

Water vapour line assignments in the 9250–26 000 cm⁻¹ frequency range

Roman N. Tolchenov^a, Olga Naumenko^b, Nikolai F. Zobov^{a,1}, Sergei V. Shirin^{a,1},
Oleg L. Polyansky^{c,1}, Jonathan Tennyson^{a,*}, Michel Carleer^d, Pierre-François Coheur^d,
Sophie Fally^d, Alain Jenouvrier^e, Ann Carine Vandaele^f

^a Department of Physics and Astronomy, University College London, London WC1E 6BT, UK

^b Russian Academy of Science, Institute of Atmospheric Optics, Tomsk 634055, Russia

^c Sektion Spektren und Strukturdokumentation, University of Ulm, D-89069 Ulm, Germany

^d Service de Chimie Quantique et Photophysique CP160/09, Université Libre de Bruxelles, 50 Av. F.D. Roosevelt, B-1050 Brussels, Belgium

^e Université de Reims Champagne-Ardenne, Groupe de Spectrométrie Moléculaire et Atmosphérique, UFR Sciences,

Moulin de la Housse B.P.1039, 51687 Reims Cedex 2, France

^f Institut d'Aéronomie Spatiale de Belgique, 3 Av. Circulaire, B-1180 Brussels, Belgium

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Abstract

Line parameters for water vapour in natural abundance have recently been determined for the 9250–13 000 cm⁻¹ region [M.-F. Mérienne, A. Jenouvrier, C. Hermans, A.C. Vandaele, M. Carleer, C. Clerbaux, P.-F. Coheur, R. Colin, S. Fally, M. Bach, J. Quant. Spectrosc. Radiat. Transfer 82 (2003) 99] and the 13 000–26 000 cm⁻¹ region [P.-F. Coheur, S. Fally, M. Carleer, C. Clerbaux, R. Colin, A. Jenouvrier, M.-F. Mérienne, C. Hermans, A.C. Vandaele, J. Quant. Spectrosc. Radiat. Transfer 74 (2002) 493] using a high-resolution Fourier-transform spectrometer with a long-path absorption cell. These spectra are analysed using several techniques including variational line lists and assignments made. In total, over 15 000 lines were assigned to transitions involving more than 150 excited vibrational states of H₂¹⁶O. Twelve new vibrational band origins are determined and estimates for a further 16 are presented.

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1. Introduction

Understanding the rotation–vibration spectrum of water is central to constructing reliable models of the transmission of light through our atmosphere. Yet, particularly at the higher frequencies crucial for many atmospheric processes, this spectrum remains difficult to both characterise and interpret. Recently, some of

us [1–3] have collaborated in a series of experiments to characterise the line parameters of water using a high-resolution Fourier-transform spectrometer with a long-path absorption cell. These studies, which considered the ranges 9250–13 000 cm⁻¹ and 13 000–26 000 cm⁻¹, covered the entire visible region as well as extending into the near-infrared. The spectra were recorded by the Bruxelles-Reims groups and are collectively referred to as BR below.

A considerable number of previous laboratory measurements have been made of the near-infrared [4–8] and visible [7–10] wavelength spectra of water vapour in natural isotopic abundance. The complexity of weak

* Corresponding author. Fax: +44 20 7679 7145.

E-mail address: j.tennyson@ucl.ac.uk (J. Tennyson).

¹ Permanent address: Institute of Applied Physics, Russian Academy of Science, Uljanov Street 46, Nizhniy Novgorod 603950, Russia.

rotation–vibration spectrum of water at these wavelengths has encouraged the development of assignment procedures based on variational nuclear motion calculations [11,12]. These have been widely applied to the spectrum of water vapour at near-infrared and visible wavelengths [13–19]. The result has been a considerable advance in our understanding of the spectroscopy of water in these regions, see [20] for example.

The new Fourier-transform spectra contained 7061 lines in the 9250–13000 cm^{-1} and 9353 lines in the 13000–26000 cm^{-1} region, although, as discussed below, these lines contain a number of blends. Quantum number assignments for these transitions were based simply on the analysis of other spectra discussed above. However, since the new systematic spectra have already been included in the 2003 edition of the GEISA database [22] and is being prepared for inclusion in the 2004 edition of HITRAN [21], it is important that as complete and correct as possible set of assignments are made to these. In this work, we present a comprehensive analysis of the line assignments performed using a variety of methods but largely based on line lists generated using variational nuclear motion calculations. One advantage of analysing the present spectra is the broad spectral range studied which significantly reduces problems with edge effects when searching for combination differences. However, for convenience, much of the analysis was performed separately for the 9250–13000, 13000–16000, and 16000–26000 cm^{-1} regions, and these three regions will be considered separately below.

2. Experimental details

Detailed descriptions of the experimental setup and procedures have been given previously [2,15]. Briefly, the absorption spectra of water vapour were recorded using a high-resolution Fourier-transform Spectrometer (Bruker IFS120M) coupled to two White multiple-reflection cells of 5 and 50 m base path. The combination of two light sources (high-pressure Xenon arc and tungsten halogen) and two detectors (Si and GaP-diode) coupled to optical filters enabled us to cover the spectral region from 27000 to 8000 cm^{-1} in three overlapping ranges. Spectra with an absorption path of 600 m and a resolution of 0.06 cm^{-1} (15 cm MOPD) were recorded in the near-UV and visible regions using the Xenon arc lamp. A resolution of 0.03 cm^{-1} (MOPD of 30 cm) and absorption paths of 100 and 600 m were chosen to re-

cord the visible-NIR region using the tungsten halogen lamp. The temperature inside the cell was measured with three platinum wire thermometers which gave a ‘room temperature’ value of 292 ± 3 K. Table 1 summarises the spectra analysed here.

