lossy compression.

7. Conclusions

SBI can incorporate full-waveform information and account for non-Gaussian noise effects in the posterior distribution. Future work will address and quantify these advantages.

• Samples are generated in the compressed space $\{t, \theta\}$, foregoing the forward model and giving $\sim 60 \times$ speed-up over MCMC using *Instaseis*.

The choice and tuning of the compression technique is important.

• More work is needed to avoid failure in the presence of modelling errors.





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