The co-addition of some 2000–4000 interferograms, leading to measurement times of 12–24 h, proved sufficient to get a signal-to-noise ratio (expressed as the maximum signal amplitude divided by twice the root-mean-square noise amplitude) better than 2500. The spectra were wavenumber calibrated using the I_2 visible spectrum. Pressure was measured with an MKS Baratron type manometer. Pure water vapour spectra were measured at various pressures from 6 to 18 hPa into the absorption cell. Additional spectra were recorded after adding N_2 or dry-air in the cell in four steps up to about 800 hPa, in order to measure the pressure effect on the lines shape and position. Data concerned with the line profiles have been presented previously [1–3] and are not re-analysed here.

Spectral line positions and other parameters were determined by fitting the lines with a Voigt profile using the procedure discussed previously [1–3]. The spectrum was treated in two separate regions: 9250–12896 and 13184–25232 cm^{-1} . No transitions were observed above this frequency. The gap between 12896 and 13184 cm^{-1} was found to contain no H_2O lines although some transitions which could be assigned to the O_2 A-band and to HDO were observed.

The experimental line list contains 7061 lines in the 9250–12896 cm^{-1} region and 9353 between 13184 and 25232 cm^{-1} . In practise, particularly at lower frequencies, some of these lines are blends of more than one transition so the assigned line list presented here is somewhat longer. On the other hand, we have chosen to be more conservative than the original work of [1] in the higher-frequency region, and to remove from the final list a set of very weak lines, which could have been misidentified. Similarly, about 25 lines around 14500 cm^{-1} have been identified as probably due to transitions of O_2 . The O_2 lines are listed in Table 2 and have been removed from our final line list.

3. Line assignments

Line assignments were performed in both London and Tomsk and iterated until an agreed set of parameters was obtained. The main means of assignment was

Table 1
Summary of experimental spectra analysed

Region	Wavenumber (cm^{-1})	Lamp	Detector	Resolution (cm^{-1})	Path (m)
Near-UV–VIS	27000–17000	Xe	GaP	0.06	600
VIS	23000–10000	Xe	Si	0.06	600
VIS–NIR	15500–8000	W	Si	0.03	100 and 600

Table 2

Transitions, in cm^{-1} , due to molecular oxygen identified in the long path-length Fourier-transform spectrum of Coheur et al. [1]

14468.8831	14468.9968	14478.2308	14480.1454	14486.9391
14495.1269	14502.8207	14504.7900	14509.9693	14518.7222
14522.7633	14531.0275	14535.7825	14535.9398	14544.0015
14545.9334	14549.2909	14550.1009	14552.2095	14554.3357
14555.3237	14555.7390	14556.5596	14557.8410	14557.9817

by comparison with line lists calculated using variational nuclear motion calculations. In the course of this work, more than three distinct line lists were used for this purpose. The first was the line list due to Partridge and Schwenke (PS) [12], which gives excellent frequency predictions for transitions below about 15000 cm^{-1} but tends to overestimate the intensity of some transitions. The second was the line list recalculated by Tashkun (private communication) using the potential and parameters of PS, and the refitted dipole surface of Schwenke and Partridge [23]. This line list, denoted SP below, gives greatly improved intensities but remains unreliable for high frequencies. The third variational line list was a newly calculated line list based on the Fit B potential energy surface of Shirin et al. [24], the wavefunctions of Barber and Tennyson [25], and the dipole surface of Schwenke and Partridge. This line list, known as BT2, gives reliable frequencies for the whole range studied and, in particular at the higher frequencies, see Dupré et al. [26] for example.

The above-mentioned synthetic line lists were used as the input information for an expert system for automatic identification of rovibrational spectra [27] which was used in the assignment process. This system is based on the Rydberg–Ritz combination rule. It searches for combination differences in a large spectral interval, then chooses the best variant and, if approved by the researcher, inserts the corresponding assignments into the database for spectrum. Up to three assignments for the same absorption line are possible. At the final stage of the spectral assignment, when the database is almost completed, it is also possible to identify single lines not included among the ground-state combination differences, if the match between observed and calculated line position and intensity is believed to be reasonable.

In the course of the theoretical analysis, it became apparent that, particularly at the lower frequencies, many of the ‘lines’ identified in the original spectral studies were actually blends of two or more transitions. This finding is consistent with the somewhat anomalous line-widths and associated pressure dependence found for many of the lines. It was therefore decided to treat as blends all ‘lines’ which, in the theoretical calculations, were not dominated by a single transition.

Blended lines are identified in our final line list by the presence of more than one transition at a given frequency. The 296 K intensity of these transitions sums to that

of the single ‘line’ given in the original BR line list but has been distributed between the lines using the theoretical intensity ratios. For this purpose, the SP and the BT2 line lists were used below and above 16000 cm^{-1} , respectively. Clearly, the use of a single frequency for all the lines in a blend is an approximation, but none of the theoretical line lists are accurate enough to be used to determine the (small) splittings between the lines. These can only be determined by careful fits to high-resolution experiments. Altogether considering blends increased the number of lines in the spectrum by about 11%.

When considering the assignment procedure below it is necessary to distinguish between assignment, which is the process of identifying the upper and lower states involved in a transition, and labelling, which is the process of associating each state with (approximate) quantum numbers. For the rotation–vibration states of H_2^{16}O the only rigorous quantum numbers are the rotational angular momentum, J , the rotational parity p and the interchange symmetry between the two H atoms, which determines whether the state is ortho or para. These quantum numbers are given rigorously by the various line lists used in this study. However, it is usual to label each state with vibrational quantum numbers, in either normal or local mode notation, and rotational quantum numbers K_a and K_c .

In the course of analysing the present spectrum, we found a number of difficulties with the labelling used in previous studies. In particular, we found cases where the correct assignments had been obtained but the labels were inconsistent so, for example, two different states had been given the same quantum number labels. We also had some initial difficulty finding the correct labels for many of our new assignments. We therefore undertook a systematic (re-)labelling exercise based on analysis of the BT2 line list [25].

First, plots of the $J=0$ vibrational wavefunctions were analysed to assign vibrational quantum numbers. Rotational levels associated with a vibrational state were identified using transition moments. To do this it was assumed that dipole transitions must be strongest between rovibrational levels belonging to the same vibrational state. Knowing labels for a J level gives candidates for the levels in the same vibrational state with $J' = J + 1$. However, the maximum $J \leftrightarrow J + 1$ transition moment does not necessarily provide the correct label

for the $J + 1$ level. As an approximate criterion it was assumed that the sum of the dipoles between levels within the same vibrational state must be the greatest. This procedure had to be iterated to get complete results.

It should be noted that one result of this procedure is the identification, discussed below, of transitions to many previous unobserved vibrational states. These new states are often characterised by only a few levels which are probably observed due to intensity stealing from a nearby bright state of the same symmetry.

The BR lines were published using previously known line assignments where available. In order to check and extend the assignments, we analysed the spectrum in three regions: 9250–13000, 13000–16000, and 16000–26000 cm^{-1} . The split between the first and second region comes from the original experiments, and the break at 16000 cm^{-1} represents approximately the point where our new line list becomes more reliable than the SP one. Table 3 summarises the results of analysis in each of these regions. The region below 11190 cm^{-1} has been the subject of studies [6,18,19,28–30] which actually probe deeper than the BR results. The results for this region are therefore of less significance. Table 3 therefore also presents separate summaries of our results for this low-frequency region and the higher-frequency region for which the BR data now form the input for the HI-TRAN database [21].

In summarising the data in Table 3 we make a distinction between a reassignment, where a line is assigned to a completely distinct transition, and a re-labelling, where the upper level in a transition is of unchanged symmetry but has different (approximate) vibrational quantum numbers. It can be seen that we have made a large number of both new assignments and re-assignments.

The BR spectra were recorded using water vapour with natural isotopic abundance. The spectrum there-

fore contains a number of lines not due to H_2^{16}O . These lines were identified by comparison with spectra of isotopically enriched spectra for H_2^{17}O [31,35], H_2^{18}O [32–35], and HDO [36–39].

Assignments are vital for determining the temperature-dependent absorption intensity of each line. They also yield information on the energy levels of the system. A complete set of the energy levels determined by analysing the BR spectra, which covers 150 vibrational states, is given in the electronic archive. Energy levels were determined from known ground-state levels [20] in two steps. First, using only unblended lines we calculated mean energies for those upper states which could be found with more than two transitions. Using the deviations from the mean values we estimated the error distribution with the line intensity and represented it as a linear function of intensity. In the second step, we calculated the final set of energy levels using this error distribution to weight the lines when evaluating the mean values. A comprehensive compilation of H_2^{16}O energy levels was recently presented by Tennyson et al. [20] and we therefore concentrate only on differences from that compilation. Table 4 gives a numerical comparison of the newly determined energy levels. A complete listing of these new levels is given in the journal's electronic archive. We note that levels determined from blended lines are inherently less accurate so they are enumerated separately.

We made new assignments to the rotational ground levels of 12 excited vibrational bands. Also assignments to low-lying rotational levels for 16 other vibrational states allow us to obtain estimates for the vibrational band origins of these states guided by systematic differences between theory and experiment. This method of estimation has proved reliable in the past [16].

As implied above, a blended line will have a somewhat different line profile than a single line of the same

Table 3
Summary of our assignments of BR long path-length Fourier-transform water vapour spectrum by spectral region, in cm^{-1} , and isotopologue

Region start (cm^{-1}):	9250	13184	16000	9250	11190	9250
Region end (cm^{-1}):	12900	16000	25300	11190	25300	25300
Original lines	7061	4755	4573	5534	10855	16389
Extra blended lines	985	409	412	851	953	1804
Total lines	8046	5162	4985	6385	11808	18193
Observed lines assigned	6573	4158	4279	5223	9787	15010
New assignments	2193	558	1325	1732	2344	4076
Reassignments	916	130	137	832	351	1183
Relabelled	111	263	579	96	857	953
De-assigned	60	41	16	43	74	117
Isotopologue	Total lines		Assigned			
H_2^{16}O	16990		15621			
H_2^{18}O	736		734			
H_2^{17}O	255		246			
HDO	212		212			

Table 4
Summary of newly determined energy levels

$v_1v_2v_3$	$mn^{\pm}v_2$	ω_0 (cm ⁻¹)	Levels		
			Ref. [20]	New ^a	New ^b
050	00 5	7542.44	105	2	4
130	10 ⁺ 3	8273.98	55	3	4
031	10 ⁻ 3	8373.85	229	5	7
210	20 ⁺ 1	8761.58	58	36	39
111	20 ⁻ 1	8807.00	209	16	23
012	11 1	9000.14	148	25	30
060	00 6	8870.6 ^c	37	11	13
140	10 ⁺ 4	9724.2	0	99	105
041	10 ⁻ 4	9833.59	144	33	42
220	20 ⁺ 2	10284.37	33	87	91
121	20 ⁻ 2	10328.73	86	56	63
022	11 2	10521.76 ^c	47	79	91
300	30 ⁺ 0	10599.69	83	54	60
201	30 ⁺ 0	10613.35	120	48	59
102	21 ⁺ 0	10868.88	73	49	55
003	21 ⁻ 0	11032.41	95	44	55
070	00 7	10086.0 ^c	3	25	28
051	10 ⁻ 5	11242.8	64	4	6
230	20 ⁺ 3	11767.39	37	8	11
131	20 ⁻ 3	11813.21	85	8	13
032	11 3	12007.78	42	9	15
310	30 ⁺ 1	12139.32	73	9	13
211	30 ⁻ 1	12151.25	122	1	4
112	21 ⁺ 1	12407.66	72	5	10
013	21 ⁻ 1	12565.01	88	3	6
160	10 ⁺ 6	12380.4 ^c	3	3	3
061	10 ⁻ 6	12586.	26	3	5
240	20 ⁺ 4	13204.8 ^c	44	25	28
141	20 ⁻ 4	13256.2	49	8	12
042	11 4	13453.6	39	4	4
320	30 ⁺ 2	13640.7	77	10	15
221	30 ⁻ 2	13652.67	96	18	21
122	21 ⁺ 2	13910.90	66	14	18
023	21 ⁻ 2	14066.19	83	13	16
400	40 ⁺ 0	13828.28	100	10	11
301	40 ⁻ 0	13830.94	121	16	19
202	31 ⁺ 0	14221.16	102	9	11
103	31 ⁻ 0	14318.81	114	6	8
180	10 ⁺ 8		0	5	6
004	22 0	14537.50	72	5	8
170	10 ⁺ 7	13660.6 ^c	6	3	4
071	10 ⁻ 7	13835.37	12	5	6
250	20 ⁺ 5		1	9	12
151	20 ⁻ 5	14647.98 ^c	23	10	13
052	11 5		1	6	6
330	30 ⁺ 3	15108.24	30	24	31
231	30 ⁻ 3	15119.03	72	8	12
132	21 ⁺ 3	15377.7	19	14	19
033	21 ⁻ 3	15534.71	64	6	8
410	40 ⁺ 1	15344.50	74	11	13
311	40 ⁻ 1	15347.96	93	8	9
212	31 ⁺ 1	15742.80	58	13	14
113	31 ⁻ 1	15832.77	85	10	12
260	20 ⁺ 6		0	5	6
161	20 ⁻ 6		0	7	7
014	22 1		0	6	6
062	11 6		0	8	8
340	30 ⁺ 4	16534.3	30	18	20
241	30 ⁻ 4	16546.32 ^c	35	20	20
142	21 ⁺ 4	16795.9 ^c	40	40	46
420	40 ⁺ 2	16823.32 ^c	51	20	21
321	40 ⁻ 2	16821.63	62	23	26
500	50 ⁺ 0	16898.4 ^c	81	10	12
401	50 ⁻ 0	16898.84	87	5	6
043	21 ⁻ 4	16967.6 ^c	18	11	11

Table 4 (continued)

$v_1v_2v_3$	$mn^{\pm}v_2$	ω_0 (cm ⁻¹)	Levels		
			Ref. [20]	New ^a	New ^b
270	20 ⁺ 7		0	6	6
222	31 ⁺ 2	17227.38 ^c	40	11	15
123	31 ⁻ 2	17312.54	45	23	28
302	41 ⁺ 0	17458.35	63	11	12
203	41 ⁻ 0	17495.53	79	10	13
024	22 2		0	5	6
104	22 ⁺ 0	17748.11 ^c	46	3	4
350	30 ⁺ 5		0	9	9
251	30 ⁻ 5		0	7	8
152	21 ⁺ 5		0	12	12
331	40 ⁻ 3	18265.82	50	12	12
430	40 ⁺ 3	18267.1 ^c	4	45	8
053	21 ⁻ 5		11	10	10
510	50 ⁺ 1	18392.97	34	25	25
411	50 ⁻ 1	18393.31	50	9	10
133	31 ⁻ 3	18758.63 ^c	23	18	21
312	41 ⁺ 1	0	10	10	
034	22 2		8	4	4
213	41 ⁻ 1	18989.96	46	7	7
360	30 ⁺ 6		0	7	7
261	30 ⁻ 6		0	7	7
440	40 ⁺ 4	19677.8	0	31	34
341	40 ⁻ 4	19679.19 ^c	31	11	11
063	21 ⁻ 6		3	5	6
600	60 ⁺ 0	19781.32 ^c	35	34	37
501	60 ⁻ 0	19781.10	53	20	20
520	50 ⁺ 2	19864.7 ^c	11	13	15
421	50 ⁻ 2	19865.28 ^c	15	4	5
044	22 4		0	11	14
322	41 ⁺ 2		0	9	9
223	41 ⁻ 2	20442.78 ^c	13	17	18
271	30 ⁻ 7		0	4	5
402	51 ⁺ 0	20533.4 ^c	25	13	13
303	51 ⁻ 0	20543.14	32	9	9
105	42 ⁻ 0		0	3	3
450	40 ⁺ 5		0	10	11
351	40 ⁻ 5		0	9	10
530	50 ⁺ 3		0	15	16
431	50 ⁻ 3	21314.45 ^c	11	10	10
610	60 ⁺ 1	21221.5 ^c	9	32	34
511	60 ⁻ 1	21221.83	22	18	19
313	51 ⁻ 1	22015.4 ^c	0	8	8
460	40 ⁺ 6		0	6	7
361	40 ⁻ 6		0	5	5
115	42 ⁻ 1	22508.3 ^c	5	25	26
700	70 ⁺ 0	22529.29	42	20	22
601	70 ⁻ 0	22529.44	37	13	13
422	51 ⁺ 2		0	4	4
403	61 ⁻ 0		0	7	7
125	42 ⁻ 2		0	9	9
710	70 ⁺ 1		0	7	7
611	70 ⁻ 1	23947.1 ^c	9	1	2
800	80 ⁺ 0	25120.3 ^c	20	9	10
701	80 ⁻ 0	25120.28	24	5	6
Total			12258 ^d	1870 ^d	2140 ^d

Given are vibrational band origin, ω_0 , number of rotational term values determined, and numbers of new levels. Band origins given to less than 2 decimal places are estimates.

^a Only strongest components of the blends have been used to calculate energy levels.

^b All blend components have been used to calculate energy levels.

^c Newly determined experimental values and estimates.

^d Values including vibrational states not tabulated.

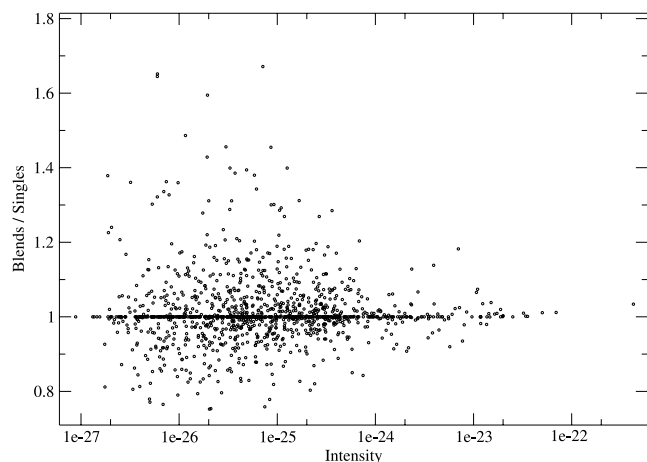


Fig. 1. Temperature dependence of the line intensity of transitions identified as blended when treated as blended compared to assigning all the line intensities to a single dominant transition. Intensities are calculated at temperature 316 °C and given in cm molecules^{-1} .

intensity. Furthermore, the intensity of a blended line will also display a temperature dependence which differs from that of the strongest line in the blend. To illustrate the effect of identifying the blended lines, Fig. 1 compares the predicted intensity of each of these lines for a temperature change of 20 °C. This sort of temperature change is typical of that found in an atmospheric column. It can be seen from the figure that for a number of lines the identification of the blend leads to significantly altered behaviour.

The splitting of the intensity of the blended lines according to the theoretical predictions of course relies heavily on the assumption that the calculated line intensities are accurate. Fig. 2 compares the line intensities

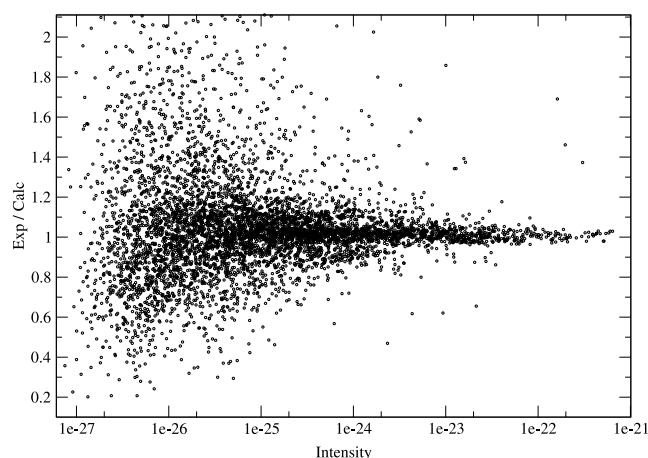


Fig. 2. Comparison of experimental line intensities with those calculated using the BT2 line list. Intensities, in cm molecules^{-1} , are for unblended transitions in the frequency range 9250–25 300 cm^{-1} and are shown as function of the observed line intensity.

calculated with the BT2 line list [25] in the frequency range 9250–25 300 cm^{-1} for transitions, which our analysis deemed unblended. It should be noted that our experience is that the transition intensities do not change significantly between calculations using the same dipole surface but different potential energy surfaces [40].

Fig. 2 shows very good agreement between the calculated and observed line strengths for the strong transitions. For transitions intensities greater than $10^{-23} \text{ cm molecules}^{-1}$, the agreement is a few percent for all but three outliers. Not surprisingly, the disagreements get larger as the lines become less intense. As the uncertainties in both theory and experiment increase as the transition intensity drops, it is unclear which is the major cause of this increased disagreement.

4. Discussion and conclusion

The Fourier-transform spectra of water vapour of [1–3] provide a uniform treatment of absorption by natural isotopic abundance water vapour throughout the visible region of the spectrum. The latest release of HITRAN [21] uses these data for the entire visible region, a key region for atmospheric physics. Our analysis has significantly increased the number of assigned transitions in the spectrum and identified over 1800 lines which are actually blends of more than one transition. These blends explain some of the highly anomalous linewidths which were reported in the original study and largely ascribed to pressure broadening. 1205 lines due to the minor isotopologues of water have also been identified and 25 lines arising from absorption by molecular oxygen have been removed from the line list. This re-analysed line list, a complete copy of which has been placed in the electronic journal archive, has been submitted for inclusion in an updated version of HITRAN.

Among newly assigned transitions are those associated with excited bending states such as (060), (061), (160), (070), (170), and (071). The energy levels of such states give insight into the intramolecular dynamics of water for high excitation of the bending vibration which probes the potential close to linear geometries and which is accompanied by the abnormal centrifugal distortion effects [41]. Direct transitions to these high bending states are weak. However, the transitions assigned here are observable due to intensity borrowing from nearby strong lines. From the perspective of effective Hamiltonian theory, this intensity borrowing occurs via high-order resonance interactions. As a result the line list contains transitions identified with vibrational states with $v_2 > 7$. Although the energies of the corresponding rovibrational levels can be found in the electronic archive these states are not presented in Table 4. Labelling of highly excited bending states remains difficult. The

labelling procedure described above works well for states with low and medium v_2 but begins to fail for $v_2 > 7$. This may well be because at these high levels of excitation the strong coupling between vibrational states makes such labels rather arbitrary [42].

Spectral analysis shows that the most intense transitions to highly excited bending states very often involve, in contrast to nearby stretching states, upper levels with high values of K_a . This point is illustrated in Fig. 3 where the experimental and calculated intensity of the strongest transition to the upper level with a given value of K_a is presented as a function of the generalised quantum number $N = J + K_a - K_c + 1$ for the $J = 9$ multiplet of the (201) stretching and the (060) bending state, respectively. It is clear from Fig. 3 that the intensity of the strongest transition to the (201) upper level belonging to the $J = 9$ multiplet smoothly decreases with increasing K_a . Conversely for the (060) state, there is a series of strong transitions (all observed in the spectrum) to the $K_a = 6, 7, 8,$ and 9 upper levels. These transitions

borrow intensity from lines to levels associated with the neighbouring (300), (201), and (022) vibrational states. These may be separated from the interacting level of (060) level by up to 50 cm^{-1} . Furthermore, there is a relatively strong predicted transition to the 9_{18} level of (060) (not observed in the spectrum as it lies under an even stronger transition) which is caused by a clear resonance with the (111) state, the energy levels lying at 10143 and 10136 cm^{-1} , respectively. The cause of such relatively strong transitions to the high K_a energy levels of the highly excited bending states is difficult to identify with confidence, since there are no energy levels of other states involving strong transitions in the vicinity of the considered bending levels. Intensity stealing from distant levels should only occur if the anharmonic resonance interaction increases strongly with both the bending v_2 and rotational K_a quantum numbers.

The aim of this and related studies was to provide as complete and accurate database of water spectral parameters as can be achieved. It is therefore worth considering how the present line list can be improved. First, although the BR spectra were of high enough quality to detect many transitions due to the minor isotopologues of water, recording these spectra reliably using natural abundance water vapour is difficult. Spectra of H_2^{17}O and H_2^{18}O were recorded using isotopically enhanced samples more than two decades ago at Kitt Peak by Chevillard and co-workers [31,32]. These spectra are only now being properly analysed [33–35] but should provide good parameters for these molecules. New HDO spectra covering the near-infrared and visible spectral regions have been recorded by the Bruxelles-Reims group, using the same Fourier-transform technique and conditions as described in this study. A line list of HDO experimental line parameters in the $11500\text{--}23000 \text{ cm}^{-1}$ region will appear in a forthcoming paper [43]. This should provide excellent parameters for HDO which are entirely missing from HITRAN for this region.

We note that 18 of our newly determined H_2^{16}O experimental energy levels were also recently observed [44] using cavity ring down spectroscopy (CRDS) in the $13310\text{--}13377 \text{ cm}^{-1}$ spectral region. The very high sensitivity provided by the CRDS technique allowed three times the number of the lines to be observed compared to this study within this limited spectral region. This CRDS study of water vapour and a similar one performed in the mid-infrared [45] have confirmed what was already suspected from theoretical analysis [46], as one looks deeper into the spectrum of water vapour more and more lines appear. There are a number of studies available [6,8,17–19,47] which probe deeper at the lower frequency end of the spectral range considered here. With the exception of the study by Brown et al. [6], these studies are most useful for characterising the position and strength of the weak lines. The studies are thus

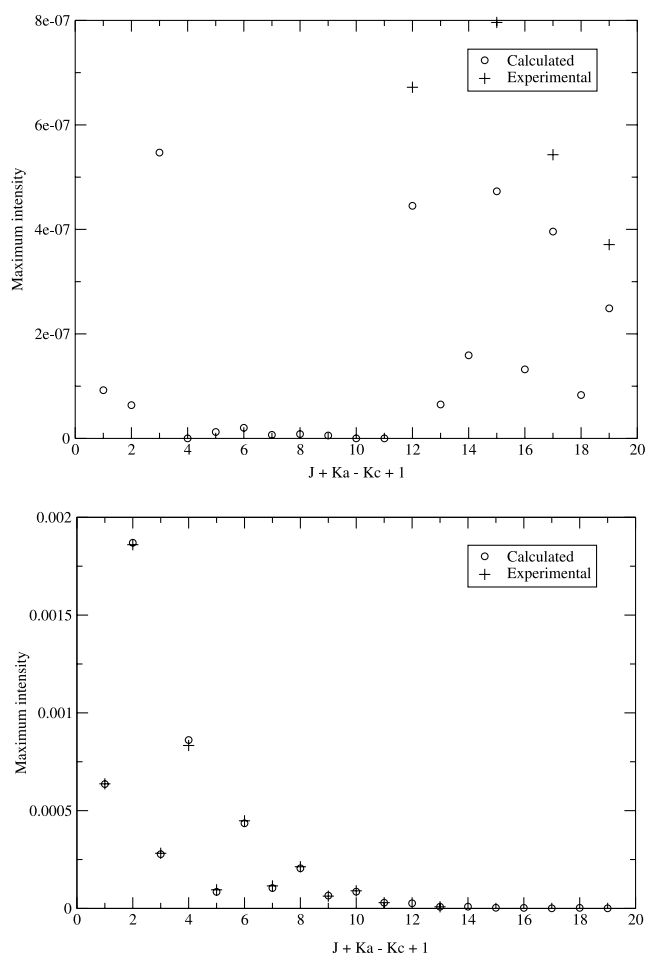


Fig. 3. Strongest transitions for $J = 9$ levels of (201) (upper panel) and (060) (lower panel) illustrating the different dependence on K_a . Intensities are given in $\text{cm}^{-2} \text{ atm}^{-1}$. Note the very different intensity scales.

largely complementary to the present work and a better database can be obtained by merging data from various sources. At present, however, the data for these weak transitions are only partial and more work needs to be done to obtain a complete picture.

This study follows a number of earlier ones using variational line lists to assign water transitions in the near-infrared and visible. The present study is very comprehensive in that there are few unassigned lines. However, it is anticipated that new spectra which probe even more deeply, such as the CRDS studies mentioned in the previous paragraph, will require further theoretical work.

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References

- [1] P.-F. Coheur, S. Fally, M. Carleer, C. Clerbaux, R. Colin, A. Jenouvrier, M.-F. Mérienne, C. Hermans, A.C. Vandaele, J. Quant. Spectrosc. Radiat. Transfer 74 (2002) 493–510.
- [2] M.-F. Mérienne, A. Jenouvrier, C. Hermans, A.C. Vandaele, M. Carleer, C. Clerbaux, P.-F. Coheur, R. Colin, S. Fally, M. Bach, J. Quant. Spectrosc. Radiat. Transfer 82 (2003) 99–117.
- [3] S. Fally, P.-F. Coheur, M. Carleer, C. Clerbaux, R. Colin, A. Jenouvrier, M.-F. Mérienne, C. Hermans, A.C. Vandaele, J. Quant. Spectrosc. Radiat. Transfer 82 (2003) 119–131.
- [4] J.-P. Chevillard, J.-Y. Mandin, J.-M. Flaud, C. Camy-Peyret, Can. J. Phys. 67 (1989) 1065–1084.
- [5] R.A. Toth, J. Mol. Spectrosc. 166 (1994) 176–183.
- [6] L.R. Brown, R.A. Toth, M. Dulick, J. Mol. Spectrosc. 212 (2002) 57–82.
- [7] R. Schermaul, R.C.M. Learner, D.A. Newnham, R.G. Williams, J. Ballard, N.F. Zobov, D. Belmiloud, J. Tennyson, J. Mol. Spectrosc. 208 (2001) 32–42.
- [8] R. Schermaul, J.W. Brault, A.A.D. Canas, R.C.M. Learner, O.L. Polyansky, N.F. Zobov, D. Belmiloud, J. Tennyson, J. Mol. Spectrosc. 211 (2002) 169–178.
- [9] C. Camy-Peyret, J.M. Flaud, J.Y. Mandin, J.P. Chevillard, J. Brault, D.A. Ramsay, M. Vervloet, J. Chauville, J. Mol. Spectrosc. 113 (1985) 208–228.
- [10] J.Y. Mandin, J.P. Chevillard, C. Camy-Peyret, J.M. Flaud, J. Mol. Spectrosc. 116 (1986) 167.
- [11] O.L. Polyansky, N.F. Zobov, S. Viti, J. Tennyson, P.F. Bernath, L. Wallace, Science 277 (1997) 346–349.
- [12] H. Partridge, D.W. Schwenke, J. Chem. Phys. 106 (1997) 4618.
- [13] O.L. Polyansky, N.F. Zobov, S. Viti, J. Tennyson, J. Mol. Spectrosc. 189 (1998) 291–300.
- [14] D.W. Schwenke, J. Mol. Spectrosc. 190 (1998) 397.
- [15] M. Carleer, A. Jenouvrier, A.-C. Vandaele, P.F. Bernath, M.F. Mérienne, R. Colin, N.F. Zobov, O.L. Polyansky, J. Tennyson, V.A. Savin, J. Chem. Phys. 111 (1999) 2444–2450.
- [16] N.F. Zobov, D. Belmiloud, O.L. Polyansky, J. Tennyson, S.V. Shirin, M. Carleer, A. Jenouvrier, A.-C. Vandaele, P.F. Bernath, M.F. Mérienne, R. Colin, J. Chem. Phys. 113 (2000) 1546–1552.
- [17] R.N. Tolchenov, J. Tennyson, J.W. Brault, A.A.D. Canas, R. Schermaul, J. Mol. Spectrosc. 215 (2002) 269–274.
- [18] O. Naumenko, A. Campargue, J. Mol. Spectrosc. 221 (2003) 221–226.
- [19] R.N. Tolchenov, J. Tennyson, S.V. Shirin, N.F. Zobov, O.L. Polyansky, A.N. Maurellis, J. Mol. Spectrosc. 221 (2003) 99–105.
- [20] J. Tennyson, N.F. Zobov, R. Williamson, O.L. Polyansky, P.F. Bernath, J. Phys. Chem. Ref. Data 30 (2001) 735–831.
- [21] L. Rothman et al., J. Quant. Spectrosc. Radiat. Transfer (in press).
- [22] N. Jacquinet-Husson et al., J. Quant. Spectrosc. Radiat. Transfer 95 (2005) 429.
- [23] D.W. Schwenke, H. Partridge, J. Chem. Phys. 113 (2000) 16.
- [24] S.V. Shirin, O.L. Polyansky, N.F. Zobov, P. Barletta, J. Tennyson, J. Chem. Phys. 118 (2003) 2124–2129.
- [25] R.J. Barber, J. Tennyson, G.J. Harris, R.N. Tolchenov, Mon. Not. R. astr. Soc. submitted.
- [26] P. Dupré, T. Germain, R.N. Tolchenov, N.F. Zobov, J. Tennyson, J. Chem. Phys. (in press).
- [27] A.D. Bykov, O.V. Naumenko, A.M. Pshenichnikov, L.N. Sinitsa, A.P. Scherbakov, Opt. Spectrosc. 94 (2003) 528–537.
- [28] J.M. Flaud, C. Camy-Peyret, A. Bykov, O. Naumenko, T. Petrova, A. Scherbakov, L. Sinitsa, J. Mol. Spectrosc. 185 (1997) 211–221.
- [29] J.M. Flaud, C. Camy-Peyret, A. Bykov, O. Naumenko, T. Petrova, A. Scherbakov, L. Sinitsa, J. Mol. Spectrosc. 183 (1997) 300–309.
- [30] R.N. Tolchenov, J. Tennyson, J. Mol. Spectrosc. 231 (2005) 35.
- [31] C. Camy-Peyret, J.-M. Flaud, J.-Y. Mandin, A. Bykov, O. Naumenko, L. Sinitsa, B. Voronin, J. Quant. Spectrosc. Radiat. Transfer 61 (1999) 795–812.
- [32] J.-P. Chevillard, J.-Y. Mandin, J.-M. Flaud, C. Camy-Peyret, Can. J. Phys. 65 (1987) 777–789.
- [33] A. Bykov, O. Naumenko, T. Petrova, A. Scherbakov, L. Sinitsa, J.-Y. Mandin, C. Camy-Peyret, J.-M. Flaud, J. Mol. Spectrosc. 172 (1995) 243–253.
- [34] M. Tanaka, J.W. Brault, J. Tennyson, J. Mol. Spectrosc. 216 (2002) 77–80.
- [35] M. Tanaka, O. Naumenko, J.W. Brault, J. Tennyson, J. Mol. Spectrosc. (to be submitted).
- [36] O. Naumenko, E. Bertseva, A. Campargue, J. Mol. Spectrosc. 197 (1999) 122–132.
- [37] O. Naumenko, A. Campargue, J. Mol. Spectrosc. 199 (2000) 59–72.
- [38] E. Bertseva, O. Naumenko, A. Campargue, J. Mol. Spectrosc. 221 (2003) 38–46.
- [39] O. Naumenko, S.M. Hu, S.G. He, A. Campargue, Phys. Chem. Chem. Phys. 6 (2003) 910–918.
- [40] A. Callegari, P. Theule, R.N. Tolchenov, N.F. Zobov, O.L. Polyansky, J. Tennyson, J.S. Muentner, T.R. Rizzo, Science 297 (2002) 993–995.
- [41] A. Bykov, O. Naumenko, L. Sinitsa, B. Voronin, J.-M. Flaud, C. Camy-Peyret, R. Lanquetin, J. Mol. Spectrosc. 205 (2001) 1–8.
- [42] J.P. Rose, M.E. Kellman, J. Chem. Phys. 105 (1994) 7348–7363.
- [43] M. Bach, S. Fally, P.-F. Coheur, M. Carleer, A. Jenouvrier, A.C. Vandaele, J. Mol. Spectrosc. (2005) in press.

- [44] S. Kassi, P. Macko, O. Naumenko, A. Campargue, *J. Mol. Spectrosc.* (to be published).
- [45] P. Macko, D. Romanini, S.N. Mikhailenko, O.V. Naumenko, S. Kassi, A. Jenouvrier, V.G. Tyuterev, A. Campargue, *J. Mol. Spectrosc.* 227 (2004) 90–108.
- [46] R.C.M. Learner, W. Zhong, J.D. Haigh, D. Belmiloud, J. Clarke, *Geophys. Res. Lett.* 26 (1999) 3609–3612.
- [47] H. Naus, W. Ubachs, P.F. Levert, O.L. Polyansky, N.F. Zobov, J. Tennyson, *J. Mol. Spectrosc.* 205 (2001) 117–121.